#### **ORIGINAL ARTICLE**



# Rapid response quality control service for the laser ranging tracking network

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Received: 11 May 2018 / Accepted: 31 August 2018 © Springer-Verlag GmbH Germany, part of Springer Nature 2018

#### **Abstract**

Quality control systems for satellite laser ranging (SLR) observations have been developed at a number of analysis institutes worldwide, using various software packages of precise orbit determination and data analysis. Satellite laser range observations, primarily from the two LAGEOS satellites but also from other satellites in low-Earth orbits and up to GNSS altitude, are being processed on a sub-daily to weekly basis. The generated quality control reports are widely used to detect various kinds of problems and quickly provide anomalous information to laser ranging stations. They have been effective in shortening the time to return to normal operations when anomalous data are detected and in quantifying the performance of laser ranging stations. Consequently, a rapid feedback loop has now been incorporated in the modern SLR operation.

Keywords Satellite laser ranging · Precise orbit determination · Quality control · Data anomaly detection

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Published online: 10 September 2018

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#### 1 Introduction

The majority of currently active satellite laser ranging (SLR) stations have achieved sub-centimeter precision in the two-way range measurements between a ground station and

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satellites at various altitudes. SLR has continuously contributed to global geodetic products, orbit validation and other areas, and, as a result, the high-quality range measurements in the optical regime have attracted many users.

The quality control process, the topic of this paper, plays a key role in the operation of the SLR network. No systems are faultless. This is especially valid for SLR, because SLR systems, although very similar in configuration, are not standardized in design and selection of components.

When an SLR station tracks a satellite and uploads the data, it cannot fully guarantee the quality. Stations have to calculate the root mean squares (RMS) of residuals generated during the normal-point formation process <sup>1</sup> (compressing all shot-by-shot data points to be represented by a single point per few seconds to a few minutes) so that they can reject obvious outliers. However, it is simply a statistical dispersion in its own observations, and generally speaking, it is not realistic for a station to locally assess the accuracy of their own observation data even at a meter or 10-meter level. Hence, SLR stations need some feedback about the data quality based on independent analysis. This paper discusses the activities in the frame of the International Laser Ranging Service (ILRS) (Pearlman et al. 2002) to generate and disseminate quality control information to the tracking stations.

In the early stage of SLR, the observation data were recorded on magnetic tapes and it took usually a few weeks or even months to get them analyzed. When a problem occurred at a station, therefore, the anomalous data continued flowing in for a long time. By the early 1990s, however, it became possible, through computer networking, for the analysts to obtain SLR data and provide feedback to the stations within a few days (Eanes et al. 1994; Ourensma and Noomen 1998). Currently, ILRS tracking stations are encouraged to release the SLR data within a few hours after the end of the pass. We can thus apply a quick-look routine analysis within a day or even less.

There are a number of ongoing quality control activities in SLR analysis institutes, such as reported by Glotov et al. (2004), Pavlis and Kuzmicz-Cieslak (2007), Otsubo et al. (2008) and Müller and Bloßfeld (2013). In this paper, we shall focus on common procedures and worldwide collaboration, without going into the analysis details of each institute.

## 2 Data flow

Continuous information flow in the SLR quality control procedure is outlined in Fig. 1. After the tracking data are transferred from each station to one of the two ILRS Operations Centers, data files are promptly organized and archived

https://ilrs.cddis.eosdis.nasa.gov/data\_and\_products/data/npt/index. html.



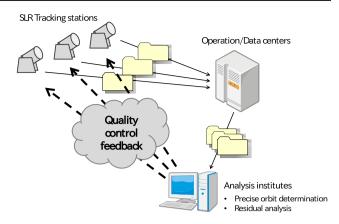


Fig. 1 SLR data flow in quality control sequence

at the ILRS Data Centers after applying a format check and some nominal verification of the reported values in various fields (Noll et al. 2018).

The quality control analysis providers listed in Table 1 have routinely contributed to the quality control of the SLR data as of 2018. They routinely fetch the SLR data files from the Data Centers and pass them through a data reduction and orbit determination process to see how the observations fit the reference orbits. As shown in Fig. 1, they send quality control feedback to the tracking stations. In particular, when anomalous observations from a certain station are detected, the station should be notified quickly.

Accuracy and reliability are the most important. Each institute has its own analysis software as listed in Table 1 and strives to improve its physical models and algorithms. Some of them have also benefited from long-term intercomparison tests through the Analysis Standing Committee of the ILRS (Luceri et al. 2018a). Every effort should be made to avoid activating a false alarm.

Rapidness is also important. Most quality control providers execute daily or weekly (Table 1) runs, and some update the analysis reports even more frequently.

The results and products from these quality control providers are described in the next section.

## **3 Quality Control**

### 3.1 Pass-by-pass analysis

Among modern space geodetic techniques, SLR is the only one that measures range directly. Every SLR station takes great care in system design, calibration, and data scrutiny to minimize measurement errors. Observational errors can be classified into two groups: random and systematic errors. Random errors will average out; systematic errors can alias the data products. The rapid quality control services for SLR

 Table 1
 Analysis institutes providing quality control information

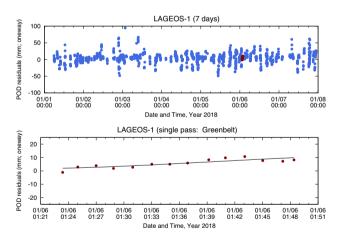
Institute	Software	Output	Satellites	Update	Duration
Astronomical Institute, University of Bern, Switzerland	Bernese 5.3	Range bias	GNSS	Daily	2000-present
Deutsches Geodätisches Forschungsinstitut, Germany	DOGS 5.4	Range & time bias	ETALON, LAGEOS and LEOs	4-hourly	2003-present
Hitotsubashi University, Japan	c5++ R889	Range & time bias	GNSS, ETALON, LAGEOS and LEOs	6-hourly	1998-present
Joint Center for Earth Systems Technology, USA	GEODYN II and SOLVE II	Range & time bias, Residual map	ETALON, LAGEOS and LEOs	Daily	2007-present
Information- Analytical Center, Russia	STARK-C 7.7	Range & time bias	LAGEOS	Daily	1997-present
NERC Space Geodesy Facility, UK	SATAN_SX	Residual plots	ETALON and LAGEOS	Daily	1997-present
Shanghai Astronomical Observatory, China	SHORD-II	Range & time bias	ETALON and LAGEOS	Weekly	1999-present
Wroclaw University of Environmental and Life Sciences, Poland	Bernese 5.2	Range bias & residual plots	GNSS	Daily	2016-present

are mainly focused on the detection of systematic errors that far exceeds the random error.

Precise orbit determination is the main tool for the quality control. Every institute involved in quality control analysis has its own orbit determination software which is automatically run on a prescribed schedule. The majority of the institutes use only SLR data to generate the reference (i.e., best-fit) orbits, but there are some institutes [Astronomical Institute, University of Bern (AIUB), and Wroclaw University of Environmental and Life Sciences (WUELS)] whose reference orbits of the GNSS (Global Navigation Satellite System) satellites are generated from microwave data.

For the purpose of quality control, the number of unknown parameters is limited. In the SLR-only analysis institutes, for instance, only orbital parameters are solved for while other parameters such as the coordinates of the stations are usually fixed to a-priori values. Therefore, the accuracy of the adopted station coordinates has a direct effect on the accuracy of quality control and and it is very important, especially for new, reinstalled, or significantly improved stations, and for stations in seismically active areas.

After iterative fits with outlier elimination, the post-fit residuals of normal points typically scatter from 0.7 to 1.5 cm weighted RMS for LAGEOS and LAGEOS-2. The



**Fig. 2** A LAGEOS post-fit residual example of a good pass. Top: the whole network throughout the 7-day arc. Bottom: one of many good behaving passes

residuals scatter more, from 1 to 4 cm weighted RMS, for low-orbit spherical satellites such as Starlette, Stella, Ajisai and LARES. The station-dependent weights are empirically based on the quality of the station data.

The top part of Fig. 2 shows typical post-fit range residuals from LAGEOS orbit, which is taken from the latest



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quality control analysis at Hitotsubashi University. For each 7-day arc, eleven parameters (six orbit (position and velocity) elements and five empirical parameters in along-track and cross-track components) are determined. In this particular span, the weighted RMS was 1.0 cm after iterative removal of outliers. The remaining noise stems from either imperfect physical models or observation errors. For instance, in Fig. 2 the normal-point residuals from a single pass observed by the Greenbelt SLR station are highlighted in red in both plots. The bottom plot is zoomed-in with a best-fit curve plotted in black, which is defined by a range bias parameter and a time-bias parameter, both of which are introduced later. Most SLR data from good performing stations align near the zero level, as in this example.

If we detect observations from a station deviating far from the zero level, i.e., the reference orbit, the station is likely to have a problem in its operation. We sometimes see the cases like the example shown in Fig. 3 where all the SLR data in a pass are deviating greatly and consistently from the reference orbit. We also see the cases as the one displayed in Fig. 4 where the residuals show a large negative-to-positive (or opposite) trend.

These two kinds of errors are called range bias  $\Delta_R$  and time bias  $\Delta_T$ , respectively. The majority of anomalous data are the result of either range bias or time bias. The conceptual diagrams for a satellite pass are shown in Fig. 5. Range bias is defined as a constant error in the range observation and expressed in a unit of length such as meter or millimeter. Time bias is defined as a constant error in the time tag and expressed in a unit of time such as microsecond or millisecond. Note that the satellites dealt within this paper orbit the Earth at the velocity of roughly 4–8 km/s. A one-microsecond time bias thus corresponds to 4–8 mm in the along-track component.

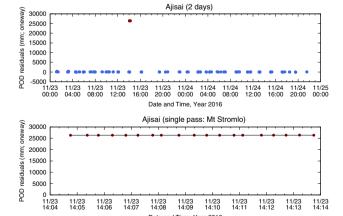
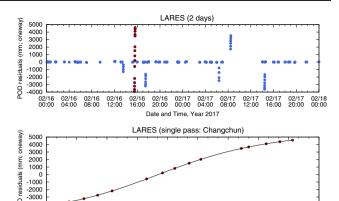


Fig. 3 An Ajisai post-fit residual example with large range bias. Top: the whole network throughout the 2-day arc. Bottom: a Mt Stromlo pass with  $+26 \,\mathrm{m}$  range bias



**Fig. 4** A LARES post-fit residual example with large time bias. Top: the whole network throughout the 2-day arc. Bottom: a Changchun pass with +1 ms time bias

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Given a sufficient tracking duration and a sufficient number of normal-point observations, these two bias parameters, range bias  $\Delta_R$  and time bias  $\Delta_T$ , can be estimated from a population of post-fit residuals whose *i*th element is defined as  $y_i$ . The observation equation is written as:

$$y_i = \Delta_R - \dot{\rho}_i \Delta_T + \epsilon_i \tag{1}$$

where  $\dot{\rho}_i$  is the range rate for the *i*th observation, and  $\epsilon_i$  is the error in the *i*th observation. This equation should be stacked for the number of observations per pass, and the two bias parameters  $\Delta_R$  and  $\Delta_T$  can be estimated in a standard least square adjustment.

In reality, it is sometimes difficult to estimate both of the two bias parameters when the tracking duration is too short or the number of observations is too few. In such cases, most analysis institutes provide only range bias, i. e. the simple mean of the residuals.

Figure 6 is an example of quality control reports of Hitotsubashi University. The providers in Table 1 whose output is 'Range & time bias' also generate and update numerical tables in similar formats. Each line corresponds to one pass. The column of 'rb' corresponds to estimated range bias, and that of 'tb' to estimated time bias, both of which are followed by their formal errors of Eq. 1 in parentheses. In this case of the Yarragadee station, the error in the orbit models, about 1 cm RMS, well exceeds the precision in the normal-point observations, about a few mm RMS, and therefore, the bias values given here scatter more than the actual observation precision. This is the typical case for good observations from good stations. If these bias values showed a huge jump which is significantly larger than its estimated error, and if the bias persisted for multiple passes, it is likely that the station has a problem with its equipment, software or operation.



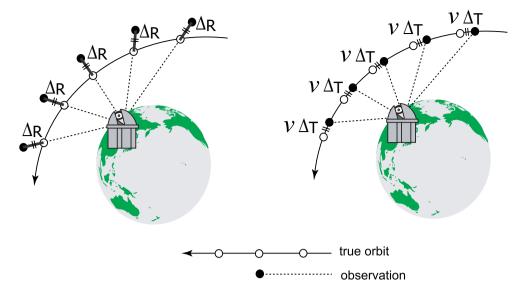


Fig. 5 Range bias (left) and time bias (right). Illustrated are positive biases in both cases. v is the velocity of the satellite

#							
# 7090 = YARRAGADEE							
#sat site date time dur	rb mm error	tb us error	prec bad total	rms	pres temp hum	sdelay shft	rms cfg r wlen
AJI1 7090 2018/01/07 18:54 7	14 (6)	-4.9 ( 2.9 )	3 0 / 17	9	979.5 295.8 31	13079 0	2 6 1 0 532
LARS 7090 2018/01/07 19:02 13	4 (2)	-0.5 ( 0.5 )	1 0 / 29	4	979.5 295.8 31	13079 0	2 6 1 0 532
GA03 7090 2018/01/07 20:48 132	-3 (6)	42.7 ( 24.3 )	3 0 / 8	8	980.6 295.4 33	21896 -0	4 6 1 0 532
LAG1 7090 2018/01/07 21:01 21	6 (3)	2.3 ( 4.3 )	3 0 / 10	5	980.2 294.9 32	13082 0	2610532
SARL 7090 2018/01/07 21:24 0	-11 ( 10 )	( )	2 0 / 4	3	980.3 294.6 33	13082 0	2610532
ETA1 7090 2018/01/07 21:48 57	1 ( 3)		6 0 / 11	14	980.8 295.2 35	21896 -0	4 6 1 0 532
SARL 7090 2018/01/07 23:02 5	8 ( 4)	-3.3 ( 0.8 )	1 0 / 13	. 4	981.0 296.4 33	13083 1	2 6 1 0 532
AJI1 7090 2018/01/07 23:09 2	20 ( 8 )		3 0 / 7	- 11	981.0 296.4 33	13083 1	2 6 1 0 532
LAG2 7090 2018/01/07 23:15 19	17 ( 4)	1.4 ( 6.5)	2 0 / 11	4	981.0 296.9 33	13083 1	2 6 1 0 532
STRL 7090 2018/01/08 00:13 11	0 ( 3 )	-3.2 ( 0. <u>7</u> )	1 0 / 15	4	981.0 298.4 28	13084 0	2 6 1 0 532
LAG1 7090 2018/01/08 00:27 33	11 ( 2 )	0.5 ( 1.5 )	2 0 / 18	. 5	980.8 299.6 28	13084 0	2 6 1 0 532
AJI1 7090 2018/01/08 01:03 17	24 ( 5 )	0.2 ( 1.0 )	3 0 / 20	ΙĬ	980.7 300.7 25	13084 0	2 6 1 0 532
CRY2 7090 2018/01/08 01:52 7	6 ( 4 )	2.2 ( 0.7 )	1 0 / 13	3	980.5 302.6 23	13084 0	2 6 1 0 532
STRL 7090 2018/01/08 02:04 7	9 5 5	6.9 ( 1.4 )	2 0 / 11	5	980.4 303.1 22	13084 0	2 6 1 0 532
AJI1 7090 2018/01/08 03:12 1	18 ( 11 )		2 0 / 4	9	979.9 305.0 20	13084 1	2 6 1 0 532
LAG2 7090 2018/01/08 03:17 32	1 4 4	5.7 ( 2.8 )	3 0 / 12	.5	979.7 305.7 19	13084 1	2 6 1 0 532
GA02 7090 2018/01/08 04:58 227	4 ( 3 )	-126.3 ( 54.0 )	6 0 / 21	14	977.8 309.6 16	21897 1	6 6 1 0 532
GL33 7090 2018/01/08 06:42 103	74 ( 17 )		/ U / 4	17	976.8 310.5 15	21898 0	6 6 1 0 532
STEL 7090 2018/01/08 07:05 7	10 \ 10 \	0.8 ( 0.9 )	2 0 / 17	6	977.1 310.8 16	13086 2	2 6 1 0 532
LAG1 7090 2018/01/08 07:19 9	10 ( 10 )	9.8 ( 17.0 )	4 0 / 6	4	976.9 310.7 16 976.8 311.1 16	13086 2	2 6 1 0 532
STRL 7090 2018/01/08 07:38 4 LAG2 7090 2018/01/08 07:47 15	1 4 4	4.0 ( 1.9 ) 5.7 ( 8.5 )	1 0 / 11	4	976.8 311.1 16 976.6 311.2 15	13086 2 13086 2	2 6 1 0 532 2 6 1 0 532
LARS 7090 2018/01/08 07:47 13	-6 ( 8)	1.4 ( 2.3 )	1 0 / 6	3	976.6 311.1 15	13086 2	2 6 1 0 532
LARO 1000 Z010/01/08 07:00 3	-0 ( 8 )	1.4 ( 2.5 )	1 07 7	3	310.0 JII.I ID	13000 2	Z 0 I U 03Z

Fig. 6 Example of a quality control report (excerpted from the 16 January 2018 report issued by Hitotsubashi University). Listed are: satellite name, station ID, observation date and time, pass duration (in min), range bias and its estimation error (in mm), time bias and its estimation error (in  $\mu$ s), estimated precision of normal point (in mm), rejected and

total data points, precision of raw ranging (in mm), pressure (in hPa), temperature (in Kelvin), humidity (in percents), applied system delay (in mm), delay shift (in mm), precision of calibration ranging (in mm), two ILRS configuration indicators, release flag and laser wavelength (in  $\mu m$ ), from left to right

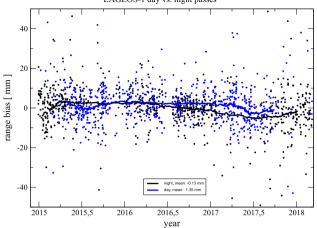
In addition to numerical tables, some institutes provide graphical information. Various kinds of plots are being generated at Deutsches Geodätisches Forschungsinsitut, Technische Universität München (DGFI-TUM). One example in Fig. 7 shows LAGEOS range bias estimates of the Herstmonceux data. Users can visually see long-term behavior of their day-time observations and the night-time observations. In this particular case, due to the high latitude (N 51°) of the station, the amount of day–night passes has a clear annual pattern, and the day–night differences of the moving average curve nicely stay within 3 mm. Graphical plots are clearly more comprehensible and more informative for users.

Historically, only the two LAGEOS satellites, i.e., LAGEOS and LAGEOS-2, had been applied to quality control since they are primarily used for geodetic analyses, their characteristics are well understood and their orbits are less prone to poorly known physical process (e.g., atmosphere and earth gravity field). Quality control providers have gradually added high-orbit and low-orbit satellites to the analysis of quality control reports. Despite the fact that the orbit determination precision of these satellites is not as good as that of the LAGEOS satellites, the use of a wide variety of satellites has great advantages as we explain in the following. Firstly, the timing of problem occurrence can be more reli-



Herstmonceux (7840) bias analysis: one per pass and running averages of 100 passes

LAGEOS-1 day vs. night passes



**Fig. 7** A plot example of LAGEOS range bias of the Herstmonceux station, UK. DGFI-TUM routinely generates not only numerical tables but also these plots for all stations

ably identified because of the increased number of passes. Secondly, we can detect problems that are not possible to be detected by analyzing only LAGEOS data because some stations have multiple hardware configurations for different satellite altitudes and because some stations are more productive in lower or higher satellites than LAGEOS. Thirdly, a range-dependent bias, often proportional to the range resulting from a problem in the clock frequency, can be clearly identified.

It should be noted that sufficient coverage during a pass is important for data quality assessment. In particular, a horizon-to-horizon pass duration of a high-orbit target, such as the two ETALON satellites and the GNSS satellites, is typically 5–8 h. We need SLR coverage for at least an hour, preferably spread over the pass, to separate the two bias parameters. However, when the SLR coverage during a pass is short, we often find that the time bias parameter cannot be determined at all, or are not well determined, for high-orbit satellite passes.

It is no surprise that AIUB provides only the range bias estimates in their routine reports dedicated to the high-orbit GNSS satellites. Reference orbits in the AIUB quality control analysis are based only on GNSS microwave data. In 2017, the number of registered passes on GNSS was almost four times larger than the number of LAGEOS passes, which allows for a rigorous analysis of SLR biases at high-altitude targets. The AIUB orbits come from the analysis at the International GNSS Service processing at the Center for Orbit Determination in Europe (Dach et al. 2009). Consequently, the reported SLR bias values include the systematic offset between the two independent techniques; Thaller et al. (2011) studied the impact of mismodelled antenna phase center issues and unconsidered range bias, and Sośnica et al.

(2015) investigated the pattens of SLR residuals in relation to SLR system types and onboard retroreflector array types.

As a result of there being many more SLR satellites in space than a decade or more ago, a ground station can now see 10 SLR satellites or even more above the horizon, while it can track only one satellite at a time. The stations are encouraged to switch targets frequently (often called "interleaving mode") rather than solid tracking from the beginning to the end of the pass, and, by doing so, they would obtain SLR data from more satellites in an efficient way.

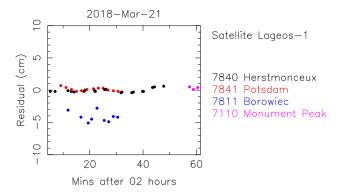
#### 3.2 Residual plots

Quality control information is reduced to just one line per pass in the previous subsection (see Fig. 6). We have shown several normal-point residual graphs (Figs. 2, 3 and 4), but they are not routine products. Residual graphs are indeed more informative and useful for a precise understanding of the behavior of observation data. NERC Space Geodesy Facility (SGF) and WUELS provide online residual graphs so that the users can visually see the precise details of normal-point residuals such as outliers and trends.

The SGF updates on its website each day interactive normal-point residual plots of LAGEOS and Etalon normal points submitted globally over the previous 7 days as a by-product of its routine analysis center activity. The SGF also separately carries out daily "short-arc" quality control, which extracts and plots a common-view pass from multiple stations in close proximity (Sinclair and Appleby 1993). Through these plots available online, we see how each normal point behaves with respect to the reference orbit. An example is shown in Fig. 8, which shows four stations tracking LAGEOS. In this particular case, the LAGEOS satellite orbits firstly over Europe and then over the USA. This plot would suggest a range bias is present in the Borowiec range measurements. These plots show solid/sparse tracking patterns and the behavior of each normal point. For instance, in the past, we frequently saw outliers at the beginning or at the end of a pass due to insufficient outlier rejection, but it is rare today thanks to alerts from this kind of service.

WUELS provides online tools for an interactive generation of SLR residual plots of GNSS satellites, based on the microwave orbits processed at AIUB in the framework of the IGS multi-GNSS Experiment (Prange et al. 2017). More than 50 active GNSS satellites are today tracked by the SLR stations, which results in a substantial number of SLR observations. The WUELS service allows for the residual analysis with respect to the date and time of data acquisition, differences from night-time and day-time tracking, dependencies of residuals as a function of the elevation and azimuth angles. All of them help in identification of various systematics. The analyses can be performed for individual stations and satellites or for a user-defined group of satellites or stations.





**Fig. 8** A residual plot example of a LAGEOS-1 common-view pass from Europe-USA region. This is taken from NERC SGF's daily shortarc quality control

# 4 Integration and communications

### 4.1 Quality control feedback

All of the quality control activities described in the previous section are intended to give prompt feedback to the tracking stations. Some stations are routinely watching the quality control reports, and the analysis institutes also send an alert message when they detect anomalous observations.

All of the analysis institutes in Table 1 routinely update the quality control reports on their websites and/or make them available through the ILRS archive and the ILRS mailing list. The long-time records of daily and weekly reports are archived not just at each institute but also aggregated on the ILRS Data Center, at Crustal Dynamics Data Information System (CDDIS) of National Aeronautics and Space Administration (NASA). The URLs of the SLR quality control services are collectively listed in Table 2.

Users can see the quality control results from multiple institutes at once. Range bias estimates of the two LAGEOS satellites from multiple institutes are easily compared by the

combined range bias reports weekly sent by AIUB. Each line of the combined report contains range bias estimates from five providers. SLR station operators can verify the reliability of their data by checking whether an anomalous result is detected undoubtedly by all institutes or it is detected just by one institute and disagree with others. It also helps analysts to see whether their results are in harmony with others. When a series of anomalous observations are clearly detected from several institutes, it very likely stems from the observations, not the analysis. A warning message is then sent to the station. The whole quality control procedure in the analysis and the data transfer is almost automated, but manual handling is still to some extent involved in detecting anomalous data and also notifying the stations.

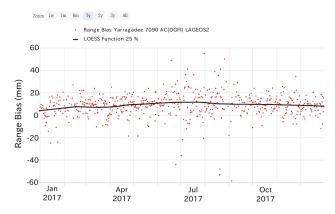
We usually use email for communication with stations. In particular, DGFI-TUM, Hitotsubashi University and JCET routinely alert to the SLR stations. In June 2011, the analysis institutes agreed to use the newly established "RapidService-Mail" mailing list implemented at DGFI-TUM to notify the stations and archive the anomaly detection events among analysis institutes.

It has been common to look at the text-based reports, but an interactive graphical web tool is more helpful. The time series of pass-by-pass range bias, time bias and other quality control parameters can be plotted in a web browser using the QC-viewer, a service provided by the Joint Center for Earth Systems Technology (JCET) as shown in Fig. 9. This service helps to give a comprehensive view of the long-and short-term behaviors of the quality of a certain station, optionally with a smoothing curve. The user has various options to choose from, such as the satellite, the station, the quantity of interest, the axis scales and so on, for customized graphs. In addition to JCET's own analysis results users can also choose those from the other institutes appearing on Table 1.

Table 2 URLs for the quality control services

Institute/service	URL			
Individual				
DGFI-TUM	https://ilrs.dgfi.tum.de/quality/weekly_biases/stations/			
Hitotsubashi University	http://geo.science.hit-u.ac.jp/slr/bias/			
NERC SGF	http://sgf.rgo.ac.uk/			
WUELS	http://www.govus.pl/			
Integrated				
ILRS Data Center, CDDIS, NASA	ftp://cddis.gsfc.nasa.gov/pub/reports/			
ILRS Global Performance Cards	https://ilrs.cddis.eosdis.nasa.gov/network/system_performance/global_report_cards/monthly/			
ILRS RapidServiceMail, DGFI-TUM	http://rapidservicemail.dgfi.tum.de/			
JCET QC-viewer	http://geodesy.jcet.umbc.edu/QC/			
JCET Visualization of ILRS Report Cards	http://geodesy.jcet.umbc.edu/ILRS_REPORT_CARD/			

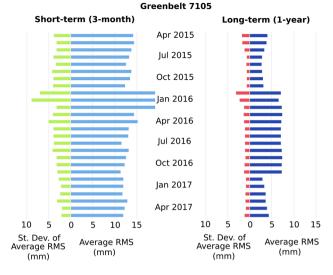




**Fig. 9** A range bias plot example of LAGEOS-2 range bias estimates of Yarragadee in 2017, plotted by JCET's online interactive tool. DGFITUM's quality control results are chosen in this case. The graph layout is rearranged for this paper

It is now possible to detect and notify a problem within a day. If a problem can be subsequently corrected afterward, the station is strongly advised to resubmit the corrected observations with an incremented release flag. We have seen that most of the problems can be solved instantly or within a few days. It minimizes the time that large bias errors are present in the data.

Pass-by-pass analysis reports from the quality control providers are gathered at the ILRS Central Bureau where a long-term statistics report is published from both qualitative and quantitative aspects. The "SLR Global Performance Report Card" covers the latest 12 months and is currently updated every month (every three months until 2012). Each report consists of two tables: one is for quantity and the other is for quality. The first table provides the ranking by the number of passes, and the number of normal-point observations, for different categories of satellites. The second table provides a summary of the five quality control reports providing a pass-by-pass bias analysis of the LAGEOS satellites. The results are averaged over time to produce a stability measure in two ways: a pass-by-pass range bias variation over the last three months ('short-term' stability) and a monthlyaverage range bias variation using over the previous 12 months ('long-term' stability). These tables are helpful to compare a station's performance with respect to others and to encourage the improvement of the quality and quantity of their data. They are also useful to monitor the evolution of a station's performance over time. To facilitate these comparisons, JCET has established a service that archives all report cards' data online and allows the user to create plots of these statistics over time for each station of interest (an example in Fig. 10). The results plotted are the average of the available short-term or long-term stability estimates from all quality control centers and the standard deviation of that estimate.



**Fig. 10** Rendering of the short-term and long-term stability measure from the quarterly report card of Greenbelt using the online interactive tool developed by JCET. The graph layout is rearranged for this paper

#### 4.2 Causes of anomalous measurements

We often receive a reply from the station about the cause of the problem. Based on our long experience in interacting with the stations, we have learned that certain patterns are connected to specific causes.

The most frequent error is constant offset range bias. Some SLR laser systems transmit multiple pulses at a fixed interval within one or a few nanoseconds, where one pulse is usually predominantly stronger than others. However, especially in the case that there is more than one strong pulse, the choice of the main pulse could be wrong. The other major cause originates from calibration. It is often caused by reflection from a point different from the calibration target or a system adjustment inadvertently made between calibration and ranging. If the internal system delay is wrongly measured, the SLR data are shifted by the same amount as its error. In both cases, a jump of the system delay is also observed. It is thus important to always keep an eye on the time series of the system delay measurement. Human input error during operation or data reduction can also be the cause.

Time bias is also detected almost as often as range bias. Note that the SLR observables are basically the combination of a time tag and a two-way range. Large time bias indicates a problem in time tagging. Time bias sometimes appears constant, and sometimes varying, i.e., the pass-by-pass time-bias estimates gradually change over time. The problem often comes from a frequency source, other timing devices and sometimes software. In the past, one second time bias was often seen after a leap second, but it is rare nowadays. Extremely large time bias of one day (wrong day) and even one year (wrong year) can be also detected.



Meteorological sensors such as a barometer and a thermometer are sometimes found to be an error source. Atmospheric delay models are constructed as a function of meteorological data and therefore a problem in these instruments manifests itself as a range error.

Analysis institutes must pay attention to avoid false alarms. Orbit determination has not always properly converged due to insufficient data, too many outliers, imperfect physical models, and orbit maneuvers. For instance, if a satellite is sparsely tracked, and one productive stations has a large time bias, then well performing nearby stations may have a spurious time-bias estimate. This is because time bias is inseparable from the along-track component of satellite orbits, and because large time bias can affect the orbital parameters if the actual anomalous data cannot be properly eliminated in the orbit determination stage. In another instance, the use of accurate station positions and velocities is essential for reliable and precise quality control results. Analvsis institutes thus obtain, or update, the station coordinates if a station is new, has undergone significant changes, or is displaced due to earthquakes. Currently, the latest version of SLRF2014 (Luceri et al. 2018b), an extension of ITRF2014, is widely adopted for quality control analyses and frequently updated.

# 5 Conclusions and ongoing/future studies

Precise orbit determination techniques and high-speed communications have been widely integrated to realize rapid quality control feedback for the SLR station network. Analysis institutes have routinely provided analysis reports in various ways and they have contributed to improving the overall quality and stability of the SLR observations.

A new ILRS Quality Control Board was organized in December 2015 to define and implement procedures to control data quality. The rapid quality control activities covered in this paper are one of the core tasks in this group, and the latest news and future plans are being discussed in monthly international teleconferences.

High-quality SLR data is beneficial not only for geodesy but also for satellite altimetry, satellite navigation and any mission that relies on SLR for its precise orbits.

This paper focused on the rapid response feedback for relatively large systematic errors, but it is also possible to look into systematic observation errors/trends at the millimeter level by accumulating longer span of SLR data (Otsubo 2014; Appleby et al. 2016). Both quality control methods, i.e., rapid and precise, should be utilized in a complementary fashion in the future so that systematic errors of any magnitude are eliminated from the data.

A new one-way ranging technology, Time Transfer by Laser Link (T2L2) demonstrated on the satellite Jason-2,

makes it possible to examine the accuracy of time tagging at SLR stations (Exertier et al. 2017) with a 1 ns accuracy. This is much more stringent than the requirement for ordinary SLR, roughly 100 ns for 1 mm ranging accuracy, and several anomalous cases have been detected. The T2L2 mission on Jason-2 was terminated in April 2018, and the future missions are strongly desired from the viewpoint of quality control.

Acknowledgements The quality control analyses are indebted not just to the authors of this paper but to a number of researchers at the analysis institutes, such as Ms. M. Kobayashi of Hitotsubashi University, Mr. C. Schwatke of DGFI-TUM, Dr. M. Kuzmicz-Cieslak of JCET, Dr. G. M. Appleby of NERC SGF, Mr. R. Zajdel of WUELS, Dr. A. Maier and Dr. T. Schildknecht of AIUB and many others. T. Otsubo acknowledges the support of JSPS KAKENHI, Grant Numbers 26400449 and 18K11632. E. C. Pavlis acknowledges the support of NASA Grant NNX15AT34A. The quality control services are made up of continuous and intensive collaboration with the ILRS tracking stations and also with the ILRS Operation Centers and Data Centers. The quality control analyses in early days were led by Dr. R. Eanes of Center for Space Research, University of Texas, and Dr. R. Noomen of Delft University of Technology. Finally, we would like to point out that the international collaboration in this field could not be achieved at this level without Prof. W. Gurtner, AIUB, who sadly passed away in 2009.

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