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1 Investigation of OSL surface exposure dating to reconstruct post-LIA 2 glacier fluctuation in the French Alps (Mer de Glace, Mont Blanc massif)

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9

10 Abstract

11 Providing quantitative constraints on late Pleistocene glacier fluctuations remains an important
12 challenge for understanding glacier response to past and future climate changes. In most mountainous
13 settings, paleo-glacier reconstructions are limited because they often lack precise temporal constraint.
14 Different geochronological methods have been developed and applied to date specific
15 geomorphological or sedimentological markers for paleo-glacier dynamics. Recently, OSL (optically
16 stimulated luminescence) surface exposure dating has been introduced and provides us with an
17 opportunity to improve paleo-glacier reconstructions. This method is based on the sensitivity of the
18 OSL signal from rock minerals to light, resulting in bleaching of the OSL signal within the upper first
19 millimeters of the exposed rock surface, a process that depends on the exposure age, the rock type and
20 the local setting (e.g. topographic shielding, bedrock orientation etc.). Here, we investigate the
21 potential of OSL surface exposure along a vertical cross-section of polished bedrock surfaces with
22 known post-LIA (Little Ice Age) exposure ages (from 3 to 137 years) along the Mer de Glace glacier
23 (Mont Blanc massif, France). The infra-red stimulated luminescence (IRSL) signals from rock slices
24 exhibit increasingly deep bleaching profiles with elevation and thus exposure age, which is consistent
25 with progressive glacier thinning since the LIA. Our results show that OSL surface exposure dating
26 can be applied to periglacial environments, and is a promising tool for high-resolution reconstruction
27 of ice-extent fluctuations, both in space and time.

28
29 **Keywords:** Optically stimulated luminescence (OSL), surface exposure dating, luminescence depth
30 profile, paleo-glacier reconstruction, Mont Blanc massif

31 1. Introduction

32 During the last ca. 5 Ma of the Earth's history, global climate cooled and evolved towards oscillating
33 climatic conditions that intensified towards the present (e.g. Zachos et al., 2001; Herbert et al., 2016).
34 This climate shift left a strong imprint on mountain topography (e.g. Penck, 1905; Broecker and
35 Denton, 1990; Molnar and Engand, 1990; Peizhen et al., 2001; Egholm et al., 2009). However,

36 understanding paleo-climatic conditions in mountainous areas over the Plio-Pleistocene epochs
37 remains difficult. Local records of successive glacial/interglacial cycles are scarce or poorly preserved
38 over such long timescales (Ehlers and Gibbard, 2007). Polar ice-sheets and marine cores are useful for
39 providing long-term global climatic records but are unable to describe regional continental climate. In
40 contrast, glaciers and their fluctuations through time provide invaluable information on past mountain
41 climatic conditions. Through mapping and dating moraine deposits and erratic boulders, it is possible
42 to reconstruct the history of ice-extent (e.g. for the European Alps: Ivy-Ochs et al., 2006; Bini et al.,
43 2009; Preusser et al., 2011; Schimmelpfennig et al. 2014; Ivy-Ochs et al., 2015; Wirsig et al., 2016).

44 Past glacier extents in the European Alps are well constrained since the Little Ice Age (LIA:
45 15th to 19th centuries). Using historical maps, survey reports and aerial photogrammetry, glacier
46 fluctuations have been precisely reconstructed over the last two centuries (e.g. Vincent et al., 2014).
47 To go further back in time into the Pleistocene, different geochronological methods can be used such
48 as lichenometry (Winkler et al., 2004), varve chronologies (Stewart et al., 2011), dendrochronology
49 (Baillie, 1995) and radiocarbon dating (Hajdas, 2008). However, organic matter can be scarce for
50 glacial/periglacial deposits because of extremely active geomorphic systems associated with glacial
51 environments. In addition to these methods, surface exposure dating of polished bedrock or erratic
52 boulders using terrestrial *in situ* cosmogenic nuclides has been developed over the last decades (Lal et
53 al., 1991; Gosse and Philips, 2001; Balco, 2011; Ivy-Ochs and Briner, 2014), and has been widely
54 used in the European Alps (see Ivy-Ochs et al., 2006; 2009 for reviews). The combination of different
55 cosmogenic nuclide pairs (e.g. ¹⁰Be and ¹⁴C: e.g. Goehring et al., 2012; Hippe et al., 2014) provides us
56 with important information on Alpine glacier paleogeography since the Last Glacial Maximum (LGM;
57 Ivy-Ochs et al., 2006; Wirsig et al., 2016). However, the cosmogenic nuclide production rate and the
58 integration of production over the first 1-2 meters below a rock surface may limit the resolution of
59 such methods for recent and/or complex exposure histories.

60 Here we investigate whether optically stimulated luminescence (OSL) surface exposure dating
61 can be used to reconstruct recent glacier fluctuation. Luminescence dating is based on the
62 accumulation of trapped electrons through time in the crystalline lattice of certain minerals (e.g. quartz
63 or feldspar). Some of these trapped electrons are sensitive to daylight exposure (Aitken, 1985; Huntley
64 et al., 1985). Luminescence dating is commonly used to date sediment burial in a range of
65 geomorphological environments (e.g. Duller, 2008; Rhodes, 2011; Fuchs and Owen, 2008) but can
66 also be used to determine rates of bedrock cooling (Guralnik et al., 2015; King et al., 2016; Brown et
67 al., 2017), and the exposure age of archaeological rock surfaces (Polikreti et al., 2003; Sohbaty et al.,
68 2011). This latter application is based on the principle that when a rock surface is exposed to light, the
69 luminescence signal, which is initially homogenous within the rock sample (at a given level or in field
70 steady-state; e.g. Valla et al., 2016), will progressively decrease at depth until being completely
71 zeroed, a phenomenon called “bleaching” (Aitken, 1998). The assumption used in this study is that the
72 longer a surface has been exposed to daylight, the deeper the signal bleaching will be (Polikreti et al.,

73 2002). In granitic and gneissic rocks, bleaching through time has been shown to occur over the first
74 few centimetres depth of the rock surface (Vafiadou et al., 2007; Sohbati et al., 2011; Freiesleben et
75 al., 2015). In alpine environments, glacier advances during the late Pleistocene to Holocene have been
76 associated with subglacial erosion of bedrock at the centimetre-scale (e.g. Goehring et al., 2011). This
77 means that only the most recent exposure history of the bedrock will be recorded, as earlier exposure
78 histories and OSL bleaching evidence will have been eroded by subsequent glacier advances. OSL
79 surface exposure dating would thus in theory enable past glacier extents to be reconstructed with a
80 high temporal resolution for both recent and complex exposure histories. Furthermore, this method is
81 attractive because of the short time required for sample preparation (Sohbati et al., 2011), although
82 one current disadvantage is the requirement for calibration of this chronometer on rock surfaces with
83 independently known exposure ages (Sohbati et al., 2012a).

84 In the following, we first introduce the study site, i.e., the Mer de Glace, and our sampling
85 strategy. We have targeted several independently dated glacially eroded bedrock surfaces, which
86 represent past elevations of the glacier surface since the LIA. We then review the basic principles of
87 the method and present the luminescence signals for six different surfaces along a vertical cross
88 section above the present-day Mer de Glace. Our results show a strong correlation between sample
89 elevation, exposure age and bleaching depth. Finally, we use this dataset to show that model
90 calibration requires multiple samples of known age to take full advantage of OSL surface exposure
91 dating in both glaciated and formerly glaciated environments.

92 **2. Setting and sampling strategy**

93 **2.1. Geomorphological setting**

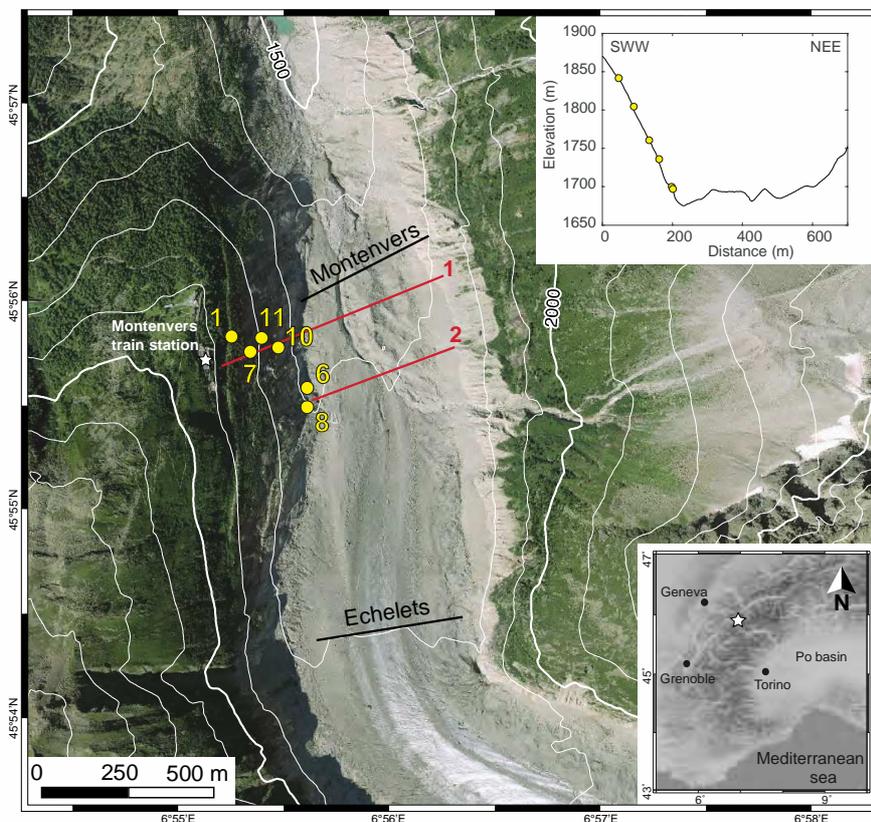
94 The Mer de Glace glacier (Fig. 1) is about 11.5 km long and is located in the Mont Blanc massif. The
95 modern glacier covers an area of 30.4 km² (excluding former tributary Talèfre Glacier) and spans an
96 elevation range from 4205 m to 1531 m.a.s.l. (data from 2008; Gardent et al., 2014). The mean
97 equilibrium line altitude (ELA), reconstructed using remote sensing methods, was about 2880 m.a.s.l.
98 between 1961 and 1990 for five of the main north-facing Mont Blanc massif glaciers, including the
99 Leschaux Glacier for the period 1984-2010 (Rabatel et al., 2013).

100 The Mer de Glace is an appropriate laboratory for validating the application of OSL surface
101 exposure dating for paleo-glacier reconstruction. Numerous studies have provided detailed
102 reconstructions of Mer de Glace fluctuations from the LGM towards the Holocene and present day
103 (Coutterand and Buoncristiani, 2006; Nussbaumer et al, 2007; Vincent et al., 2014; LeRoy et al.,
104 2015). The Montenvers site (Fig. 1) was chosen as an optimal study site as the evolution of the glacier
105 thickness since the LIA has been reconstructed by Vincent et al. (2014) using historical maps, aerial
106 photogrammetry and satellite-derived digital elevation models (see Section 2.3 for details).
107 Furthermore, the rock type is generally homogenous along the valley flank (i.e. orthogneiss; Dobmeier

108 et al., 1998), avoiding any lithological dependency of the OSL surface exposure dating approach
 109 although occasional granitic lenses are exposed in the lower part of the profile (see Section 2.2 for
 110 details).

111 Our sampling strategy was to collect glacially polished bedrock surfaces with the best-
 112 preserved erosion patterns (glacial striations, roches moutonnées; Fig. 3) to ensure that sample
 113 bleaching profiles reflect the period of time since post-LIA deglaciation. The samples were also
 114 selected to have low topographic shielding and vegetation cover (e.g. lichen). Steep slopes were
 115 selected (i.e. above 30°) to limit any potential snow cover effects. In particular, we focused on rock
 116 surfaces exhibiting striations parallel to the Mer de Glace flow line in order to avoid the potential
 117 influence of tributary glaciers.

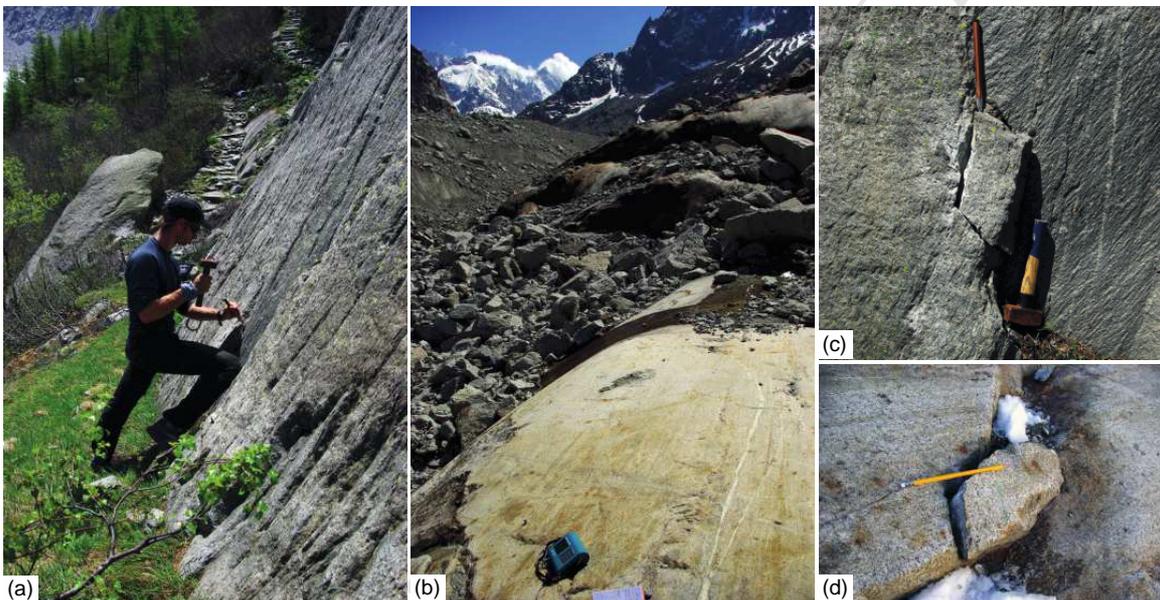
118



119
 120 **Figure 1:** Sampling map of the Montanvers site, Mer de Glace. The orthorectified aerial photograph
 121 of the Mer de Glace was acquired in 2016 (source: www.geoportail.gouv.fr). The black lines show the
 122 two cross-sections produced by Vincent et al. (2014) which we interpolated to reconstruct glacier
 123 surface elevations at two different locations (red lines 1 and 2, see Section 2.3 and Supplementary Fig.
 124 A1) where samples were collected (yellow dots with numbers). Upper right inset represents the
 125 collected samples projected along cross-section 1. Bottom right inset shows location of the study area
 126 within the western Alps.

127 **2.2 Sample description**

128 We collected six samples along the Monteners profile during several field campaigns (2015-2016),
 129 ranging in elevation from 1841 to 1696 m.a.s.l. (Fig. 1 and Table 1). Samples MBMV1, MBMV7,
 130 MBMV8, MBMV10 and MBMV11 consist of coarse-grained orthogneiss, typical for the Aiguilles
 131 Rouges massif (Dobmeier et al., 1998). These rocks mainly comprise coarse K-feldspar crystals,
 132 quartz, biotite and muscovite. Only MBMV6 was collected from a granitic lens, which consists of
 133 bigger quartz and feldspar crystals than the orthogneiss (Fig. 3b). Because differences in crystals
 134 properties may influence light penetration, i.e. due to both crystal size and distribution, sample
 135 MBMV6 is used to explore any potential lithological effect on the OSL surface exposure dating
 136 approach.

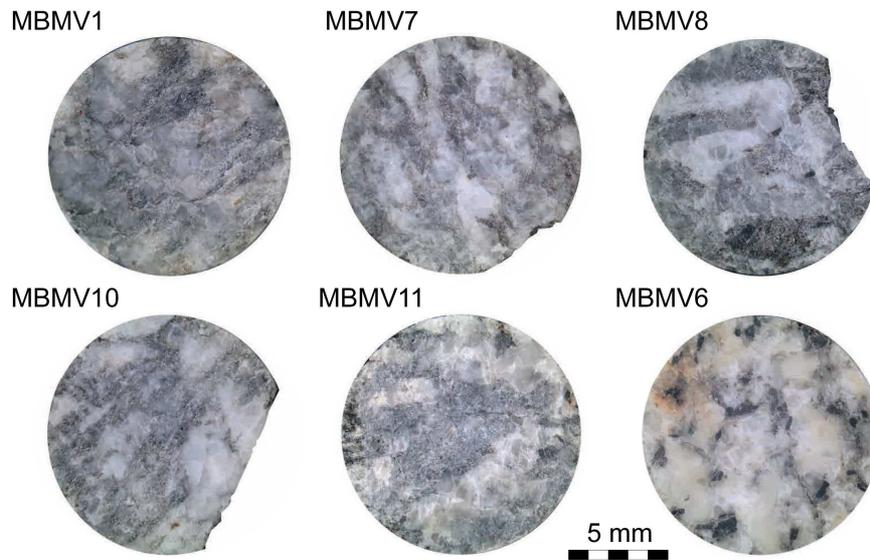


137 (a) (b) (c) (d)
 138 **Figure 2:** Sampling sites and sample details at the Monteners site. (a-b) Sampling sites for MBMV7
 139 and MBMV6. (c-d) Outcrops and samples MBMV7 and MBMV6.

140
 141 **Table 1:** Sample characteristics from the Monteners cross-section. Estimated exposure ages were
 142 reconstructed using differential GPS and ice-thickness reconstruction as shown in section 2.3.
 143 Shielding factors were calculated with the geometric shielding calculator (CRONUS-Earth project).
 144 Note that all estimated exposure ages are referenced from 2015 (date of the first field campaign).

Sample ID	Latitude	Longitude	Elevation	Lithology	Estimated Exposure Age	Topographic	Surface
	WGS 84		[m.a.s.l.]		[Year before 2015]	shielding factor	orientation
MBMV1	45°55'54.0"	06°55'07.7"	1841	Gneiss	137	0.81	N8 55°E
MBMV6	45°55'48.9"	06°55'17.7"	1696	Granite	2	0.92	N0 30°E
MBMV7	45°55'52.7"	06°55'09.9"	1804	Gneiss	69	0.79	N374 60°E
MBMV8	45°55'47.7"	06°55'18.5"	1699	Gneiss	3	0.81	N13 54°E
MBMV10	45°55'54.0"	06°55'14.1"	1735	Gneiss	18	0.79	N0 60°E
MBMV11	45°55'54.3"	06°55'11.5"	1760	Gneiss	30	0.88	N355 75°E

145
146



147

148 **Figure 3:** Pictures of selected rock slices (see section 3.1 for details), showing the difference in
149 composition and texture between orthogneiss (MBMV1, MBMV7, MBMV8, MBMV10, and
150 MBMV11) and granite (MBMV6).

151 2.3 Independent age calibration from glacier thickness reconstruction

152 We use the historical post-LIA reconstruction of the Mer de Glace thickness (Vincent et al., 2014)
153 available for two cross-sections: Montenvers and Echelets (Fig. 1). Samples MBMV1, MBMV7,
154 MBMV10 and MBMV11 were collected from the same profile located 290 m upstream of the
155 Montenvers cross-section and 690 m downstream of the Echelets cross-section (cross-section 1, Figs.
156 1 and 4). Samples MBMV6 and MBMV8 were taken along a profile (cross-section 2, Fig. 1) located
157 200 m upstream of the cross-section 1. Because glacial thinning would progressively expose bedrock
158 surfaces at lower elevations, we can use the relationship between exposure age and sample elevation
159 to constrain the temporal evolution of glacial thickness (Fig. 1 and 4).

160 Post-LIA thickness reconstructions of the Mer de Glace for cross-sections 1 and 2 have been
161 interpolated from the Montenvers and Echelets cross-sections. Exposure ages from 2 to 137 years
162 were obtained for the different samples, using either cross-section 1 (MBMV1, MBV7, MBMV10 and
163 MBMV11) or cross-section 2 (MBMV6 and MBMV8) (see Supplementary Material A1). All
164 exposure ages are relative to the first sampling campaign in summer 2015.

165 3. Methodology: OSL surface exposure dating

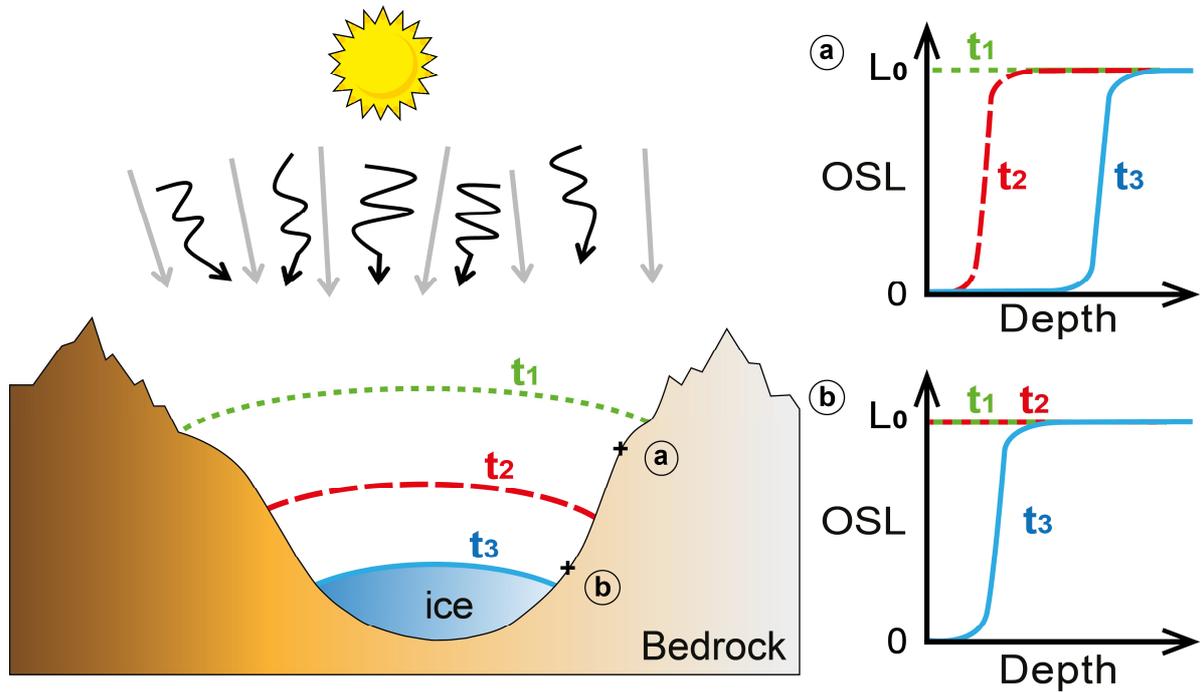
166 3.1. Theoretical approach

167 Minerals such as quartz and feldspar naturally contain defects or impurities in their crystal lattice.
168 Energy released by ambient radiation (i.e. cosmic rays and the flux of high-energy solar particles
169 or/and radioactive decay in the rock-matrix) excites electrons from their equilibrium state (valence
170 band), and these can become trapped at higher energy levels within the crystal. Because of the finite
171 number of traps, electron filling occurs until saturation is reached. By giving energy to the system in
172 the form of light or heat (natural bleaching processes), electrons are released and return to their
173 equilibrium state, producing photons. This phenomenon is called luminescence and the intensity of a
174 given luminescence signal is thus proportional to the number of trapped electrons (Aitken, 1985;
175 1998). In a rock surface continuously exposed to daylight, the progressive bleaching of the
176 luminescence signal is expected to propagate deeper into the surface with time (Habermann et al.,
177 2000; Polikreti et al., 2002; Laskaris and Liritzis, 2011).

178 Rock surface dating was first used in archaeology, and was based on thermally-stimulated
179 measurements, i.e. thermoluminescence (TL; Liritzis et al., 1994; Richards et al., 1994; Theocaris et
180 al., 1997; Polikreti et al., 2002; 2003). More recently, optically stimulated luminescence dating (OSL;
181 e.g. Habermann et al. 2000; Vafiadou et al., 2007) has been introduced to date surface exposure,
182 which benefits from improved measurement reproducibility and more rapid signal bleaching following
183 exposure to daylight than typically-used TL signals (e.g. the 325°C TL peak in quartz). The potential
184 of OSL for dating exposure events in geomorphological (Freiesleben et al., 2015; Sohbaty et al., 2015)
185 and archaeological (Liritzis, 2011) contexts has recently been investigated, and a range of applications
186 including relative sea-level changes and coastal geomorphology (Simms et al., 2011; Simkims et al.,
187 2013) have been published. However, OSL surface exposure dating has not yet been applied to
188 glacially polished bedrock surfaces.

189 In mountainous environments, OSL dating can be used to evaluate the exposure age of a
190 polished bedrock surface as described in Figure 4. At the initial condition (t_1 in Fig. 4), the glacier has
191 reached its maximum thickness. Ice and periglacial sediments cover the bedrock surface, and the
192 luminescence signals of bedrock minerals are in field steady-state and uniform in the rock column.
193 When the glacier retreats, freshly-eroded surfaces are exposed to daylight (point **a** at time t_2 , Fig. 4).
194 The initial luminescence signals start to bleach for these exposed surfaces, while the sample at lower
195 elevation is still covered by the glacier and its luminescence signals remain uniform in the rock (point
196 **b** at time t_2 , Fig. 4). As the glacier continues to thin, the lower part of the bedrock flanks are uncovered
197 (t_3 , Fig. 4) and the luminescence signals start to bleach for the lower-elevation surfaces. Therefore, in a
198 setting affected by progressive glacier retreat and thinning, there is a direct correlation between the
199 elevation of the studied site and the exposure age, with the assumption that the longer a surface is

200 exposed to daylight, the deeper into the rock the luminescence signal is bleached (Freiesleben et al.,
 201 2015; Sohbati et al., 2011).



202
 203 **Figure 4:** Sketch linking glacier thinning and OSL signal evolution for two bedrock surfaces located
 204 at different elevations along the valley flank. Straight arrows (grey) represent cosmic rays and high-
 205 energy solar particle flux; this radiation, together with radioactive decay in the rock matrix build up
 206 the latent luminescence signal. Other arrows (black) represent low energy electromagnetic radiation
 207 from the sun; this radiation bleaches the latent luminescence signal. At the initial time t_1 , the glacier is
 208 at its maximum extent and the OSL signals for both surfaces are in field steady-state and uniform
 209 within the rocks, L_0 . At time t_2 , the glacier has retreated and exposed the surface (a), the OSL signal
 210 begins to bleach whilst surface (b) remains covered with its luminescence signal unchanged. In the
 211 final step t_3 , the glacier size has shrunk, surface (a) remains exposed and its OSL signal is bleached at
 212 greater depth while surface (b) has just been exposed to daylight and its OSL signal has been bleached
 213 just below the exposed surface.

214 3.2. Modelling approach

215 To assess rock surface exposure durations to daylight from a luminescence depth profile, we use the
 216 model proposed by Sohbati et al. (2011; 2012a,b) who provide an in-depth review of each parameter.
 217 When a rock surface is exposed to daylight, both detrapping (due to the release of energy by daylight)
 218 and trapping (due to absorption of energy from ambient radiation) occur simultaneously. The trapped-
 219 charge concentration during light exposure is given by the following differential equation:

220

$$221 \frac{\partial n(x,t)}{\partial t} = -E(x) n(x,t) + F(x) [N(x) - n(x,t)] \quad (1)$$

222

223 Where $n(x, t)$ is the trapped charge concentration [m^{-3}] at time t [s] and depth x [m], $N(x)$ is the
 224 concentration of sites [m^{-3}] available for trapping at depth x , $E(x)$ is the charge detrapping rate [s^{-1}],
 225 and $F(x)$ is the trap filling rate [s^{-1}]. The charge detrapping rate, $E(x)$, is itself given by:

226

$$227 \quad E(x) = \overline{\sigma\varphi_0} e^{-\mu x} \quad (2)$$

228

229 where $\varphi_0(\lambda, x)$ is the photon flux [$\text{cm}^{-2} \text{s}^{-1}$] describing the rate of incoming photons that can bleach the
 230 trap of interest. $\sigma(\lambda)$ is the photoionization cross section [cm^2] describing the probability of this
 231 specific trap to be excited by light stimulation. It is averaged over the wavelengths present in the solar
 232 spectrum at the surface ($x = 0$). Here, we assume that the photon flux does not fluctuate through time,
 233 and we are only concerned with the product of the two parameters, which is given by $\overline{\sigma\varphi_0}$ [s^{-1}] (i.e. the
 234 effective decay rate of luminescence; Sohbati et al., 2011). Equation (2) also includes a decay term for
 235 light attenuation with depth. The light attenuation coefficient μ [m^{-1}] describes how deep into the rock
 236 a photon will penetrate and affect the luminescence signal. μ is assumed to be independent of
 237 wavelength in the spectral range of interest (Sohbati et al., 2011).

238 For surface exposure dating of terrestrial surfaces, the effect of trap filling during daylight exposure
 239 over short timescales (i.e. centuries) is often negligible (i.e. $F(x) \approx 0$) (see Supplementary Material
 240 A2). The trapped charge population at a given depth (x) can then be approximated by:

241

$$242 \quad n(x) = n_0 e^{-E(x)t} \quad (3)$$

243

244 where n_0 is the initial charge population [m^{-3}] assumed to be constant with depth within the rock
 245 column prior to bleaching. Assuming that the luminescence signal (L) is proportional to n , Eq. (3)
 246 becomes:

247

$$248 \quad L = \frac{Lx}{Tx} = L_0 e^{-\overline{\sigma\varphi_0}t} e^{-\mu x} \quad (4)$$

249

250 where Lx/Tx is the normalized natural luminescence signal measured at depth x [m] after exposure
 251 age t [s]. L_0 is the normalized natural luminescence signal before bleaching (Fig. 4), which is sample
 252 dependent and can be constrained in the laboratory.

253

254 Equation 4 can predict the rock luminescence profiles for different exposure ages, however the
 255 mean photon flux φ_0 , the photoionization cross-section σ , and the attenuation coefficient μ must first
 256 be quantified. φ_0 is mainly controlled by the latitude and the cloudiness; and it is broadly correlated to
 257 elevation (Blumthaler et al., 1997). The solar irradiance is fluctuating over decadal timescales (Lean,
 1987) making the independent determination of the photon flux impossible without knowing the time

258 of exposure. The photoionization cross-section σ is depending on both the mineral and the trap
 259 targeted (Bailey, 2004). For samples coming from the same region and from a similar lithology, $\overline{\sigma\varphi_0}$
 260 is assumed to be uniform and μ is expected to be of the same order of magnitude between samples, but
 261 not necessarily equal.

262 The OSL-depth profile of exposed rock surfaces with independently constrained exposure
 263 durations can be used to calibrate the $\overline{\sigma\varphi_0}$ and μ parameters by fitting the luminescence signal
 264 bleaching with depth (Singarayer, 2002; Sohbati et al., 2012a). These constrained parameters can then
 265 be used to determine the exposure histories of unknown-age surfaces from the same region.

266 Here, our objective is to demonstrate the validity of the proposed model (Eq. 4) on polished
 267 bedrock surfaces and to calibrate the model parameters on surfaces with known exposure age. To do
 268 so, the unknown $\overline{\sigma\varphi_0}$ and μ parameters are inverted for each sample using a probability density
 269 function of the model parameters, given the observed OSL-depth profile data. This includes a least
 270 absolute deviation regression \mathcal{L}_{sample} (Eq. 5), in which we randomly prescribed a range of different
 271 $\overline{\sigma\varphi_0}$ and μ values. From the residual likelihood \mathcal{L}_{sample} obtained, we select the maximum likelihood
 272 values of $\overline{\sigma\varphi_0}$ and μ . The modelled luminescence signals $\left(\frac{Lx}{Tx}\right)_m$ are calculated for each rock slice of a
 273 given sample using the known exposure age of each sampling site, giving:

$$275 \quad \mathcal{L}_{sample} = \exp\left(-\frac{1}{2a} \sum_{i=1}^n \left| \left(\frac{Lx}{Tx}\right)_{obs}^{(i)} - \left(\frac{Lx}{Tx}\right)_m^{(i)} \right| \right) \quad (5)$$

276
 277 where n is the number of rock slices per sample, $\left(\frac{Lx}{Tx}\right)_m^{(i)}$ is the luminescence signal calculated using
 278 Eq. (4), $\left(\frac{Lx}{Tx}\right)_{obs}^{(i)}$ is experimentally measured for each rock slice i and a is the uncertainty. Given that
 279 the scatter of the plateau signal (L_0) for every independent sample is larger than the analytical error,
 280 we use the standard deviation around the plateau value L_0 to estimate a . Then, we compute the
 281 combined likelihood for a number of samples p using:

$$283 \quad \mathcal{L}_{combined} = \prod_{j=1}^p \mathcal{L}_{sample(j)} \quad (6)$$

284
 285 This approach provides the most likely common values of $\overline{\sigma\varphi_0}$ and μ . Once the parameters of the
 286 model are determined as shown above, it is possible to invert the exposure age for other rock surfaces
 287 using the constrained $\overline{\sigma\varphi_0}$ and μ values (cf. Eq. (5)).

288
 289 **Table 2:** Summary of symbols.

290

Symbol	Unit	Definition
n	m^{-3}	Concentration of trapped charge
x	m	Depth
t	s	Time
N	m^{-3}	Concentration of sites available too trap charge
E	s^{-1}	Charge detrapping rate due to solar radiation
F	s^{-1}	Charge trapping due to ionising radiation
σ	cm^2	Photionisation cross-section
φ	$\text{cm}^{-2} \text{s}^{-1}$	Photon flux
$\overline{\sigma\varphi}_n$	s^{-1}	Charge detrapping rate
μ	m^{-1}	Attenuation coefficient
Lx	Counts	Regenerated luminescence signal
Tx	Counts	Test dose signal
L	Counts	Luminescence

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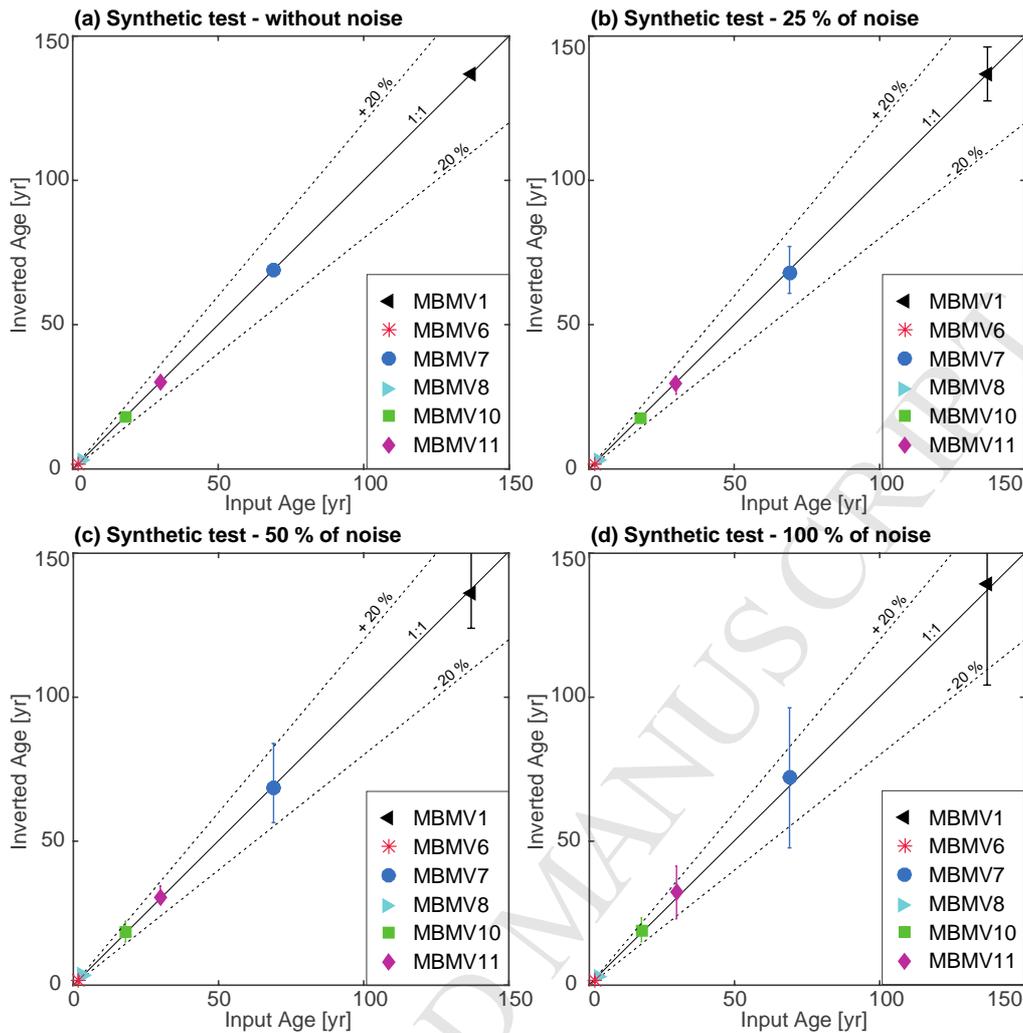
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In order to verify our modelling approach, we show a synthetic inversion. We produce a synthetic luminescence signal (Lx/Tx for depths in between 0 and 14 mm) using Eq. (4) and sample-specific $\overline{\sigma\varphi}_0$ and μ parameters (obtained from initially fitting every sample using their independent age control, see Section 5.2 for details) and assuming a constant μ value (i.e. homogenous lithology with rock depth). The first step of the synthetic test is to invert parameters $\overline{\sigma\varphi}_0$ and μ knowing the exposure age t for each individual sample as presented above. Then, using these $\overline{\sigma\varphi}_0$ and μ parameters, we subsequently invert for the exposure age t using Eq. (4). In order to study the effect of potential uncertainties from the experimental data on the exposure age determination, we reproduce this synthetic test with white noise on the luminescence signal, with four different amplitudes between 0 and 100 % (Fig. 5). Our synthetic results show that our inversion approach can recover the exposure age with 0 to 50% noise. The synthetic test with 100% noise on the luminescence signal provided age outcomes with larger uncertainties (>20%). The best results are obtained using the best-fit of $\overline{\sigma\varphi}_0$ and μ and the median value of the predicted exposure ages. The resulting uncertainties are correlated with the magnitude of the noise, however any potential variability in the luminescence signal does not appear to produce a significant bias on the inverted exposure age.



308
 309 **Figure 5:** Results (median value) of inverted exposure age from the synthetic test, (a) without noise on
 310 the luminescence signal, (b) with 25%, (c) 50%, and (d) 100% noise. Error bars represent $\pm 2\sigma$ on the
 311 inverted age.

312 4. Sample preparation and analysis

313 The bedrock samples were cored to 30 mm depth using a Husqvarna DM220 drill, with 10-mm
 314 diameter. Cores were then sliced into 0.7-mm thick rock slices with a BUEHLER IsoMet low speed
 315 saw equipped with a 0.3-mm thick diamond blade. The samples were drilled and sliced under wet
 316 conditions (water and lubricant, respectively) to avoid any heating that could potentially reset the OSL
 317 signal. Sample preparation was done under subdued red light conditions. The thickness of each rock
 318 slice was measured to determine the precise depth of each luminescence measurement.

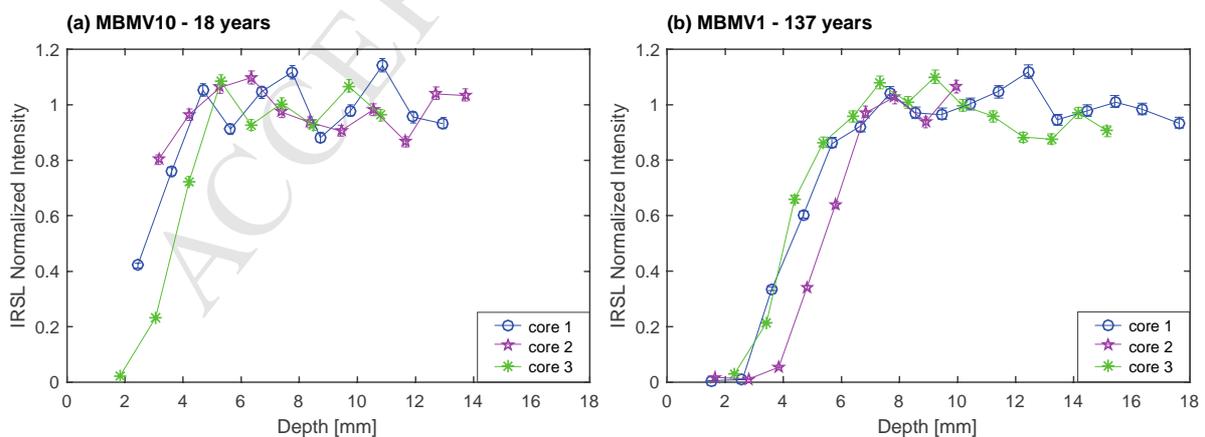
319 All luminescence measurements were performed using Risø TL-DA 20 TL/OSL luminescence
 320 readers (Bøtter-Jensen et al., 2010) equipped with ^{90}Sr beta sources at the University of Lausanne
 321 (Switzerland). The readers have dose rates of ~ 0.1 and $\sim 0.2 \text{ Gy s}^{-1}$ and measurement reproducibility of
 322 1.14 % and 1.26 % respectively. We first perform a preheat at 250 °C before giving infrared (IR)
 323 stimulation (870 nm, FWHM 40 nm) at 50 °C. Luminescence signals are detected through a filter

324 combination of a Schott BG-3 and Schott BG-39. A uniform test dose was used (27.2 Gy) to measure
 325 the subsequent luminescence response (T_x) and to normalize the natural infrared stimulated
 326 luminescence (IRSL) signal (L_x) for every rock slice. Infrared stimulated luminescence was measured
 327 for 200 s and signals were integrated over the first 6 seconds whereas the background signal was
 328 integrated between 70-100 seconds. Measurements were analysed using Analyst v.3.22b (Duller,
 329 2005). All thermal treatments and stimulations at temperatures greater than 200°C (i.e. preheat step)
 330 were carried out in nitrogen atmosphere. The experimental approach was validated using a dose
 331 recovery and preheat plateau test (see Supplementary Material A3; Murray and Wintle, 2000; Wintle
 332 and Murray, 2006).

333 5. Results

334 5.1. Experimental results

335 Figure 6 shows the luminescence measurements for representative samples MBMV1 and MBMV10
 336 (results of the others samples are presented in Fig. 8). Three replicates (i.e. individual cores) per
 337 sample were sliced in a way that a depth and an IRSL signal can be attributed to each rock slice
 338 (unique colour/symbol for each individual rock slice in Fig. 6). The results show similar behaviour
 339 between the different cores for a given sample (Fig. 6). The IRSL signal is bleached near the surface
 340 and reaches a plateau at depth. Furthermore, and more importantly, the transition from a bleached
 341 signal to the plateau varies with the exposure age. The three core measurements reproduce well for
 342 both samples illustrated in Figure 6, with the mean standard deviation between the three cores ranging
 343 from 7 to 27% for all the studied samples. These results confirm experimentally that cores extracted
 344 from one individual sample record the same exposure history, supporting the proposed approach.
 345



346
 347 **Figure 6:** Infrared stimulated luminescence (IRSL) signal with depth for samples (a) MBMV10 and
 348 (b) MBMV1. Each coloured data point represents an individual rock slice. IRSL signals were
 349 normalized by L_0 , which was determined by taking the average of the luminescence measurements

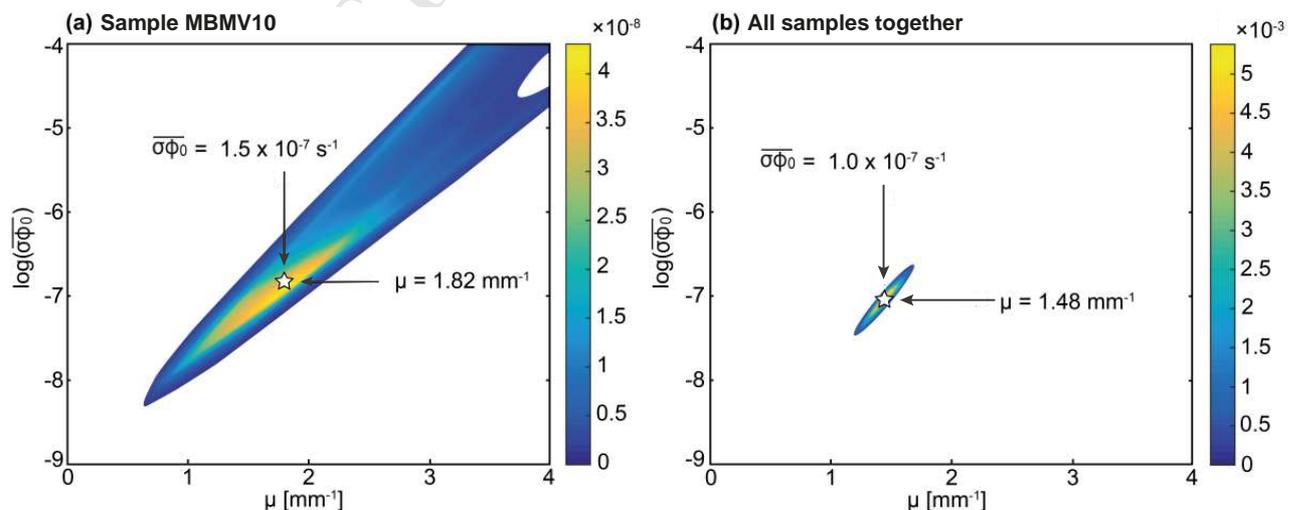
350 along the plateau. The plateau was defined when the luminescence signal is fluctuating by less than
 351 20%.

352 5.2. Independent parameter determination

353 In this section, we determine the $\overline{\sigma\phi_0}$ and μ parameters individually for each sample in order to study
 354 their potential variability from one rock surface to another (Table 3). As explained in Section 3.2,
 355 bedrock surfaces from the same location should share a common $\overline{\sigma\phi_0}$ parameter (i.e. same order of
 356 magnitude; Blumthaler et al., 1997). Similarly, we expect that the μ parameter should be similar for
 357 samples from a uniform lithology. The determined parameters are then used in the inversion of the
 358 exposure ages for each sample individually (see Section 5.5). All inversion outcomes are summarized
 359 in Table 3. Samples MBMV1, MBMV8, MBMV10 and MBMV11 share similar effective decay rates
 360 ($\overline{\sigma\phi_0}$) with the same order of magnitude (from $1.4 \cdot 10^{-8}$ to $2 \cdot 10^{-7} \text{ s}^{-1}$) and show attenuation coefficients
 361 (μ) between 1.07 and 1.89 mm^{-1} . Samples MBMV6 and MBMV7 behave differently with much lower
 362 effective decay rates ($\overline{\sigma\phi_0}$ of $2.0 \cdot 10^{-6}$ and $4.2 \cdot 10^{-6} \text{ s}^{-1}$, respectively), and different attenuation
 363 coefficients (μ of 0.92 and 2.50 mm^{-1} , respectively).

364 5.3. Parameter determination from joint probability estimates

365 We evaluate now the parameter determination from joint probability estimates in order to illustrate the
 366 benefit of having several known-age calibration samples. Figure 7a presents modelled results for
 367 sample MBMV10, which is representative of the other samples (except MBMV6 and MBMV7, see
 368 Section 5.2). The results show that $\log(\overline{\sigma\phi_0})$ and μ co-vary, which we attribute to measurement
 369 uncertainties and variability between the different cores. Figure 7b shows the area of acceptable fits
 370 when all the gneissic samples are included (i.e. excluding the granitic sample MBMV6).
 371



372
 373

374 **Figure 7:** Relationship between the $\overline{\sigma\varphi_0}$ and μ parameters (a) for sample MBMV10, and (b) for all of
 375 the gneiss samples (i.e. excluding MBMV6) enabling determination of the shared $\overline{\sigma\varphi_0}$ and μ
 376 parameters ($1.0 \cdot 10^{-7} \text{ s}^{-1}$ and 1.48 mm^{-1} respectively). For both figures, the colour scale shows the
 377 likelihood between modelled and experimental data (Eq. 6, note the differences in scaling between the
 378 two panels), and the star is the best-fit parameter values. Zero probability is not shown for clarity.

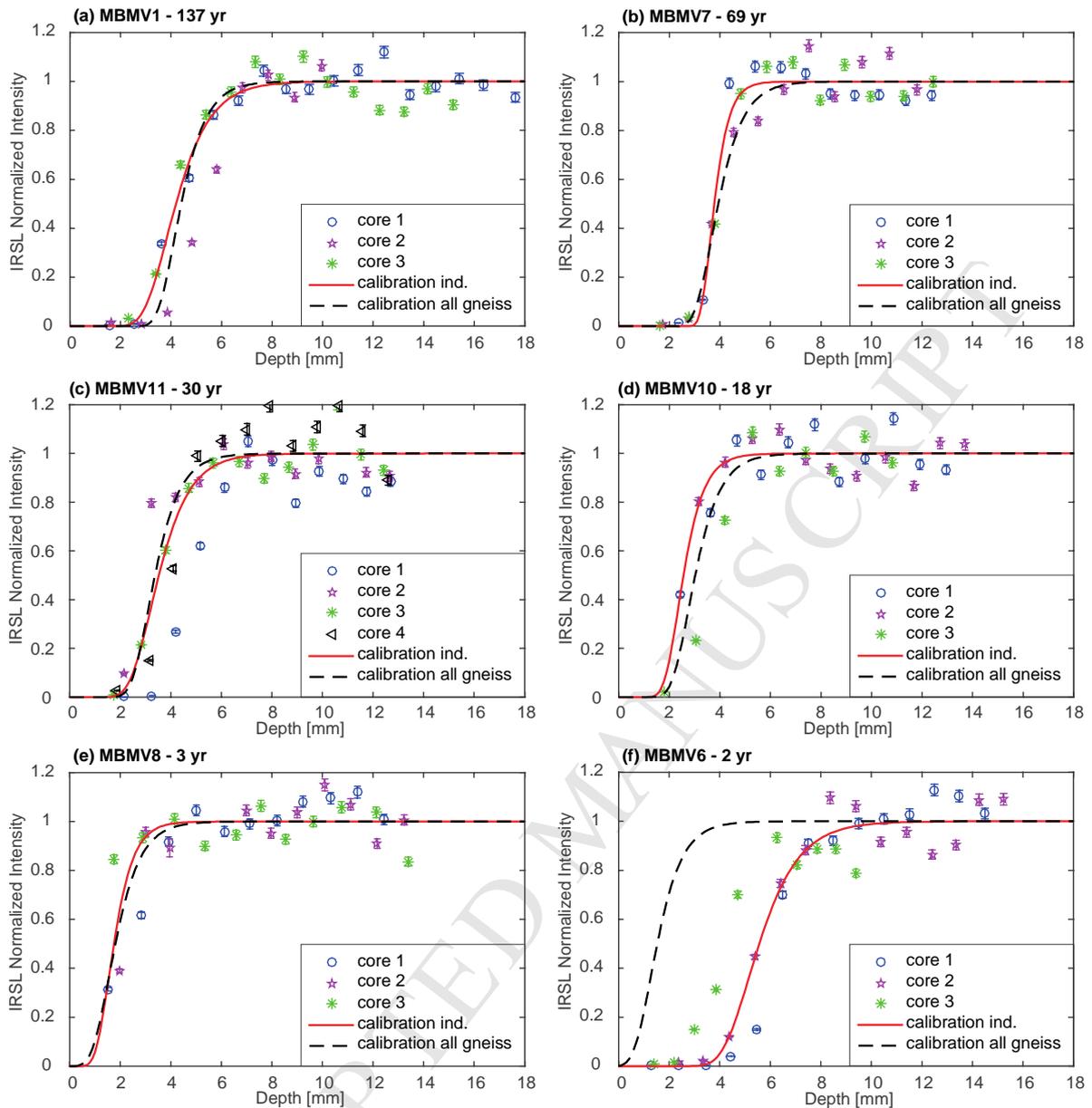
380 We then contrasted individual estimates of $\overline{\sigma\varphi_0}$ and μ for each sample, using different
 381 combinations of samples to estimate these parameters. The results are summarized in Tables 3 and A3.
 382 When inverting the model parameters with any combinations of three samples, all estimates of the
 383 effective decay rates are between $6.6 \cdot 10^{-8}$ and $1.4 \cdot 10^{-7} \text{ s}^{-1}$, and all estimates of the attenuation
 384 coefficients are between 1.33 and 1.57 mm^{-1} . Combinations of four samples provide $\overline{\sigma\varphi_0}$ values
 385 ranging from $7.2 \cdot 10^{-8}$ to $1.2 \cdot 10^{-7} \text{ s}^{-1}$ and μ values between 1.38 and 1.53 mm^{-1} . According to Table 3
 386 and Figure 7b, a common likelihood exists for all the gneissic samples calibrated together, giving an
 387 effective decay rate of $1.0 \cdot 10^{-7} \text{ s}^{-1}$ and an attenuation coefficient of 1.47 mm^{-1} .

388 Figure 8 depicts the normalized IRSL signals measured for all samples and their individually-
 389 constrained best-fit models (red lines) as described previously and illustrated in Figure 7a. The
 390 obtained outcomes show that the proposed model accurately describes the luminescence bleaching
 391 process through depth and time. The best-fit model calibrated with all of the gneissic samples together
 392 (black dashed lines, parameters in Table 3) fits close to the best-fit model determined for each sample
 393 individually (except MBMV6). These results confirm a key objective of the study, which is the
 394 possibility to calibrate the model parameters using different surfaces along a vertical profile, with the
 395 same lithology and different (independently-determined) exposure ages.

397 **Table 3:** Best-fit values of $\overline{\sigma\varphi_0}$ and μ determined for every sample individually and for all samples
 398 excluding MBMV6. Combinations of three or four samples are presented in Table A3 (Supplementary
 399 Material).

Parameter	$\overline{\sigma\varphi_0} [\text{s}^{-1}]$			$\mu [\text{mm}^{-1}]$		
	Best-fit	+1 σ	-1 σ	Best-fit	+1 σ	-1 σ
Samples						
Individually						
MBMV1	$1.4 \cdot 10^{-8}$	$1.5 \cdot 10^{-8}$	$1.3 \cdot 10^{-8}$	1.07	1.08	1.05
MBMV6	$2.0 \cdot 10^{-6}$	$2.2 \cdot 10^{-6}$	$1.8 \cdot 10^{-6}$	0.92	0.95	0.92
MBMV7	$4.0 \cdot 10^{-6}$	$5.0 \cdot 10^{-6}$	$3.6 \cdot 10^{-6}$	2.52	2.56	2.46
MBMV8	$2.2 \cdot 10^{-7}$	$2.4 \cdot 10^{-7}$	$1.5 \cdot 10^{-7}$	1.89	1.98	1.70
MBMV10	$1.5 \cdot 10^{-7}$	$1.5 \cdot 10^{-7}$	$1.1 \cdot 10^{-7}$	1.82	1.87	1.75
MBMV11	$4.2 \cdot 10^{-8}$	$5.3 \cdot 10^{-8}$	$3.9 \cdot 10^{-8}$	1.21	1.22	1.13
All samples together excluding MBMV6	$1.0 \cdot 10^{-7}$	$1.1 \cdot 10^{-7}$	$9.5 \cdot 10^{-8}$	1.48	1.50	1.44

401



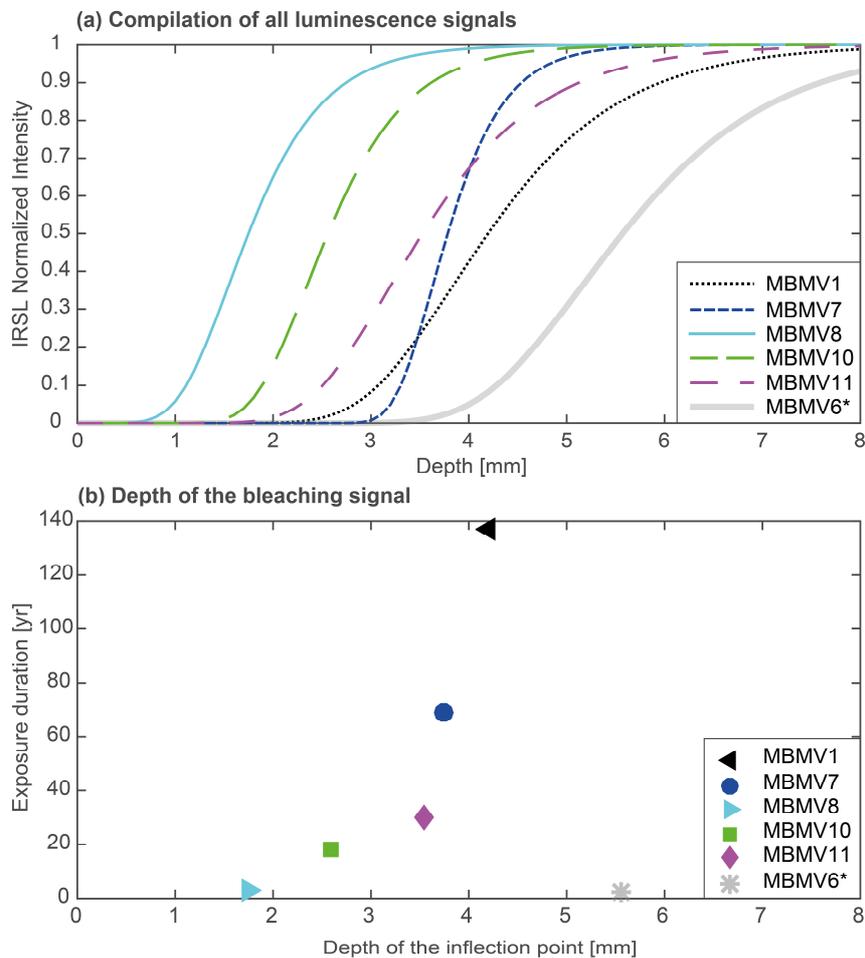
402

403 **Figure 8:** Normalized infrared stimulated luminescence (IRSL) profiles with depth and best-fit
 404 models. Coloured data point represents individual rock slice (each symbol/colour represents one core).
 405 The red lines show the best-fit model for each sample taken individually. The dashed black lines
 406 represent the best-fit model from a common calibration of the parameters using all gneiss samples
 407 together ($\overline{\sigma\varphi_0} = 1.0 \cdot 10^{-7} \text{ s}^{-1}$ and $\mu = 1.48 \text{ mm}^{-1}$, cf. Fig. 7b and Table 3). Raw IRSL data are presented
 408 in Table A3(Supplementary Material).

409 5.4. Evolution of the luminescence signal through time

410 Compiling the best-fit models determined for each sample individually, a positive correlation between
 411 the exposure age and the depth at which the natural IRSL signal is zeroed can be clearly observed for
 412 samples within the same lithology (Fig. 9a). If we consider the inflection point of each individual
 413 model ($Lx/Tx = 0.5$ on Fig. 9a) as a proxy for the bleaching depth, this value ranges between 1.7 and

414 4.2 mm for 3 and 137 years of daylight exposure, respectively (Fig. 9b). The granitic sample MBMV6
 415 does not follow this correlation, its bleaching depth being at 7 mm after 2-yr exposure to daylight.
 416



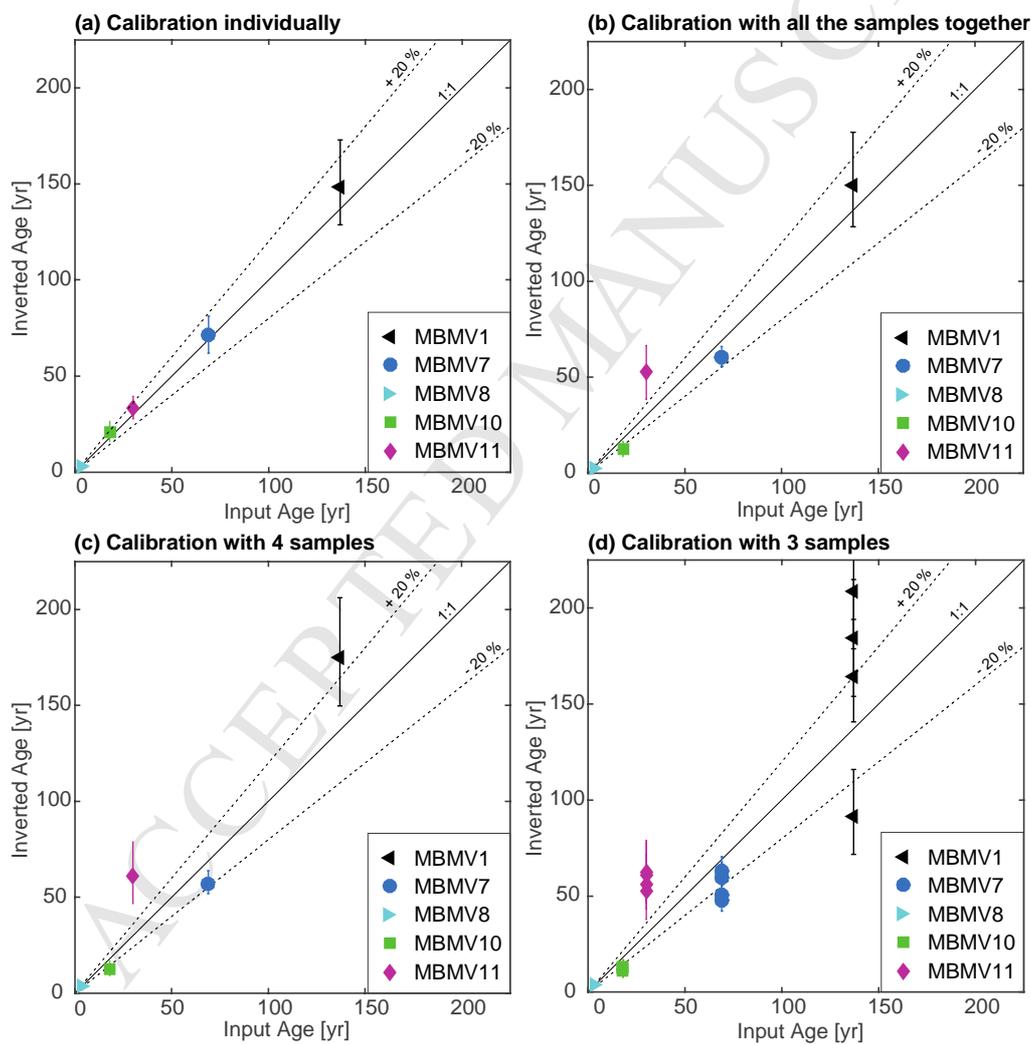
417 **Figure 9:** (a) Compilation of the best-fit models for each individual sample (cf. red lines in Fig. 8). (b)
 418 Correlation between the IRSL bleaching depth (i.e. the inflection point of the models presented in (a))
 419 and the exposure age of each individual sample. The * symbol indicates the granitic sample
 420 (MBMV6); all the other samples are gneiss.
 421

422 5.5. Inversion for exposure age

423 Once the model parameters have been determined by different sample combinations, it is possible to
 424 subsequently invert the exposure age as explained above. Figure 10 compares the exposure ages
 425 inverted from the different sample combinations, with the observed exposure age (all results are
 426 compiled in the Supplementary Table A4 and Figure A3). Figure 10a shows that our modelling
 427 approach is able to recover the observed exposure ages using parameters determined for each
 428 individual sample (<10% difference). When the exposure ages are inverted using the parameters
 429 determined for all of the gneissic samples together (as shown in Fig. 7b), there are slight differences
 430 between the inverted exposure age and independent age control (Fig. 10b, Table A4 and Figure A3).

431 The inverted ages are almost all within 20% of the observed ages except for sample MBMV11, which
 432 is overestimated by 90%.

433 Taking different calibration combinations with four (Fig. 10c) or three (Fig. 10d) samples also
 434 results in different performance regarding age predictions. For all gneissic samples, except MBMV11,
 435 the inverted exposure ages at 2σ are still within 20% of the observed ages. Note that in our approach
 436 the inverted exposure ages with four and three samples calibrations are only shown when the specific
 437 sample is not part of the calibration combination (grey shadow in the Supplementary Table A4).
 438 Although the match between the inverted and observed ages, as well as the trend between samples, is
 439 preserved independent of the calibration approach, our results show that the higher the number of
 440 calibration sites is, the better the inversion of exposure ages would be.
 441



442
 443 **Figure 10:** Correlation between inverted (median values) and observed exposure ages resulting from
 444 different calibration combinations to constrain the model parameters. The error bars on the inverted
 445 exposure ages are $\pm 2\sigma$ as presented in Table 3 (all results presented in the Supplementary Table A4).

446 **6. Discussion**

447 Our results from the Mer de Glace glacier have allowed to validate, over post-LIA timescales (i.e. over
448 2-137 years), the assumption that the longer a rock surface has been exposed to daylight, the deeper
449 the luminescence signal has been bleached (Polikreti et al., 2002; 2003; Sohbati et al., 2011; 2012).
450 Using the mathematical model propose by Sohbati et al. (2011), we accurately describe the time
451 evolution of luminescence within a rock column. The different combinations of samples used to
452 calibrate the model give parameter values ($\overline{\sigma\phi_0}$ and μ) that are on the same order of magnitude for
453 samples within similar regions and lithology, and which agree with literature values (Sohbati et al.,
454 2011; 2012a,b).

455 We also observe that the evolution of luminescence signals with both time and depth within
456 bedrock is mainly controlled by rock characteristics (lithology, texture, weathering and mineral
457 composition). These rock properties will govern the light attenuation and penetration into rocks
458 (parameter μ in Eq. 4), and thus the net bleaching effect on the luminescence signal. At the regional
459 scale, the lithology should preferably be uniform to enable model calibration on some known-age
460 surfaces (through independent dating) before application to reconstruct the exposure history of other
461 bedrock surfaces with unknown exposure age. We see that in a granitic rock, comprising coarse quartz
462 and feldspar grains (translucent minerals), the luminescence-bleaching front will propagate much
463 faster than in gneiss bedrock.

464 Our inversion approach to constrain rock surface exposure ages from OSL data, reveals that
465 the number of calibration samples is critical for constraining the model parameters and thus obtaining
466 accurate exposure ages. Fortunately, calibration rock surfaces in periglacial environments can often be
467 found from historical or remote-sensing paleo-glacier reconstructions. Other types of bedrock surfaces
468 can be used for independent constraint, e.g. anthropogenic structures such as road-cut outcrops (e.g.
469 Sohbati et al., 2012a) or landslide scars. The combined investigation of OSL systems with other
470 surface exposure dating methods such as terrestrial *in situ* cosmogenic nuclides will also enable us to
471 quantitatively assess the method's accuracy over longer timescales such as the late Pleistocene.

472 Experimental luminescence data presented in Figure 8 confirms that each individual sample's
473 exposure history has been recorded in its luminescence depth profile. For the six bedrock surfaces
474 studied here, each luminescence profile exhibits a fully-bleached signal at shallow depth (i.e. from 1 to
475 7 mm depending on both the exposure age and lithology, Fig. 9), followed by a sharp transition to a
476 plateau of intensity deeper into the rock. These simple and homogeneous luminescence profiles can be
477 compared with complex profiles previously observed following multi-stage exposure histories
478 obtained from buried cobbles (Freiesleben et al., 2015; Sohbati et al., 2015). This confirms that the
479 glacially-polished surfaces we sampled along the Montenvers cross-sections have experienced a
480 simple exposure history. Furthermore, field evidence for surface preservation with glacial features
481 (striations, flutes) indicate that the bedrock surfaces have been eroded and polished by subglacial

482 processes before deglaciation. Weathering or mechanical erosion may lead to an underestimation of
483 the true exposure age. Thereby, the inferred exposure history from these well-preserved rock surfaces
484 can be used to reconstruct the paleo-glacier thickness and extent since the LIA.

485 Bleaching of the OSL signal has occurred at less than 1 cm depth below the exposed surface
486 after more than 137 years of daylight exposure, highlighting the high temporal resolution of this novel
487 method for paleo-glacier reconstruction. In mountainous locations such as the Mont Blanc massif,
488 where the glacial history has been complex with several glacier fluctuations during the late Pleistocene
489 to Holocene (recurrent retreat/advance cycles; e.g. LeRoy et al., 2015), the application of absolute
490 dating methods such as terrestrial *in situ* cosmogenic nuclides are difficult due to potential inheritance
491 from previous exposure events (e.g. Goehring et al., 2011). One of the main advantages of OSL
492 surface exposure dating is that daylight bleaching of the OSL signal occurs within the first few
493 millimetres below the exposed rock surface. Short glacier re-advances over the late Holocene (e.g.
494 LeRoy et al., 2015) would have easily eroded the first centimetres of bedrock, consequently resetting
495 the OSL system before the post-LIA glacier retreat. We have thus shown in this study that well-
496 preserved polished bedrock surfaces can be used for the application of OSL surface exposure dating in
497 order to constrain the timing of the last glacial retreat from the LIA to present day, improving our
498 temporal resolution for glacier reconstruction. Over such timescales, the contribution of the trap filling
499 rate ($F(x)$ in Section 2.3) from radioactive decay in gneissic or granitic rock can be assumed to be
500 negligible (see Supplementary material A2). However, this contribution may have to be taken into
501 account when extending paleo-glacier reconstruction using OSL surface exposure dating to longer
502 timescales, e.g. since the Last Glacial Maximum or further back into the Quaternary. Over the same
503 timescales, weathering and erosion of the surface are likely to play a significant role.

504 7. Conclusions

505 In this study, we have investigated the potential of OSL surface exposure dating for quantitatively
506 reconstructing post-LIA glacier retreat. We worked along an altitudinal cross-section of the Mer de
507 Glace glacier (Mont Blanc massif, France), and collected glacially-polished bedrock surfaces with
508 known exposure ages (from 3 to 137 years) along the Montenvers profile from around 1841 m.a.s.l.
509 elevation to the present-day glacier position (1696 m.a.s.l.). We have developed a statistical approach
510 to calibrate the bleaching model parameters from known-age samples. Experimental IRSL depth-
511 profile data for five different polished bedrock surfaces show an increase of the luminescence signal
512 bleaching depth with exposure age. We conclude that OSL surface exposure dating can be applied to
513 glacial and periglacial environments, and is a promising tool for high-resolution reconstruction of
514 recent ice-extent and thickness fluctuations, both in space and time. However, we find that several
515 calibration samples must be used to calibrate the model parameters before inferring exposure ages on
516 bedrock surfaces within a specific area, taking into account the potential variation in bedrock

517 lithology. We also find that measurement uncertainties, intrinsic data noise or both can result in large
518 uncertainties on inverted ages, especially when applying this method over 10^3 - 10^4 yr timescales.

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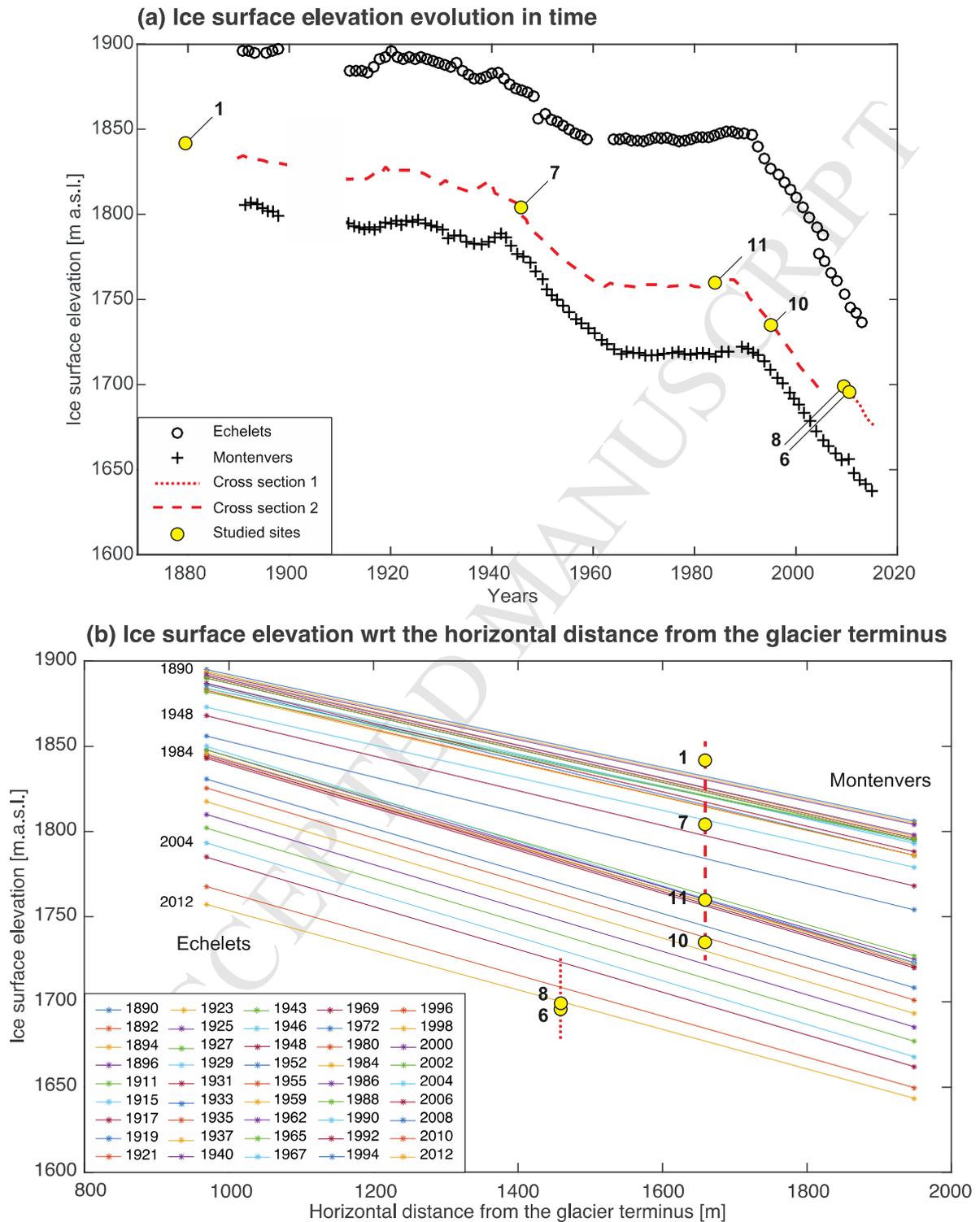
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- 679

680 **Supplementary Material**

681

682 **A1. Ice surface reconstruction**

683

684 **Figure A1:** (a) Reconstruction of averaged ice-surface elevation [m.a.s.l.] through time along the Mer

685 de Glace glacier (see locations of cross-sections on Fig.1). Averaged ice-surface elevations at the

686 Monteners (crosses) and Echelets (circles) cross-sections. The cross-sections 1 and 2 (red dashed

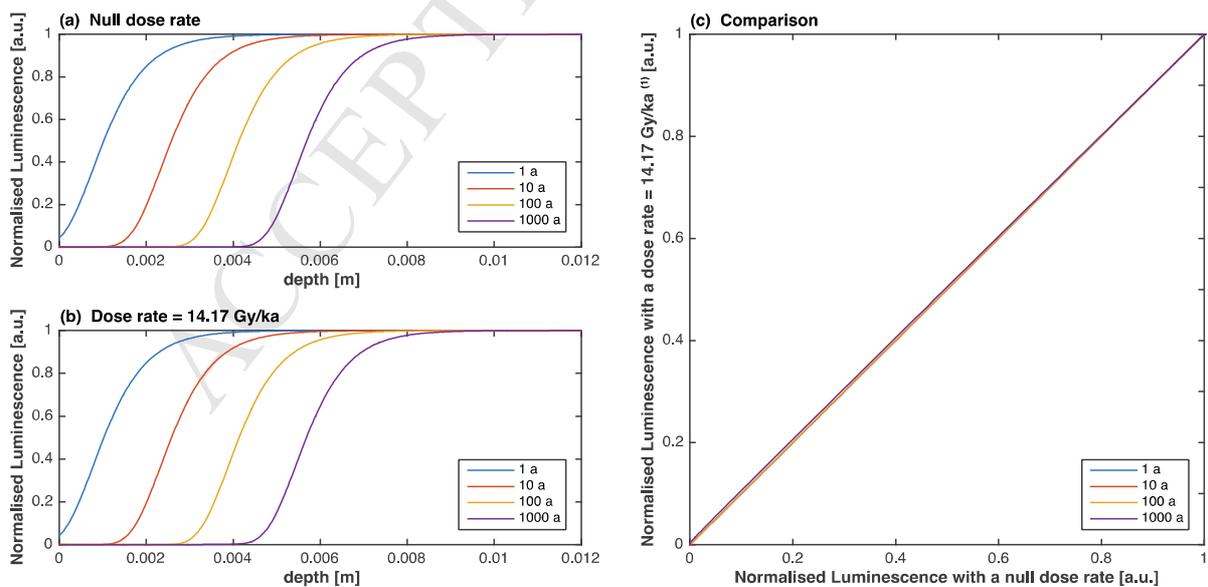
687 lines) have been interpolated from the Montenvers and Echelets cross-sections, and used to project the
 688 studied samples (yellow circles). (b) Ice-surface elevation with respect to the horizontal distance from
 689 the glacier terminus used for the interpolation of the Montenvers and Echelets cross-sections. The ice
 690 surface elevations have been reconstructed from historical maps, survey reports and aerial
 691 photogrammetry (modified from Vincent et al., 2014). The dataset was kindly provided by the French
 692 glacier observatory GLACIOCLIM (<http://www-igge.ujf-grenoble.fr/ServiceObs/index.htm>).
 693

694 A2. Dose rate sensitivity

695 The sensitivity of luminescence signal evolution to the dose rate is tested after four different daylight
 696 exposure times (1, 10, 100 and 1000 years of exposure), with a null dose rate and an extremely high
 697 dose rate ($\dot{D} \approx 14 \text{ Gy ka}^{-1}$, King et al., 2016). We used an equation developed by Sohbaty et al. (2012)
 698 describing the luminescence evolution $L(x)$ as a function of the exposure time t [s], depth x [mm],
 699 charge detrapping rate $\bar{\sigma}\bar{\varphi}_0$ [s^{-1}], attenuation factor μ [mm^{-1}], a sample-dependent constant that
 700 characterises filling rate D_0 [Gy] and the natural dose rate \dot{D} [Gy s^{-1}].

$$L(x) = \frac{\bar{\sigma}\bar{\varphi}_0 e^{-\mu x} e^{-t[\bar{\sigma}\bar{\varphi}_0 e^{-\mu x} + \frac{\dot{D}(x)}{D_0}] + \frac{\dot{D}(x)}{D_0}}}{\bar{\sigma}\bar{\varphi}_0 e^{-\mu x} + \frac{\dot{D}(x)}{D_0}} \quad (\text{A1})$$

701
 702 The resulting comparison shows that the luminescence signal is not sensitive to dose rate over
 703 millennial timescales. We thus consider the dose rate as negligible for our applications of OSL surface
 704 exposure dating, and do not take it into account in the luminescence evolution equation.
 705
 706



707
 708 **Figure A2:** Evolution of the normalised luminescence signal through time and depth for 1, 10, 100 and
 709 1000 years of daylight exposure, taking into account (a) a null dose rate and (b) an extremely high

710 dose rate of ~14 Gy/ka (King et al., 2016). Inset (c) shows the comparison between the result with two
 711 different dose rates.

712

713 **A3. Luminescence measurement tests**

714 The purpose of the following tests is to find the most appropriate infrared stimulated luminescence
 715 (IRSL) measurement conditions for analysis of the collected samples from the Montenvers site. We
 716 first performed a residual dose determination. Under natural daylight conditions, luminescence signals
 717 of feldspar may not be completely reset, leaving a residual dose. The residual test allows the
 718 evaluation of this remaining natural dose (which may also originate from other sources e.g. thermal
 719 transfer). We first reset the luminescence signal by exposing rock slices (3 slices for a representative
 720 sample of Montenvers site) to daylight for about 3 hours before analysing both slide sides. We then
 721 measured the residual dose using infrared stimulation at 50°C (IRSL₅₀) following different preheat
 722 temperatures (during 60 s). The residual dose after a preheat temperature of 250°C is 0.25 ± 0.45 Gy.
 723 With preheat temperature equal to 275°C, the residual dose is 0.85 ± 0.43 Gy. For preheat temperature
 724 equal to 300°C and 325°C, the residual doses increase to 2.94 ± 0.41 Gy and 2.10 ± 0.52 Gy
 725 respectively.

726 We then proceeded to a dose recovery test with preheat-plateau to determine the most
 727 appropriate preheat temperature. Thereby we quantified the recovered doses with IRSL₅₀ for the same
 728 range of preheat temperatures explored in the residuals test. We analyzed 3 rock slices with a
 729 laboratory beta dose of 27.25 Gy after complete optical bleaching (both disk sides exposed to daylight
 730 for about 3 hours). The samples were not heated prior the daylight bleaching. Preheat temperatures
 731 250°C, 275°C, 300°C, 325°C were investigated, and results are corrected for the residual dose values
 732 reported in Table A1. For preheat temperatures of 250°C, 275°, 300°C and 325°C, we obtained dose
 733 recovery ratios of 0.90 ± 0.10 , 0.87 ± 0.17 , 0.77 ± 0.23 and 0.85 ± 0.15 , respectively (Table A1). The
 734 optimal preheat temperature for both the residual dose and dose recovery is thus 250°C, and was used
 735 in all subsequent experiments.

736 **Table A1:** Results of the residual test and the dose recovery preheat plateau test after a given beta
 737 dose of 27.25 Gy.

738

Preheat Temperature (°C)	250	275	300	325
Residual dose (Gy)	0.25 ± 0.45	0.85 ± 0.43	2.94 ± 0.41	2.1 ± 0.52
Dose recovery ratio	0.9 ± 0.10	0.87 ± 0.17	0.77 ± 0.23	0.85 ± 0.15

739

740

741 **Table A2:** Sensitivity corrected luminescence signal intensities with depth. The depth x (cm) is
 742 measured during core slicing with a high-precision numerical micrometre. IRSL measurements

752

753 **Table A4:** Inversion results for exposure age using the different calibration combinations of bedrock

754 samples. Grey shading shows the inverted results for a specific sample when not included in the

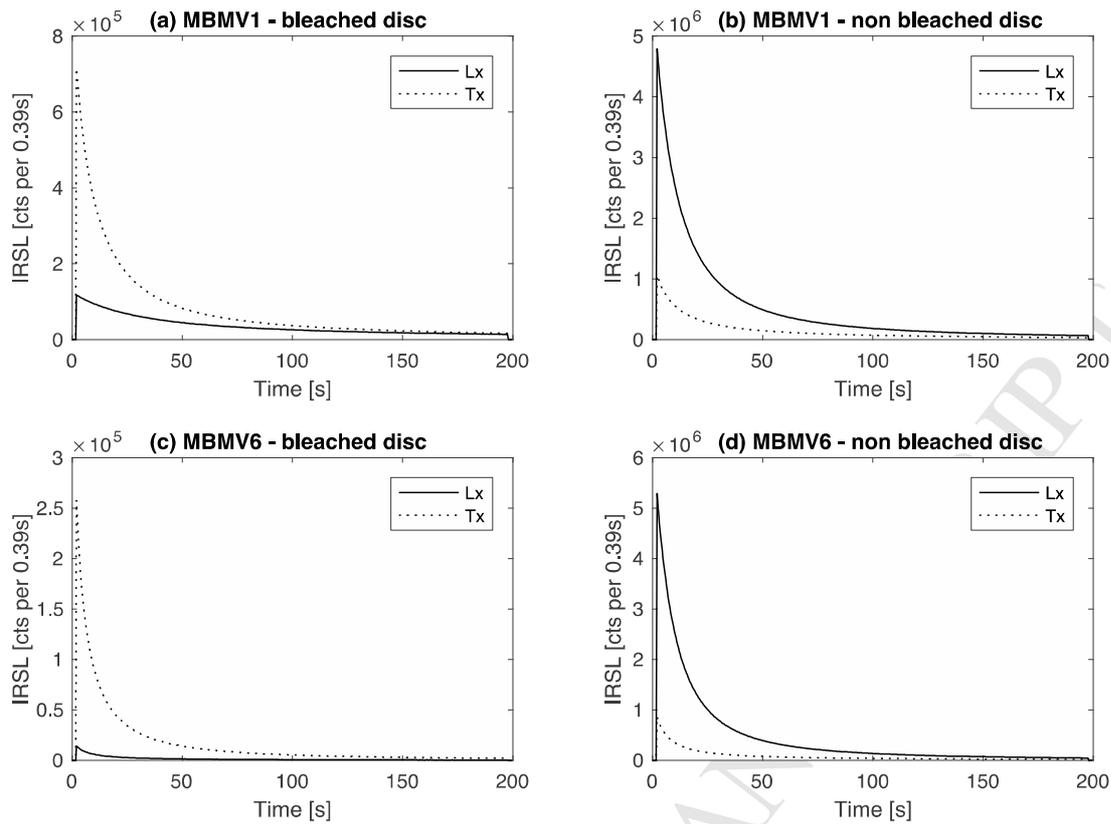
755 calibration combination (i.e. the exposure age of the specific sample has not used for the calibration of

756 the model). These results are used to produce Figures 10c-d.

757

Calibration combination	MBMV1 - 137 yrs				MBMV7 - 69 yrs				MBMV8 - 3 yrs				MBMV10 - 18 yrs				MBMV11 - 30 yrs						
	Best fit	Median	+2 σ	-2 σ	Best fit	Median	+2 σ	-2 σ	Best fit	Median	+2 σ	-2 σ	Best fit	Median	+2 σ	-2 σ	Best fit	Median	+2 σ	-2 σ			
Individually	146	152	172	126	89	79	88	67	2	3	3	2	19	19	23	14	50	42	50	34			
with all the sample excluding MBMV6	164	161	185	132	61	63	68	57	3	3	4	3	17	14	18	9	62	57	70	41			
with 3 samples, MBMV...																							
1	7	8	184	166	190	136	56	62	67	56	3	3	3	2	9	13	17	9	53	54	69	40	
1	7	10	194	183	213	146	77	80	88	71	5	5	5	4	21	18	23	13	90	69	88	50	
1	7	11	118	126	149	103	56	56	61	50	3	3	4	3	15	12	16	9	57	50	61	35	
1	8	11	122	135	159	104	61	63	70	56	3	4	4	3	13	15	18	10	55	53	68	39	
1	8	10	106	113	134	92	50	50	55	45	3	3	3	2	14	12	15	8	50	43	55	31	
1	10	11	174	150	181	119	67	71	80	63	4	4	5	4	17	17	21	12	44	61	77	44	
7	8	10	181	211	242	173	78	74	82	64	3	3	3	2	16	13	17	9	78	65	78	48	
7	8	11	165	162	189	136	66	62	67	56	3	3	3	2	14	13	17	9	40	54	68	40	
7	10	11	206	189	216	155	65	65	73	57	3	3	3	2	15	12	15	8	65	58	70	43	
8	10	11	116	93	116	71	46	50	55	43	3	3	4	3	11	12	15	9	52	42	52	29	
with 4 samples, MBMV...																							
1	7	8	10	176	170	194	139	63	67	73	61	4	4	4	3	10	15	19	10	68	61	75	43
1	7	8	11	177	152	174	125	63	59	64	54	3	3	3	3	15	13	17	9	63	52	66	38
1	7	10	11	129	129	154	98	58	62	69	54	4	4	4	3	13	14	18	10	38	52	66	38
1	8	11	10	117	134	155	106	58	58	64	52	4	3	4	3	16	13	17	9	42	52	64	36
7	10	11	8	159	168	196	141	63	63	69	56	2	3	3	2	16	13	16	9	52	55	69	41

758



759

760 **Figure A3:** Natural infrared stimulated signal (solid line, Lx) and test dose (27.25 Gy) subsequent
 761 luminescence response (dashed line, Tx) for a bleached signal (surface disc) (a) and (b) and for non-
 762 bleached signal (inside core disc) (b) and (c). (a) and (b) are IRSL signal representative for gneissic
 763 lithology (sample MBMV1). (c) and (d) for granitic lithology (sample MBMV6).

764 **References**

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