



Combined Analysis of Neutrino and Antineutrino Oscillations at T2K

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T2K reports its first results in the search for *CP* violation in neutrino oscillations using appearance and disappearance channels for neutrino- and antineutrino-mode beams. The data include all runs from January

2010 to May 2016 and comprise 7.482×10^{20} protons on target in neutrino mode, which yielded in the far detector 32 e -like and 135 μ -like events, and 7.471×10^{20} protons on target in antineutrino mode, which yielded 4 e -like and 66 μ -like events. Reactor measurements of $\sin^2 2\theta_{13}$ have been used as an additional constraint. The one-dimensional confidence interval at 90% for the phase δ_{CP} spans the range $(-3.13, -0.39)$ for normal mass ordering. The CP conservation hypothesis ($\delta_{CP} = 0, \pi$) is excluded at 90% C.L.

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Introduction.—A new source of CP violation beyond the Cabibbo-Kobayashi-Masakawa quark mixing matrix is necessary to explain observations of baryon asymmetry in the Universe. In the lepton sector the Pontecorvo-Maki-Nakagawa-Sakata framework [1,2] allows for CP violation. The first indication of nonzero θ_{13} [3] followed by its discovery [4–6] and then the discovery of $\nu_\mu \rightarrow \nu_e$ oscillation by T2K [7] have opened the possibility to look for CP violation in neutrino oscillation.

In this Letter we present the first joint fit of neutrino and antineutrino $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$ and $\nu_\mu \rightarrow \nu_e$ oscillation at T2K. The mixing of neutrinos in the three-flavor framework is represented by the unitary PMNS matrix, parameterized by three mixing angles, θ_{12} , θ_{13} , and θ_{23} , and a CP -violating phase δ_{CP} [8]. The probability for $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$ oscillation, as a function of neutrino propagation distance L and energy E , can be written

$$P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e) \approx 4c_{13}^2 s_{13}^2 s_{23}^2 \sin^2 \phi_{31} \left[1 + \frac{2a}{\Delta m_{31}^2} (1 - 2s_{13}^2) \right] \\ + 8c_{13}^2 c_{12} c_{23} s_{12} s_{13} s_{23} \sin \phi_{32} \sin \phi_{31} \sin \phi_{21} \sin \delta_{CP} \\ + 8c_{13}^2 s_{13}^2 s_{23}^2 (1 - 2s_{13}^2) \frac{aL}{4E} \cos \phi_{32} \sin \phi_{31} \\ + (CP\text{-even, solar terms}), \quad (1)$$

where $s_{ij} = \sin \theta_{ij}$, $c_{ij} = \cos \theta_{ij}$, $\phi_{ij} = \Delta m_{ij}^2 L / 4E$, and $\Delta m_{ij}^2 = m_i^2 - m_j^2$ represents the neutrino mass-squared difference between mass eigenstates i and j . Matter effects are included to first order in the terms $a[\text{eV}^2/c^4] = 7.56 \times 10^{-5} E[\text{GeV}] \rho[\text{g/cm}^3]$. Our analyses use the complete probability calculation, without approximating matter effects. The $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$ survival probability is dominated by the parameters $\sin^2 \theta_{23}$ and Δm_{32}^2 , as given in [9]. Comparing electron neutrino and antineutrino

appearance probabilities allows a direct measurement of CP violation at T2K. The asymmetry variable $[A_{CP} = P(\nu_\mu \rightarrow \nu_e) - P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)] / [P(\nu_\mu \rightarrow \nu_e) + P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)]$ and the ν_μ ($\bar{\nu}_\mu$) component of the expected T2K flux without oscillations are shown in Fig. 1. At the flux peak energy, A_{CP} can be as large as 0.4, including a contribution of around 0.1 due to matter effects.

The T2K experiment.—The T2K experiment [10] uses a 30-GeV proton beam from the J-PARC accelerator facility to produce a muon (anti)neutrino beam. The proton beam strikes a graphite target to produce charged pions and kaons, which are focused by three magnetic horns. Depending on the polarity of the horn current, either positively or negatively charged mesons are focused, resulting in a beam largely composed of muon neutrinos or antineutrinos. A 96-m decay volume lies downstream of the magnetic horns, followed by the beam dump and muon monitor [11]. The neutrino beam is measured by detectors placed on axis and off axis at 2.5° relative to the beam direction. The off-axis neutrino energy spectrum peaks at 0.6 GeV, and has a reduced $\bar{\nu}_\mu$ contamination and smaller backgrounds from higher-energy neutrinos than the on-axis spectrum. Two detectors located 280 m from the target are used to measure the beam direction, spectrum, and composition, as well as

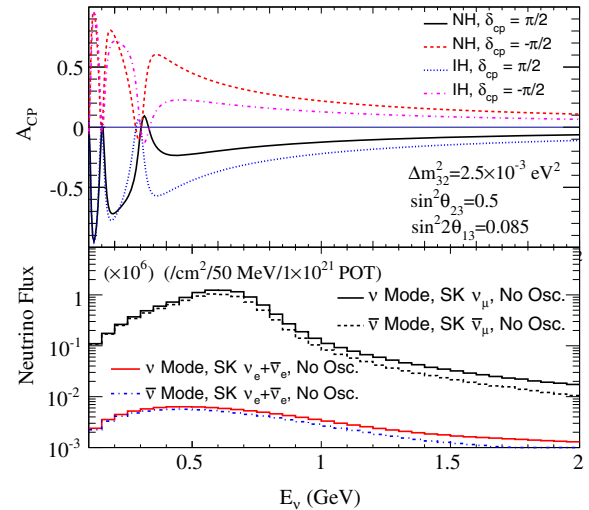


FIG. 1. The leptonic CP asymmetry, $A_{CP} = [P(\nu_\mu \rightarrow \nu_e) - P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)] / [P(\nu_\mu \rightarrow \nu_e) + P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)]$, as a function of energy for maximal CP -violation hypotheses (top) and the ν_μ ($\bar{\nu}_\mu$) and $\nu_e + \bar{\nu}_e$ components of the unoscillated neutrino flux in neutrino and antineutrino modes (bottom).

the event rate: INGRID (on axis) [12] and ND280 (off axis), which is housed inside a 0.2-T magnet. The Super-Kamiokande (Super-K) 50-kt water Cherenkov detector [13], located off axis and 295 km from the neutrino production point, is used to detect oscillated neutrinos.

Data sets.—The results presented here are based on data collected in two periods: one in which the beam operated solely in neutrino mode, January 2010–May 2013, and one in which the beam operated mostly in antineutrino mode, May 2014–May 2016. This comprises a neutrino beam exposure of 7.482×10^{20} protons on target (POT) in neutrino mode and 7.471×10^{20} POT in antineutrino mode for the far-detector analysis, and an exposure of 5.82×10^{20} POT in neutrino mode and 2.84×10^{20} POT in antineutrino mode for the near-detector analysis.

Analysis strategy.—The analysis strategy is similar to that of previous T2K results [7,9,14,15]: oscillation parameters are estimated by comparing predictions and observations at the far detector. A tuned prediction of the oscillated spectrum at the far detector, with associated uncertainty, is obtained by fitting samples of charged-current interactions at ND280. The analysis presented here differs from previous results in that both neutrino and antineutrino samples are fitted at both ND280 and Super-K. Including antineutrino data at ND280 ensures that the interaction model is consistent between neutrinos and antineutrinos. Additionally, the use of a magnetized near detector with charge-selected samples in both neutrino and antineutrino beams allows a constraint on wrong-sign contaminations in the beam.

Neutrino flux model.—The T2K neutrino and antineutrino fluxes at near and far detectors, and their correlations, are calculated [16] using a data-driven hybrid simulation with FLUKA 2011 [17] used to simulate hadronic interactions and transport particles inside the target, while GEANT3 [18] with G4CALOR [19] is used to simulate the rest of the neutrino beam line. The interactions of hadrons in both FLUKA 2011 and G4CALOR are tuned using thin target hadron production data, including measurements of the total cross section for particle production, and π^\pm , K^\pm , p^+ , Λ , and K_S^0 production with 30-GeV protons on a graphite target by the NA61/SHINE experiment [20]. Dominant systematic error sources include uncertainties on the NA61/SHINE hadron production measurements, hadronic interaction length measurements from NA61/SHINE and other experiments, the initial proton beam trajectory, and the horn currents. The total uncertainty on the flux near the peak energy is $\sim 9\%$. The ν_μ ($\bar{\nu}_\mu$) component of the predicted fluxes without oscillations are shown in Fig. 1. At the far detector and in the absence of oscillations, we predict that 94.1% (92.3%) of the T2K neutrino-mode (antineutrino-mode) beam below 1.25 GeV is ν_μ ($\bar{\nu}_\mu$). The $\bar{\nu}_\mu$ flux in antineutrino mode is reduced by $\sim 20\%$ relative to the ν_μ flux in neutrino mode due to the smaller production cross section for π^- relative to π^+ in 30 GeV $p + C$ interactions.

Neutrino interaction model.—The interactions of neutrinos and antineutrinos with nuclei in the near and far detectors are modeled with the NEUT [21] neutrino interaction generator. The charged-current quasielastic (CCQE) interactions are modeled with a relativistic Fermi gas nuclear model with relativistic corrections for long-range correlations using the random phase approximation (RPA) as applied by Nieves *et al.* [22]. The choice of the CCQE nuclear model was made based on fits to external CCQE-like data [23] from the MiniBooNE [24,25] and MINERvA [26,27] experiments. Interactions on more than one nucleon are modeled with an implementation of the $2p$ - $2h$ model developed by Nieves *et al.* [28,29]. These interactions are characterized by multinucleon ejection and no final state pions; hence, they may be confused for CCQE interactions in a water Cherenkov detector. The single-pion production model in NEUT has been tuned using form factors from Graczyk and Sobczyk [30] and with a reanalysis of Argonne National Laboratory and Brookhaven National Laboratory bubble chamber data sets [31]. The coherent pion production model has been tuned to reproduce data from MINERvA [32] and T2K [33]. At the T2K peak energy, the antineutrino cross section is ~ 3.5 times smaller than the neutrino cross section.

The parameterization of uncertainties in the neutrino interaction model is largely unchanged from previous measurements [14,15]. Parameters that vary the binding energy, Fermi momentum, $2p$ - $2h$ normalization, and charged current (CC) coherent pion production cross-section normalization are applied separately for interactions on carbon and oxygen. To cover the different predictions by Nieves *et al.* [28,29] and Martini *et al.* [34,35] of the relative $2p$ - $2h$ interaction rates for neutrinos and antineutrinos, the normalizations of $2p$ - $2h$ interactions for neutrinos and antineutrinos are allowed to vary independently.

Only the interactions of ν_μ and $\bar{\nu}_\mu$ are explicitly constrained by near-detector measurements in this analysis. Since the oscillation signals include ν_e and $\bar{\nu}_e$ interactions, it is necessary to assign uncertainties on the cross-section ratios $\sigma_{\nu_e}/\sigma_{\nu_\mu}$ and $\sigma_{\bar{\nu}_e}/\sigma_{\bar{\nu}_\mu}$. Following the treatment in [36], separate parameters for $\sigma_{\nu_e}/\sigma_{\nu_\mu}$ and $\sigma_{\bar{\nu}_e}/\sigma_{\bar{\nu}_\mu}$ are introduced with a theoretical uncertainty of 2.8% for each. A correlation coefficient of -0.5 is assumed for these two parameters.

Fit to near-detector data.—The systematic parameters in the neutrino flux and interaction models are constrained by a fit to CC candidate samples in the ND280 [10] near detector. The data sets used consist of reconstructed interactions in two fine-grained detectors (FGDs) [37] with particle tracking in three time projection chambers (TPCs) [38]. FGD2 contains six 2.54-cm-thick water panels, allowing systematic parameters governing neutrino interactions on H_2O , the same target as Super-K, to be directly constrained. The CC candidate samples in ND280 are divided into categories based on the beam mode (neutrino vs antineutrino), the FGD in which the interaction takes

place, the muon charge, and the final-state multiplicity. For data taken in neutrino mode, only interactions with a negatively charged muon are considered. For data taken in antineutrino mode, there are separate categories for events with positively charged (right-sign) and negatively charged (wrong-sign) muon candidates. The wrong-sign candidates are included because the larger neutrino cross section leads to a non-negligible wrong-sign background in antineutrino mode. In neutrino mode, there are three categories for reconstructed final states: no pion candidate in the final state (CC0 π), one pion candidate in the final state (CC1 π), and all other CC candidates (CC other). In antineutrino mode, events are divided into two categories based on the final states: only the muon track exits the FGD to enter the TPC (CC 1-track) and at least one other track enters the TPC (CC N -track).

When fitting, the data are binned according to the momentum of the muon candidate, p_μ , and $\cos\theta_\mu$, where θ_μ is the angle of the muon direction relative to the central axis of the detector, roughly 1.7° away from the incident (anti)neutrino direction. A binned maximum likelihood fit is performed in which the neutrino flux and interaction model parameters are allowed to vary. Nuisance parameters describing the systematic errors in the ND280 detector

model—the largest of which is pion interaction modeling—are marginalized in the fit.

The fitted p_μ and $\cos\theta_\mu$ distributions for the FGD2 CC0 π and CC 1-track categories are shown in Fig. 2. Acceptable agreement between the postfit model and data is observed for both kinematic variables, with a p value of 0.086. The best-fit fluxes are increased with respect to the original flux model by 10%–15% near the flux peak. This is driven by the prefit deficit in the prediction for the CC0 π and CC other samples. The fitted value for the axial mass in the CCQE model is $1.12 \text{ GeV}/c^2$, compared to $1.24 \text{ GeV}/c^2$ in a previous fit where the $2p$ - $2h$ model and RPA corrections were not included [14]. The lower axial mass decreases the interaction rate, driving the increased flux prediction. The fit to ND280 data reduces the uncertainty on the event-rate predictions at the far detector due to uncertainties on the flux and ND280-constrained interaction model parameters from 10.9% (12.4%) to 2.9% (3.2%) for the ν_e ($\bar{\nu}_e$) candidate sample.

Far-detector data.—At the far detector, events are extracted that lie within $[-2, 10] \mu\text{s}$ relative to the beam arrival. Fully contained events within the fiducial volume are selected by requiring that no hit cluster is observed in the outer detector volume, that the distance from the reconstructed vertex to the inner detector wall is larger than 2 m, and that the total observed charge is greater than the equivalent quantity for a 30-MeV electron. The CCQE component of our sample is enhanced by selecting events with a single Cherenkov ring. The $\nu_\mu/\bar{\nu}_\mu$ CCQE candidate samples are then selected by requiring a μ -like ring using a particle identification likelihood, zero or one decay electron candidates, and muon momentum greater than $200 \text{ MeV}/c$ to reduce pion background. Post selection, 135 and 66 events remain in the ν_μ and $\bar{\nu}_\mu$ candidate samples, respectively, while if $|\Delta m_{32}^2| = 2.509 \times 10^{-3} \text{ eV}^2/c^4$ and $\sin^2\theta_{23} = 0.528$ (i.e., maximal disappearance), 135.5 and 64.1 events are expected. The $\nu_e/\bar{\nu}_e$ CCQE candidate samples are selected by requiring an e -like ring and zero decay electron candidates, not π^0 -like and reconstructed energy less than 1.25 GeV . The total number of events remaining in these samples is presented in Table I with their respective expectation for different values of δ_{CP} , $\sin^2\theta_{13} = 0.085$, $|\Delta m_{32}^2| = 2.509 \times 10^{-3} \text{ eV}^2/c^4$, and

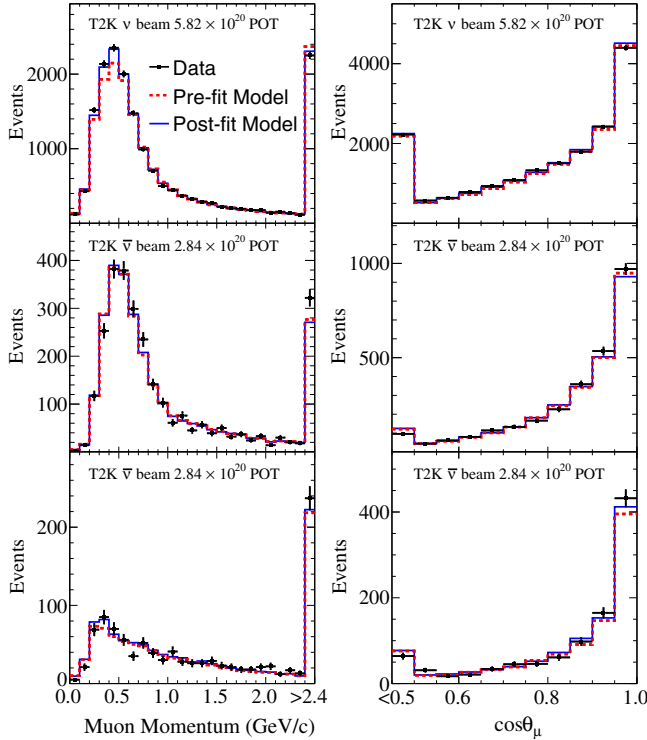


FIG. 2. The FGD2 data, prefit predictions and postfit predictions binned in p_μ (left) and $\cos\theta_\mu$ (right) for the neutrino mode CC0 π (top), antineutrino mode CC 1-track μ^+ (middle) and antineutrino mode CC 1-track μ^- (bottom) categories. The overflow bins are integrated out to $10000 \text{ MeV}/c$ for p_μ and -1.0 for $\cos\theta_\mu$ respectively.

TABLE I. Number of ν_e and $\bar{\nu}_e$ events expected for various values of δ_{CP} and both mass orderings compared to the observed numbers.

Normal	$\delta_{CP} = -\pi/2$	$\delta_{CP} = 0$	$\delta_{CP} = \pi/2$	$\delta_{CP} = \pi$	Observed
ν_e	28.7	24.2	19.6	24.1	32
$\bar{\nu}_e$	6.0	6.9	7.7	6.8	4
Inverted	$\delta_{CP} = -\pi/2$	$\delta_{CP} = 0$	$\delta_{CP} = \pi/2$	$\delta_{CP} = \pi$	Observed
ν_e	25.4	21.3	17.1	21.3	32
$\bar{\nu}_e$	6.5	7.4	8.4	7.4	4

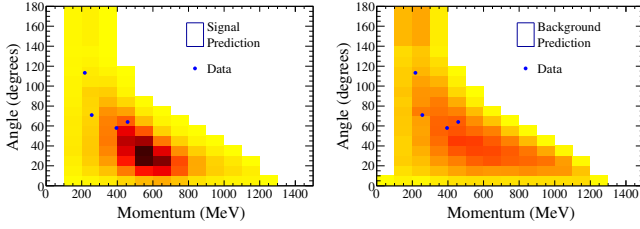


FIG. 3. The reconstructed lepton momentum and angle relative to the beam at the far detector for the $\bar{\nu}_e$ sample signal (left) and background (right) expectation with the data overlaid (blue points).

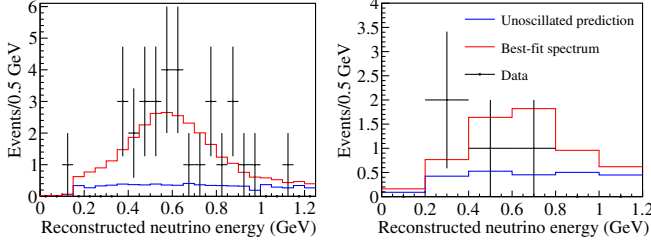


FIG. 4. The reconstructed neutrino energy at the far detector for the ν_e (left) and $\bar{\nu}_e$ (right) candidate samples is shown together with the expected distribution without oscillation (blue histogram) and the best fit (red histogram).

$\sin^2 \theta_{23} = 0.528$. The ν_e ($\bar{\nu}_e$) contamination in the $\bar{\nu}_e$ (ν_e) sample is 17.4% (0.5%), and the proportion of the sample expected to correspond to oscillated $\bar{\nu}_e$ (ν_e) events is 46.4% (80.9%) for $\delta_{CP} = -\pi/2$. A more detailed description of the candidate event selections can be found in previous publications [14]. The $\bar{\nu}_e$ signal events are concentrated in the forward direction with respect to the beam, unlike the backgrounds (Fig. 3). Therefore, incorporating reconstructed lepton angle information in the analysis increases the sensitivity. The reconstructed neutrino energy spectra for the ν_e and $\bar{\nu}_e$ samples is shown in Fig. 4.

The systematic errors concerning the detector behavior are estimated using atmospheric neutrino and cosmic-ray muon events. A sample of hybrid data-Monte Carlo events is also used to evaluate uncertainties regarding π^0 rejection.

TABLE II. Systematic uncertainty on the predicted event rate at the far detector.

Source (%)	ν_μ	ν_e	$\bar{\nu}_\mu$	$\bar{\nu}_e$
ND280-unconstrained cross section	0.7	3.0	0.8	3.3
Flux and ND280-constrained cross section	2.8	2.9	3.3	3.2
Super-Kamiokande detector systematics	3.9	2.4	3.3	3.1
Final or secondary hadron interactions	1.5	2.5	2.1	2.5
Total	5.0	5.4	5.2	6.2

Correlations between the uncertainties for the four samples are taken into account in the fits.

The fractional variation of the number of expected events for the four samples owing to the various sources of systematic uncertainty are shown in Table II. A more in-depth description of the sources of systematic uncertainty in the fit is given in [14], although this reference does not cover the updates discussed in previous sections.

Oscillation analysis.—The oscillation parameters $\sin^2 \theta_{23}$, Δm_{32}^2 , $\sin^2 \theta_{13}$, and δ_{CP} are estimated by performing a joint maximum-likelihood fit of the four far-detector samples. The oscillation probabilities are calculated using the full three-flavor oscillation formulas [39]. Matter effects are included with an Earth density of $\rho = 2.6 \text{ g/cm}^3$ [40].

As described previously, the priors for the beam flux and neutrino interaction cross-section parameters are obtained from the fit with the near-detector data. The priors [8] for the solar neutrino oscillation parameters—whose impact is almost negligible—are $\sin^2 2\theta_{12} = 0.846 \pm 0.021$, $\Delta m_{21}^2 = (7.53 \pm 0.18) \times 10^{-5} \text{ eV}^2/c^4$, and in some fits we use $\sin^2 2\theta_{13} = 0.085 \pm 0.005$ [8], called the “reactor measurement.” Flat priors are used for $\sin^2 \theta_{23}$, Δm_{32}^2 , and δ_{CP} .

We use a procedure analogous to [15]: we integrate the likelihood over the prior probability density function of the nuisance parameters and we obtain the marginal likelihood which depends only on the relevant oscillation parameters. We define $-2\Delta \ln \mathcal{L} = -2 \ln[\mathcal{L}(\mathbf{o})/\mathcal{L}_{\max}]$ as the ratio between the marginal likelihood at the point \mathbf{o} of the relevant oscillation parameter space and the maximum marginal likelihood.

We have conducted three analyses using different far-detector event quantities and different statistical approaches. All of them use the neutrino energy reconstructed in the CCQE hypothesis (E_{rec}) for the $(\bar{\nu}_\mu^-)$ samples. The first analysis uses E_{rec} and the reconstructed angle between the lepton and the neutrino beam direction, θ_{lep} , of the $(\bar{\nu}_e^-)$ candidate samples and provides confidence intervals using a hybrid Bayesian-frequentist approach [41]. These results are shown in the following figures. The second analysis is fully Bayesian and uses the lepton

momentum, p_{lep} , and θ_{lep} for the $(\bar{\nu}_e^-)$ samples to compute credible intervals using the posterior probability. The third analysis uses only E_{rec} spectra for the $(\bar{\nu}_e^-)$ samples and a Markov chain Monte Carlo method [42] to provide Bayesian credible intervals. This analysis performs a simultaneous fit of both the near- and far-detector data, providing a validation of the extrapolation of the flux, cross section, and detector systematic parameters from the near to far detector. All three methods are in good agreement.

An indication of the sensitivity to δ_{CP} and the mass ordering can be obtained from Table I. If CP violation is maximal ($\delta_{CP} = \pm\pi/2$), the predicted variation of the total number of events with respect to the CP conservation

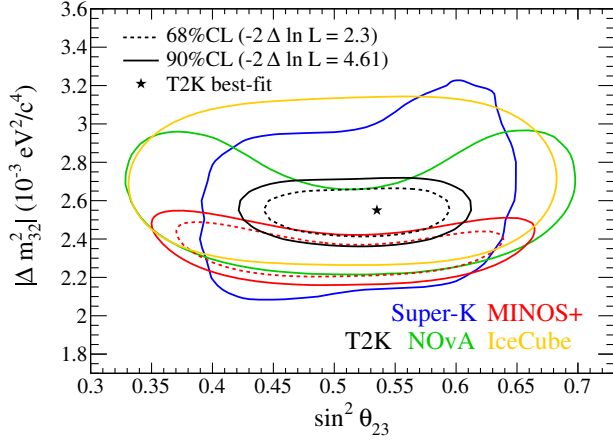


FIG. 5. The 68% (90%) constant $-2\Delta \ln L$ confidence regions for the $\sin^2 \theta_{23} - |\Delta m_{32}^2|$ plane assuming normal ordering, alongside NOvA [43], MINOS+[44], SK [45], and IceCube [46] confidence regions.

hypothesis ($\delta_{CP} = 0, \pi$) is about 20%. The different mass orderings induce a variation of the number of expected events of about 10%. Matter effects are negligible for the ν_μ and $\bar{\nu}_\mu$ candidate samples, while they affect the number of events in the ν_e and $\bar{\nu}_e$ candidate samples by about 6% and 4%, respectively, for maximal CP violation.

A series of fits are performed where one or two oscillation parameters are determined and the others are marginalized. Confidence regions are set using the constant $-2\Delta \ln L$ method [8]. In the first fit confidence regions in the $\sin^2 \theta_{23} - |\Delta m_{32}^2|$ plane (Fig. 5) were computed using the reactor measurement of $\sin^2 \theta_{13}$. The best-fit values are $\sin^2 \theta_{23} = 0.532$ and $|\Delta m_{32}^2| = 2.545 \times 10^{-3} \text{ eV}^2/c^4$ ($\sin^2 \theta_{23} = 0.534$ and $|\Delta m_{32}^2| = 2.510 \times 10^{-3} \text{ eV}^2/c^4$) for

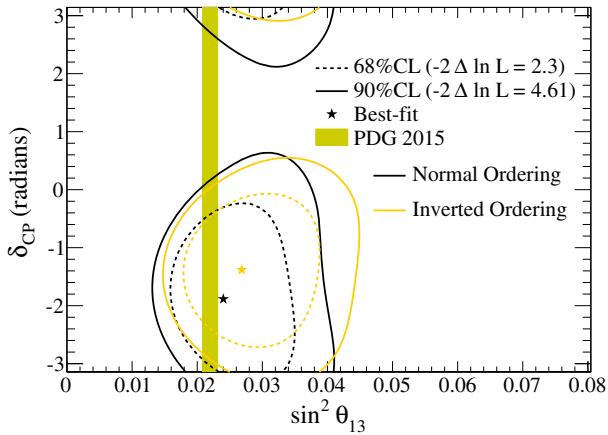


FIG. 6. The 68% (90%) constant $-2\Delta \ln L$ confidence regions in the $\delta_{CP} - \sin^2 \theta_{13}$ plane are shown by the dashed (continuous) lines, computed independently for the normal (black) and inverted (yellow) mass ordering. The best-fit point is shown by a star for each mass-ordering hypothesis. The 68% confidence region from reactor experiments on $\sin^2 \theta_{13}$ is shown by the yellow vertical band.

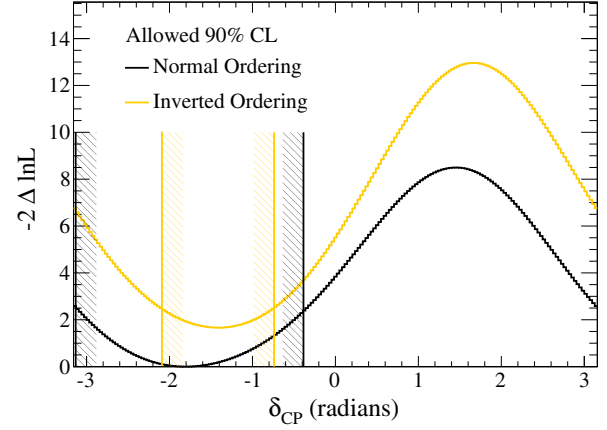


FIG. 7. $-2\Delta \ln \mathcal{L}$ as a function of δ_{CP} for the normal (black) and inverted (yellow) mass ordering. The vertical lines show the corresponding allowed 90% confidence intervals, calculated using the Feldman-Cousins method. $\sin^2 \theta_{13}$ is marginalized using the reactor measurement as prior probability.

the normal (inverted) ordering. The goodness of fit for all three analyses is better than 80%. The result is consistent with maximal disappearance. The T2K data weakly prefer the second octant ($\sin^2 \theta_{23} > 0.5$) with a posterior probability of 61%.

Confidence regions in the $\sin^2 \theta_{13} - \delta_{CP}$ plane are computed independently for both mass-ordering hypotheses (Fig. 6) without using the reactor measurement. The addition of antineutrino samples at Super-K gives the first sensitivity to δ_{CP} from T2K data alone. There is good agreement between the T2K result and the reactor measurement for $\sin^2 \theta_{13}$. For both mass-ordering hypotheses, the best-fit value of δ_{CP} is close to $-\pi/2$.

Confidence intervals for δ_{CP} are obtained using the Feldman-Cousins method [47]. The parameter $\sin^2 \theta_{13}$ is marginalized using the reactor measurement. The best-fit value is obtained for the normal ordering and $\delta_{CP} = -1.791$, close to maximal CP violation (Fig. 7). For inverted ordering the best-fit value of δ_{CP} is -1.414 . The hypothesis of CP conservation ($\delta_{CP} = 0, \pi$) is excluded at 90% C.L. and $\delta_{CP} = 0$ is excluded at more than 2σ . The δ_{CP} confidence intervals at 90% C.L. are $(-3.13, -0.39)$ for normal ordering and $(-2.09, -0.74)$ for inverted ordering. The Bayesian credible interval at 90%, marginalizing over the mass ordering, is $(-3.13, -0.21)$. The normal ordering is weakly favored over the inverted ordering with a posterior probability of 75%.

Sensitivity studies show that, if the true value of δ_{CP} is $-\pi/2$ and the mass ordering is normal, the fraction of pseudoexperiments where CP conservation ($\delta_{CP} = 0, \pi$) is excluded with a significance of 90% C.L. is 17.3%, with the amount of data used in this analysis.

Conclusions.—T2K has performed the first search for CP violation in neutrino oscillations using $\nu_\mu \rightarrow \nu_e$ appearance and $\nu_\mu \rightarrow \nu_\mu$ disappearance channels in neutrino and

antineutrino modes. The one-dimensional confidence interval at 90% for δ_{CP} spans the range $(-3.13, -0.39)$ in the normal mass ordering. The CP -conservation hypothesis ($\delta_{CP} = 0, \pi$) is excluded at 90% C.L. The data related to the measurements and results presented in this Letter can be found in Ref. [48].

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Note added.—Recently, a paper by the NO ν A Collaboration has appeared [49], in which $\sin^2 \theta_{23} = 0.5$ is disfavored by the data at 2.6σ . Considering their measurement, $\sin^2 \theta_{23} = 0.404^{+0.030}_{-0.022}$ for normal ordering, and the fact that the 68% C.L. interval for T2K measurement [50] extends

from 0.464 to 0.578, with a best fit of 0.532, the tension between the two measurements is rather mild (1.7σ). Several systematic effects (including additional smearing effects on the reconstructed energy) might produce a bias in the $\sin^2 \theta_{23}$ measurement and they must be studied with care. We have investigated them, including possible multi-nucleon knockout in neutrino-nucleus interactions [15]. This last effect is not a significant uncertainty source at the present statistical precision.

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- [1] Z. Maki, M. Nakagawa, and S. Sakata, *Prog. Theor. Phys.* **28**, 870 (1962).
- [2] B. Pontecorvo, *Sov. Phys. JETP* **26**, 984 (1968).
- [3] K. Abe *et al.* (T2K Collaboration), *Phys. Rev. Lett.* **107**, 041801 (2011).
- [4] F. P. An *et al.*, *Phys. Rev. Lett.* **108**, 171803 (2012).
- [5] J. K. Ahn *et al.*, *Phys. Rev. Lett.* **108**, 191802 (2012).
- [6] Y. Abe *et al.* (Double Chooz Collaboration), *Phys. Rev. Lett.* **108**, 131801 (2012).
- [7] K. Abe *et al.* (T2K Collaboration), *Phys. Rev. Lett.* **112**, 061802 (2014).
- [8] K. A. Olive *et al.* (Particle Data Group Collaboration), *Chin. Phys. C* **38**, 090001 (2014); and <http://pdg.lbl.gov/2015/>.
- [9] K. Abe *et al.* (T2K Collaboration), *Phys. Rev. Lett.* **111**, 211803 (2013).
- [10] K. Abe *et al.* (T2K Collaboration), *Nucl. Instrum. Methods Phys. Res., Sect. A* **659**, 106 (2011).
- [11] K. Suzuki *et al.* (T2K Collaboration), *Prog. Theor. Exp. Phys.* (2015) 053C01.
- [12] K. Abe *et al.*, *Nucl. Instrum. Methods Phys. Res., Sect. A* **694**, 211 (2012).
- [13] Y. Fukuda *et al.*, *Nucl. Instrum. Methods Phys. Res., Sect. A* **501**, 418 (2003).
- [14] K. Abe *et al.* (T2K Collaboration), *Phys. Rev. D* **91**, 072010 (2015).
- [15] K. Abe *et al.* (T2K Collaboration), *Phys. Rev. Lett.* **116**, 181801 (2016).
- [16] K. Abe *et al.* (T2K Collaboration), *Phys. Rev. D* **87**, 012001 (2013); **87**, 019902 (2013).
- [17] A. Ferrari, P. R. Sala, A. Fasso, and J. Ranft, Reports No. CERN-2005-010, SLAC-R-773. and INFN-TC-05-11, 2005.
- [18] R. Brun, F. Bruyant, F. Carminati, S. Giani, M. Maire, A. McPherson, G. Patrick, and L. Urban, Report No. CERN-W5013, CERN, 1994.
- [19] C. Zeitnitz and T. A. Gabriel, *Nucl. Instrum. Methods Phys. Res., Sect. A* **349**, 106 (1994).

- [20] N. Abgrall *et al.* (NA61/SHINE Collaboration), *Eur. Phys. J. C* **76**, 84 (2016).
- [21] Y. Hayato, *Acta Phys. Pol. B* **40**, 2477 (2009). Version 5.3.2 of NEUT library is used, which includes (i) the multinucleon ejection model of Nieves *et al.* [23] and (ii) nuclear long-range correlations for CCQE interactions, treated in the random phase approximation [24].
- [22] J. Nieves, J. E. Amaro, and M. Valverde, *Phys. Rev. C* **70**, 055503 (2004); **72**, 019902(E) (2005).
- [23] C. Wilkinson *et al.*, *Phys. Rev. D* **93**, 072010 (2016).
- [24] A. A. Aguilar-Arevalo *et al.* (MiniBooNE Collaboration), *Phys. Rev. D* **81**, 092005 (2010).
- [25] A. A. Aguilar-Arevalo *et al.* (MiniBooNE Collaboration), *Phys. Rev. D* **88**, 032001 (2013).
- [26] G. Fiorentini *et al.* (MINERvA Collaboration), *Phys. Rev. Lett.* **111**, 022502 (2013).
- [27] L. Fields *et al.* (MINERvA Collaboration), *Phys. Rev. Lett.* **111**, 022501 (2013).
- [28] J. Nieves, I. R. Simo, and M. J. Vicente Vacas, *Phys. Rev. C* **83**, 045501 (2011).
- [29] R. Gran, J. Nieves, F. Sanchez, and M. J. Vicente Vacas, *Phys. Rev. D* **88**, 113007 (2013).
- [30] K. M. Graczyk and J. T. Sobczyk, *Phys. Rev. D* **77**, 053001 (2008); **79**, 079903(E) (2009).
- [31] C. Wilkinson, P. Rodrigues, S. Cartwright, L. Thompson, and K. McFarland, *Phys. Rev. D* **90**, 112017 (2014).
- [32] A. Higuera *et al.* (MINERvA Collaboration), *Phys. Rev. Lett.* **113**, 261802 (2014).
- [33] K. Abe *et al.* (T2K Collaboration), *Phys. Rev. Lett.* **117**, 192501 (2016).
- [34] M. Martini, M. Ericson, G. Chanfray, and J. Marteau, *Phys. Rev. C* **80**, 065501 (2009).
- [35] M. Martini and M. Ericson, *Phys. Rev. C* **87**, 065501 (2013).
- [36] M. Day and K. S. McFarland, *Phys. Rev. D* **86**, 053003 (2012).
- [37] P. Amaudruz *et al.* (T2K ND280 FGD Collaboration), *Nucl. Instrum. Methods Phys. Res., Sect. A* **696**, 1 (2012).
- [38] N. Abgrall *et al.* (T2K ND280 TPC Collaboration), *Nucl. Instrum. Methods Phys. Res., Sect. A* **637**, 25 (2011).
- [39] V. D. Barger, K. Whisnant, S. Pakvasa, and R. J. N. Phillips, *Phys. Rev. D* **22**, 2718 (1980).
- [40] K. Hagiwara, N. Okamura, and K.-i. Senda, *J. High Energy Phys.* **09** (2011) 082.
- [41] R. D. Cousins and V. L. Highland, *Nucl. Instrum. Methods Phys. Res., Sect. A* **320**, 331 (1992).
- [42] W. K. Hastings, *Biometrika* **57**, 97 (1970).
- [43] P. Adamson *et al.* (NO ν A Collaboration), *Phys. Rev. D* **93**, 051104 (2016).
- [44] P. Adamson *et al.*, *Phys. Rev. Lett.* **25**, 181801 (2013).
- [45] R. Wendell, *Proc. Sci.*, ICRC2015 (2015) 1062.
- [46] M. Aartsen *et al.*, *Nucl. Phys.* **B908**, 161 (2016).
- [47] G. J. Feldman and R. D. Cousins, *Phys. Rev. D* **57**, 3873 (1998).
- [48] K. Abe *et al.* (T2K Collaboration), <http://t2k-experiment.org/results/t2kdata-nu-antinu-joint-analysis-2016>.
- [49] L. Adamson *et al.* (NO ν A Collaboration), following Letter, *Phys. Rev. Lett.* **118**, 151802 (2017).
- [50] See Supplemental Material at <http://link.aps.org/supplemental/10.1103/PhysRevLett.118.151801> for the figures of the reconstructed energy for the far-detector muon neutrino candidate samples together with the best fit.