

# Advancing the orbit model for Galileo satellites during eclipse seasons

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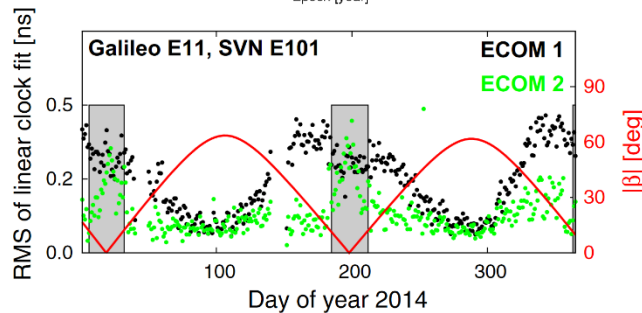
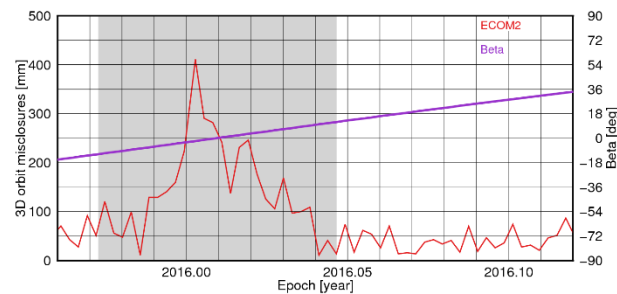
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# Motivation

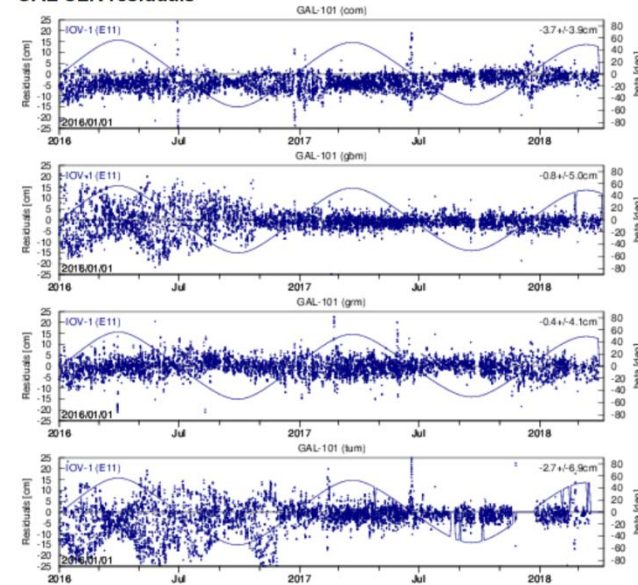
- Poor (Galileo) orbit modelling during eclipse seasons using Empirical CODE Orbit Model (ECOM2; Arnold et. al., 2015):
  - elevated orbit misclosures at day boundaries;
  - artifacts in SLR residuals at low  $\beta$  angles;
  - elevated RMS of linear clock fits during eclipses.

Gal E11



(Prange et al., 2017)

GAL SLR residuals



([http://mgex.igs.org/analysis/slres\\_GAL.php](http://mgex.igs.org/analysis/slres_GAL.php))

# Possible reasons

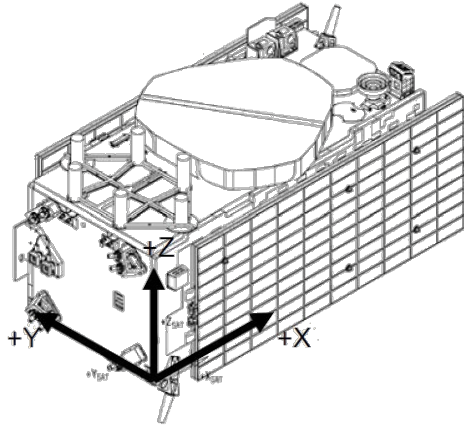
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- **Incorrect modelling of satellite attitude (nominal instead of the “true”).**
  - **Corrected thanks to the metadata of Galileo IOV and FOC satellites published by GSA.**
- **Insufficient SRP model parameterization.**
  - **More demanding to the modelling due to low satellite weight, but reasonably solved by ECOM2.**
- **Thermal effects are not fully absorbed (e.g., during eclipse seasons).**
  - **All empirical (SRP) parameters are switched off in eclipses.**

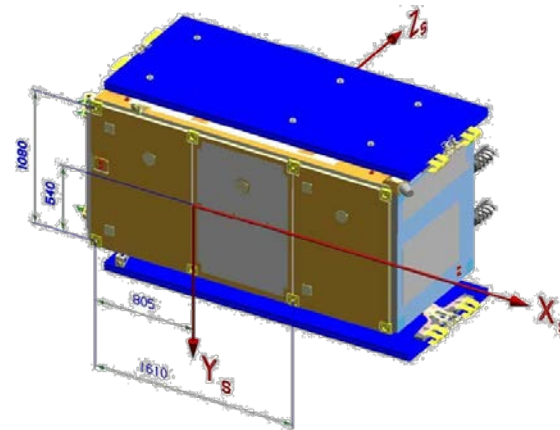
# Galileo Satellites

- From the metadata\* published by GSA:
  - thermal radiators on +X, +Y, -Y, -Z (FOC only) faces of the satellite body;
  - Galileo satellite mass  $\sim 700$  kg.

IOV



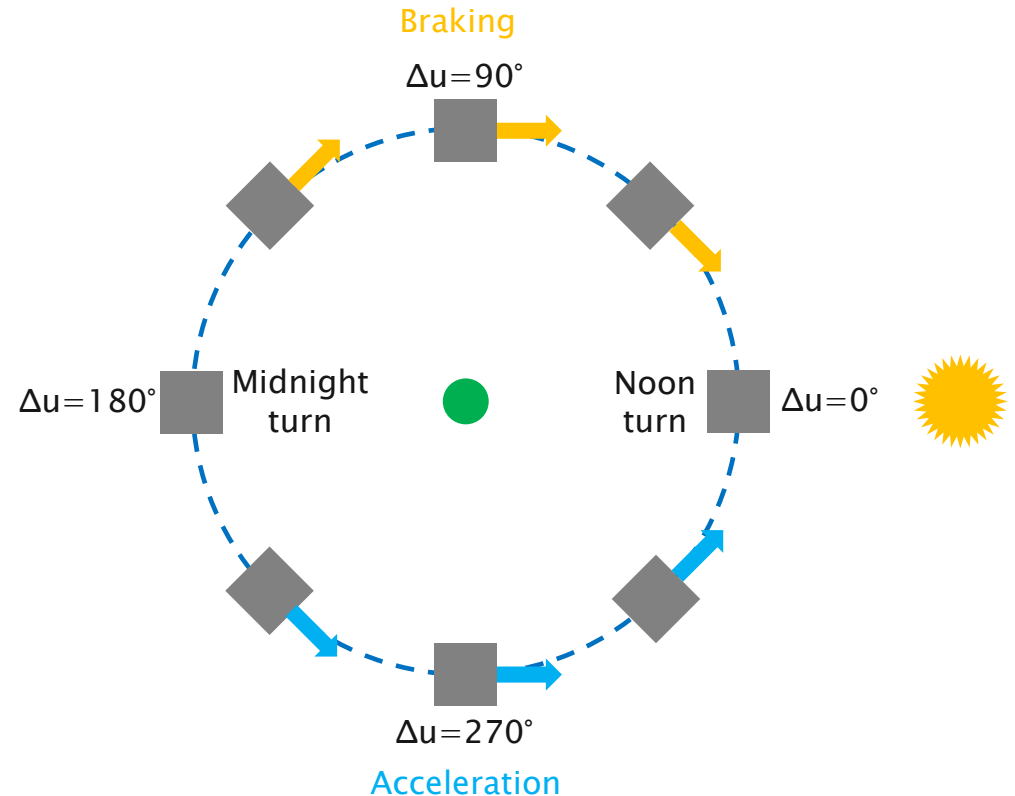
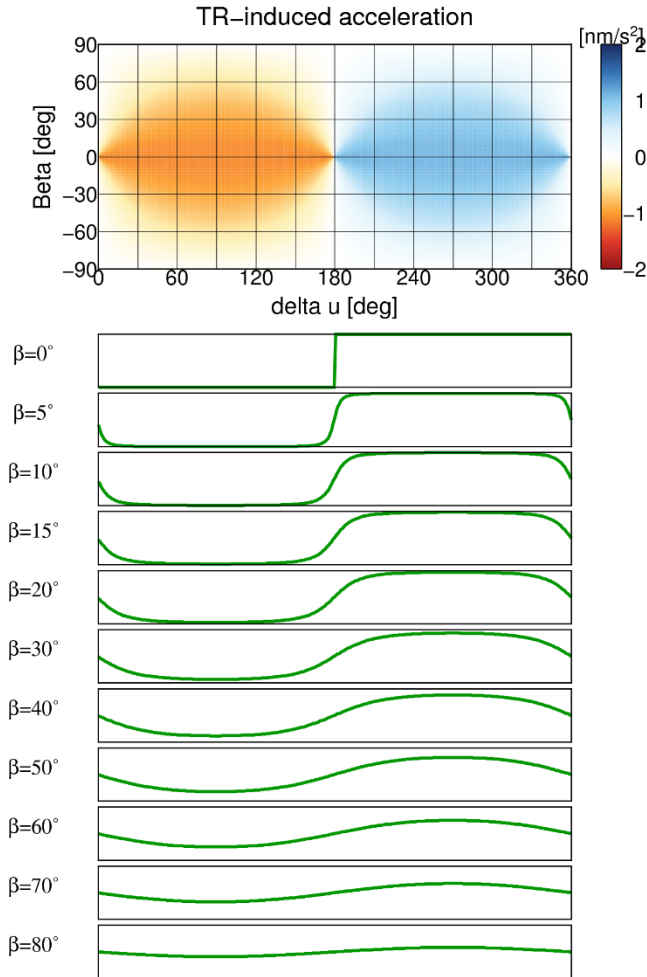
FOC



(\*) <https://www.gsc-europa.eu/support-to-developers/galileo-satellite-metadata>

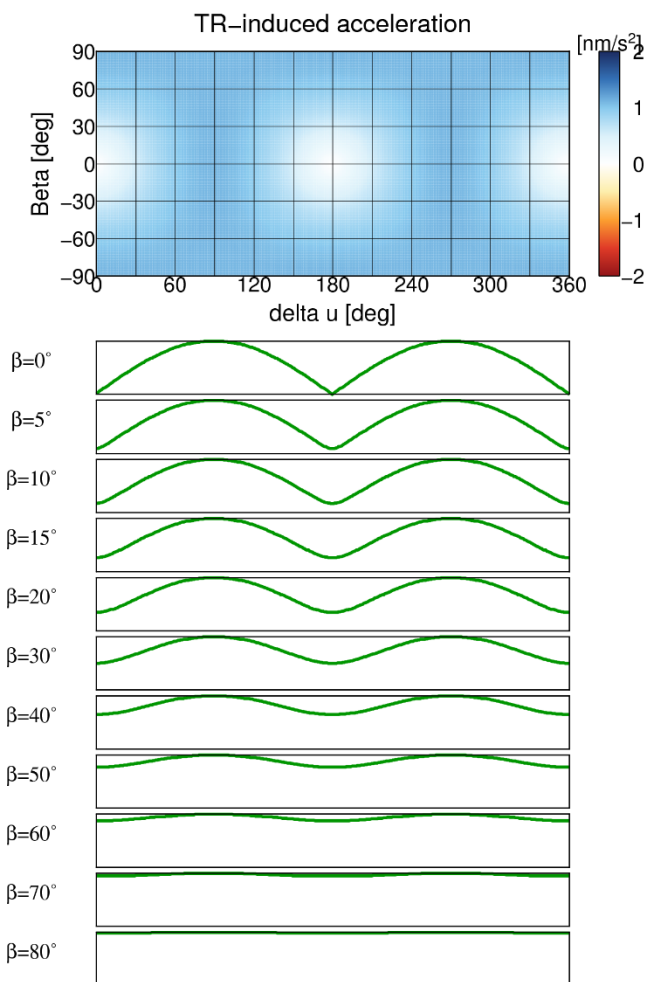
# Simulations of +X radiator effects

## Along-track component



# Additional terms in ECOM2 (D1S)

Projection on ECOM2  $D$  axis (satellite–Sun)



To be accounted by ECOM2:

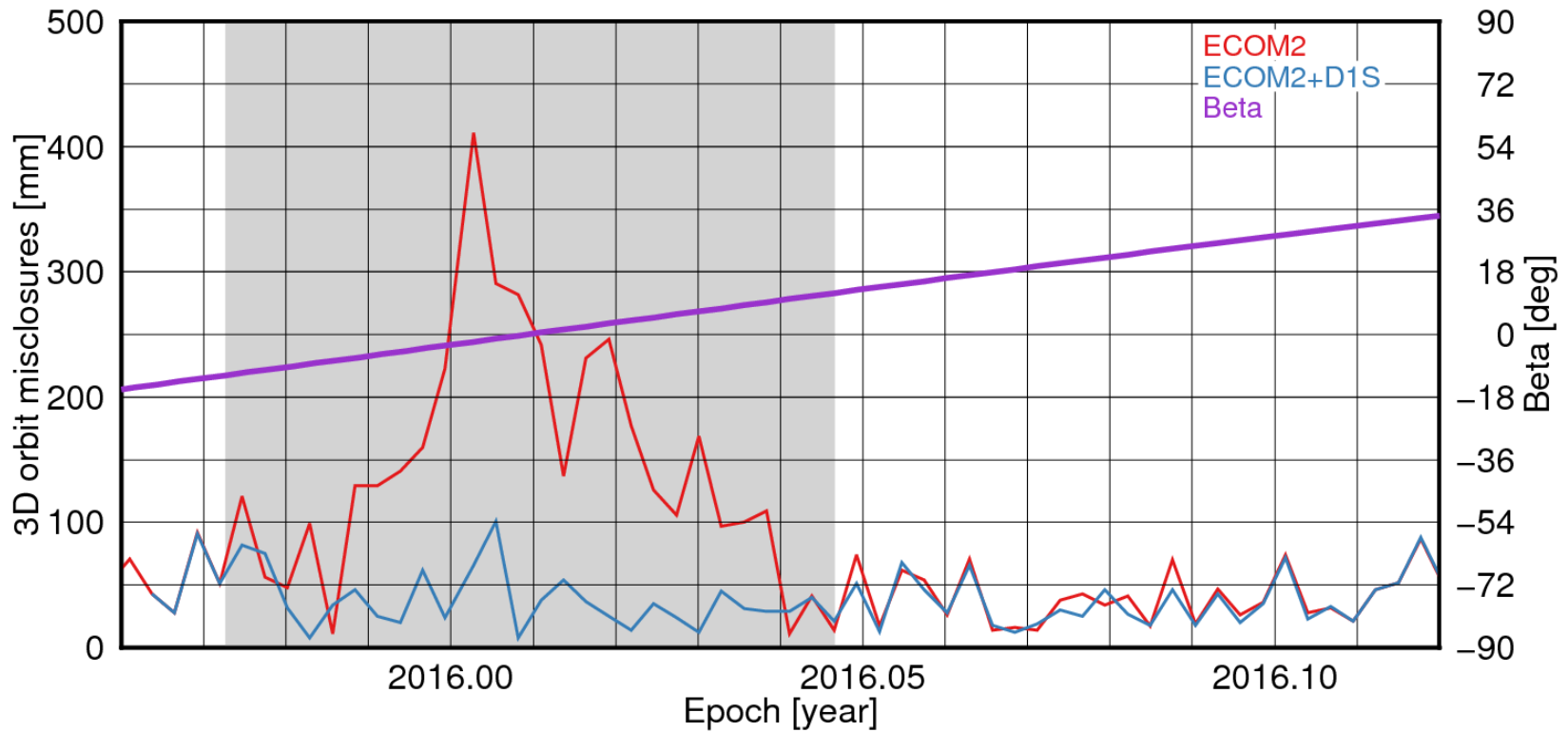
- for low  $\beta$  angles requires a once-per-rev sine term in  $D$ ,
- for high  $\beta$  angles a constant term in  $D$  is sufficient.

Actions taken:

- introduced D1S for  $|\beta| < 12^\circ$  for Galileo satellites,
- reprocessed the data from one eclipse season for Galileo.

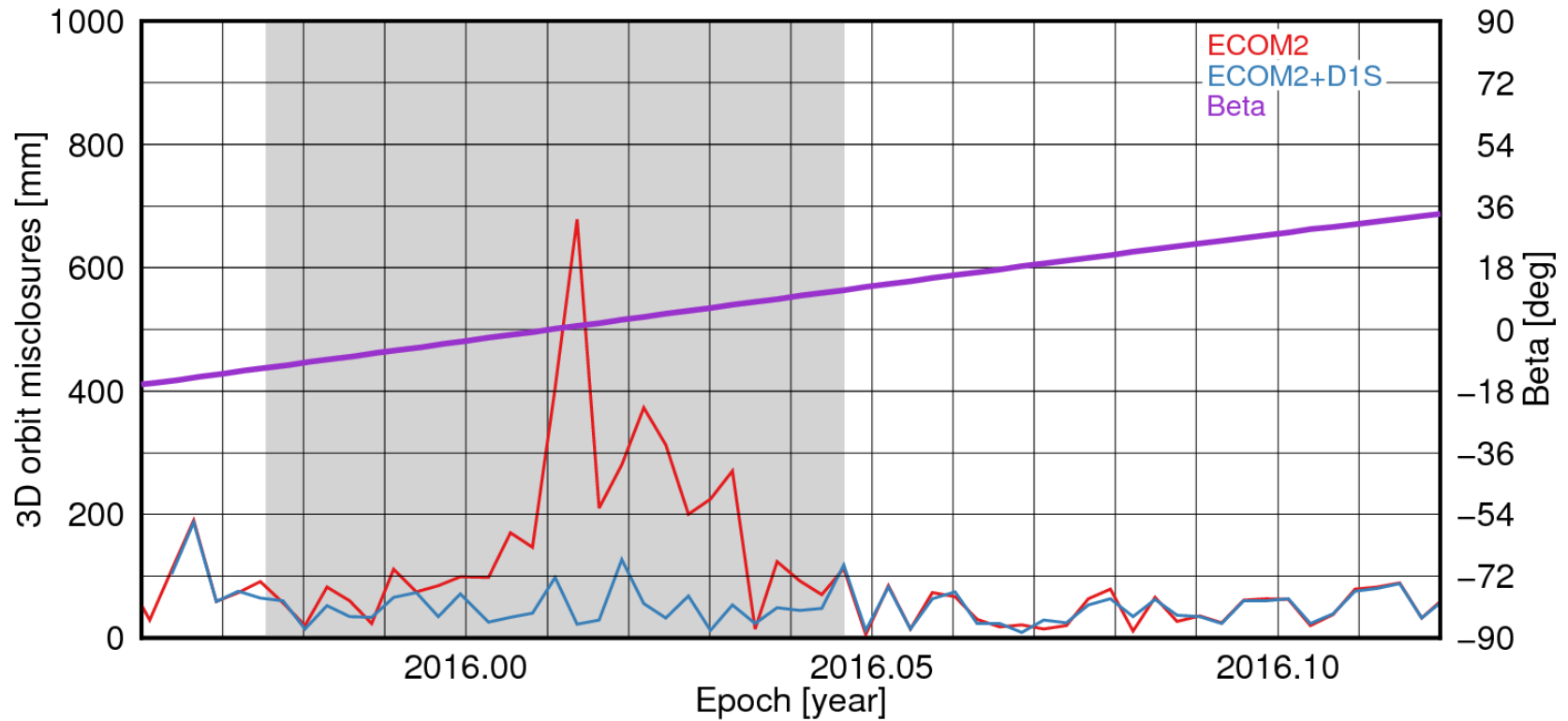
# Results: Orbit Misclosures

- Orbit misclosures for E11 during eclipse phase in Dec 2015 – Jan 2016:



# Results: Orbit Misclosures

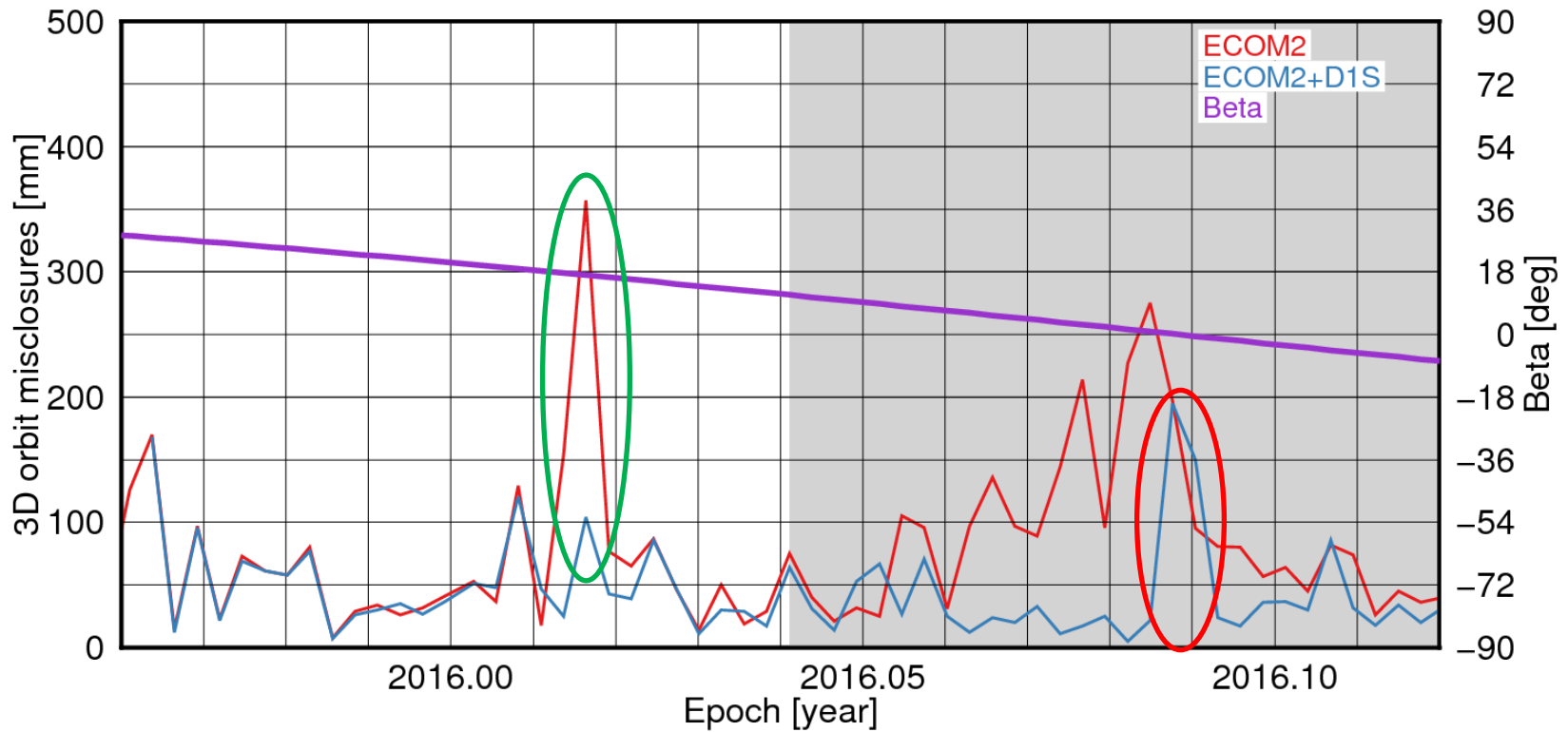
- Orbit misclosures for E26 during eclipse phase in Dec 2015 – Jan 2016:





# Results: Orbit Misclosures

- Orbit misclosures for E30 during eclipse phase in Jan - Feb 2016:



# Results: SLR Residuals

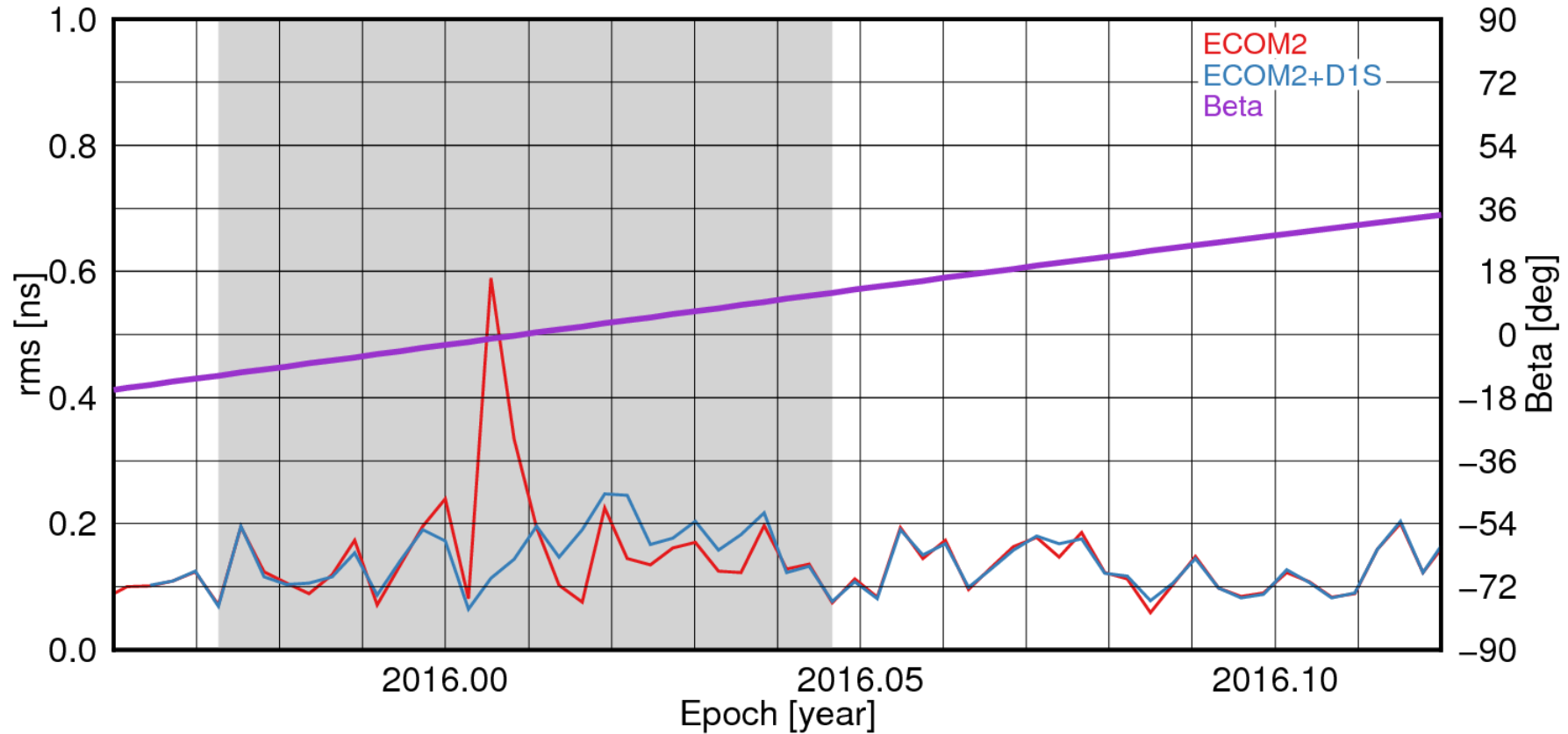
## Summary on the SLR residuals:

- the pattern is left unchanged (shrinking at orbit noon and expansion at orbit midnight);
- the scatter of the SLR residuals is reduced during eclipse phases in Dec 2015 – Feb 2016:

	ECOM2	ECOM2+D1S
IOV	$-12.7 \pm 57.3$	$-16.7 \pm 53.8$
FOC	$-9.7 \pm 49.0$	$-11.5 \pm 46.7$
IOV+FOC	$-10.6 \pm 52.3$	$-13.4 \pm 49.6$

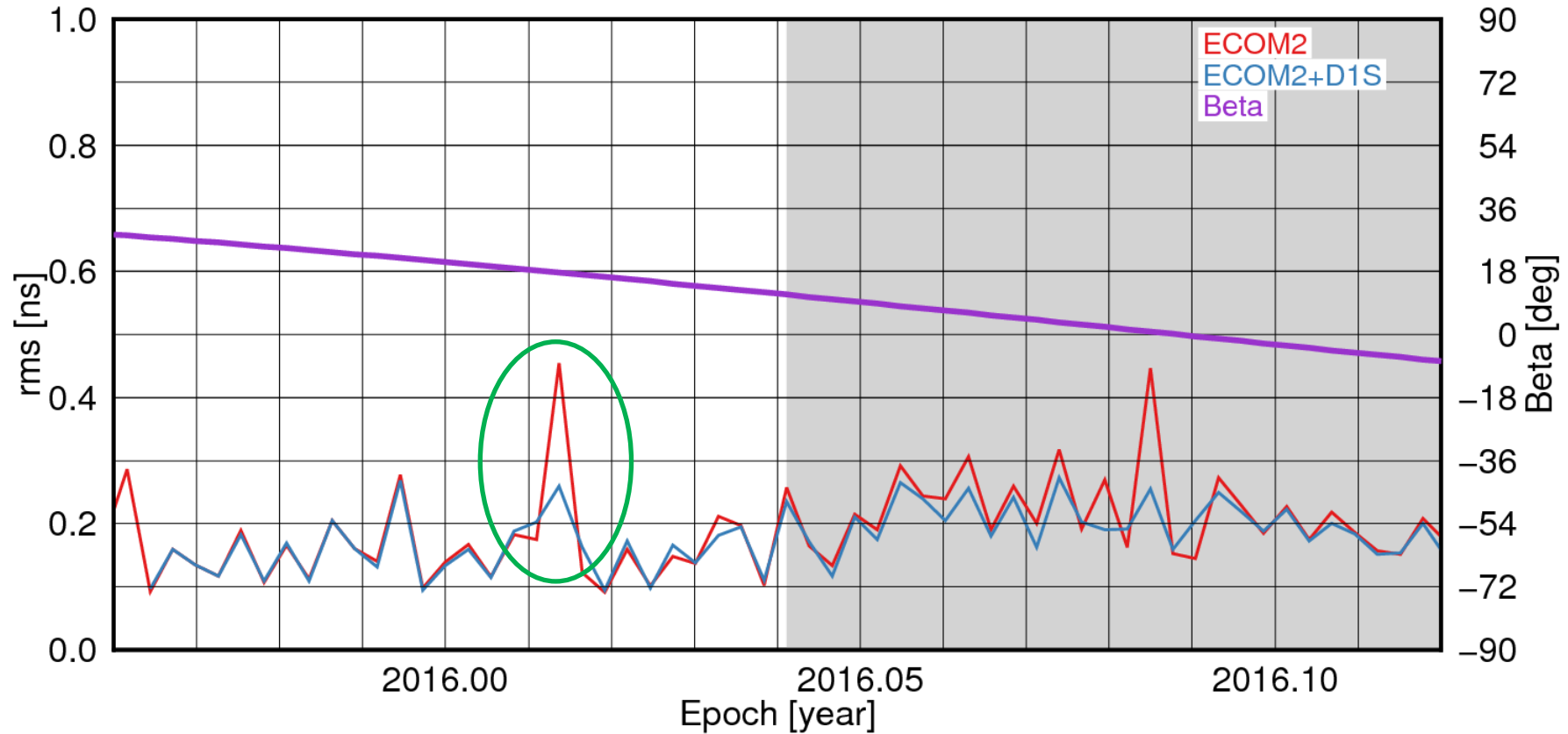
# Results: Satellite Clocks

- RMS of the linear clock fit for E11 in Dec 2015 – Feb 2016:



# Results: Satellite Clocks

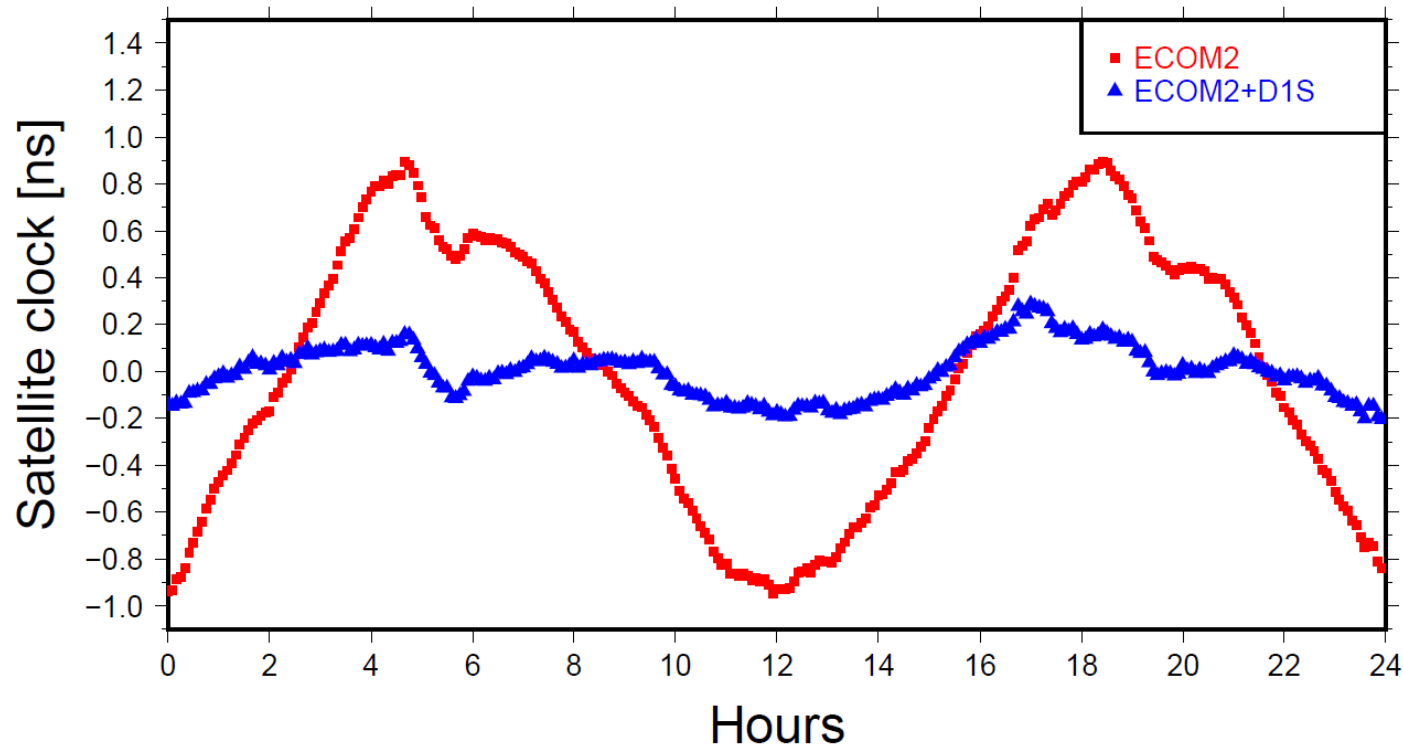
- RMS of the linear clock fit for E30 in Dec 2015 – Feb 2016:



# Results: Satellite Clocks

- Estimated satellite clocks (extreme case):

## E11 clock on 02 Jan 2016



# Conclusion

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- The recently published Galileo metadata shed light on how to model shadow crossings of the satellites, e.g.,
  - attitude control,
  - complete antenna correction models,
  - surface properties.
- Details on the internal temperature management of the satellites are appreciated.
- The unaccounted thermal effects may significantly deteriorate the estimated orbit.
- Addition of once-per-rev sine term in  $D$  to ECOM2 during eclipses significantly improves orbit modelling of Galileo satellites (should be added only for small  $\beta$  angles).