

**Strangeness production in the new version of the Liège intranuclear cascade model**

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The capabilities of the new version of the Liège intranuclear cascade model (labeled INCL ++6 from now on) are presented in detail. This new version of INCL is able to handle strange particles, such as kaons and the  $\Lambda$  and  $\Sigma$  hyperons, and the associated reactions and also allows extending nucleon-nucleon collisions up to about 15–20 GeV incident energy. Compared to the previous version, new observables can be studied, e.g., kaon, hyperon, and hypernuclei production cross sections (with the use of a suitable de-excitation code) as well as aspects of kaon-induced spallation reactions. The main purpose of this paper is to present the specific ingredients of the new INCL version and its new features, notably the newly implemented variance reduction scheme. We also compare, for some illustrative strangeness production cases, theoretical results calculated using this version of INCL with experimental data.

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## I. INTRODUCTION

Spallation reactions have applications in many different domains, such as medicine, astrophysics, nuclear physics, etc. This large range of applications explains why it was, and still is, so important to fully understand and model this type of spallation reactions as stand-alone models or embedded in transport codes. Spallation models are commonly used in the energy range 100 MeV–3 GeV, the higher energies being treated with string models. However, spallation models can in principle be used in a wider energy range. Notably, they can be used at energies up to 15–20 GeV if the new open channels are correctly treated.

The intranuclear cascade model of Liège (INCL), which was recognized by the IAEA benchmark in 2010 [1] as one of the best spallation models when combined with the ABLA07 [2] de-excitation code, has been considerably improved during the past decade. One can mention the light-ion-induced reactions [3], the improvements for few nucleons removal [4,5], and the extension to high energy [6,7]. The last point was motivated by various reasons. A notable example is the need to make a better transition between string models like the FTF model [8], for which the few GeV scale represents the low-energy part, and intranuclear cascade (INC) models like INCL where the few GeV scale represents the high-energy part. This energy range is crucial notably for cosmic rays interactions. We can also mention an interest in exotic physics, like processes involving strangeness that appear in this new region with new particles and hypernuclei formation.

The high-energy improvement of INCL involved three successive stages: multipion production in the final state of binary nucleon-nucleon and pion-nucleon collisions [6], production

of  $\eta$  and  $\omega$  particles [7], and production of strange particles described for the first time in this paper. This last step is based on a previous work, which reunified the main ingredients needed for the implementation of strange particles [9]. Here, we present the new version of INCL and the calculation results related to strange particle production. The observables, which will be discussed, are generated only or almost only by the intranuclear cascade. Observables depending on both the intranuclear cascade and the de-excitation stages will be discussed in a forthcoming publication.

The paper is organized as follows. In Sec. II, we present the main changes in INCL since the previous public version, and in Sec. III we explain in detail the variance reduction methods introduced in INCL to facilitate the study of strangeness production physics. Next, in Sec. IV, we compare the results obtained with the new version of INCL to experimental data of strange particle production for different initial conditions. Finally, Sec. V is devoted to the conclusions and perspectives.

## II. THE PRESENT INCL VERSION

Our starting point is the INCL ++ version, described in Ref. [3]. It has essentially the same physics content, concerning nucleon-nucleus and pion-nucleus interactions, as the Fortran version INCL4.6 (see Ref. [10] and references therein). Only minor differences exist. However, one relevant difference between the INCL ++ and the Fortran version is the capability to treat reactions induced by light ions ( $A \leq 18$ ). Since the publication of Ref. [3], two main aspects were improved: the few nucleon removal channels, which are discussed in detail in Refs. [4,5], and the high-energy part [6,7,9].

The extension of INCL to high energies allows simulations up to  $\approx 15$  GeV instead of the former upper limit of 2–3 GeV of the previous version. This high-energy extension was performed in three steps: multipion emission in binary collisions, which is the main contribution at high energies [6]; implementation of  $\eta$  and  $\omega$  particles [7]; and finally, implementation of kaons and hyperons, described for the first time in this paper. In this section, we describe the different input data and implementations linked to processes involving strangeness. Cross sections are crucial ingredients. They are briefly discussed below, but the reader will find all details in Ref. [9].

The implementation of strangeness was divided into four main parts: first, the new particles to implement were selected. Then, their characteristics (mass, charge, isospin, decay modes, and average nuclear potential) were provided. Third, the interactions of these particles were described essentially with the use of experimental measurements and isospin symmetry. This includes the determination of production, scattering, and absorption elementary cross sections as well as the phase space distribution of outgoing particles in binary collisions involving strange particles in the initial state and/or in the final state. The last part was the post-cascade treatment of the new particles.

### A. Particles

The very first step before strangeness could be implemented was to determine which particles are to be added in INCL. This was based on two criteria. The first obvious criterion is the production rate in  $NN$  and  $\pi N$  collisions, which are the most common reaction types involved in intranuclear cascades. However, we also considered the quantity of available information in addition to *a priori* knowledge. For those reasons, it was decided to consider the kaons ( $K^0$  and  $K^+$ ), the antikaons ( $\bar{K}^0$  and  $K^-$ ), the  $\Sigma$ 's ( $\Sigma^-$ ,  $\Sigma^0$ , and  $\Sigma^+$ ), and the  $\Lambda$ . The difference between kaons and antikaons is relevant in this paper because the opposite hypercharge leads to significant differences, notably in the production, scattering, and absorption reactions. The heavier strange particles were not considered, due to lower production rates (they might be added in future extensions). The  $\Phi$  meson ( $s\bar{s}$ ) was not explicitly taken into account due to the lack of experimental data, even if it may play a role in the kaon-antikaon pair production. Some of its contributions are, however, hidden in the cross sections considered below, as the kaon-antikaon production cross sections (this may be the case for other resonances).

### B. Average nuclear potentials and characteristics

The mass, charge, isospin, decay modes, and half-life of each considered particle are taken from the Particle Data Group review [11]. The average nuclear potential is much more difficult to determine. Only some of the implemented particles are relatively well known ( $\Lambda$ ,  $K^+$ , and  $K^-$ ). Studies on the  $\Sigma$ 's potential are rather sparse and it is not yet clear whether the  $\Sigma$ 's potential is attractive or repulsive.

Because few experimental measurements exist, the average potentials for strange particles in INCL are considered as a constant over the nuclear volume, for sake of simplicity. In the case of the  $\Lambda$  particle, a dependence on the nuclear asymmetry [12] is introduced to improve hypernuclei physics. Typically, the potential used for the  $\Lambda$  is a 28 MeV attractive potential for the symmetric nuclei that grows up to a 41 MeV attractive potential for the largest asymmetries [ $(A - 2Z)/A = 0.25$ ]. The  $K^+$  and  $K^-$  average potentials have been relatively well studied and it is commonly accepted that the  $K^+$  potential is slightly repulsive and the  $K^-$  is strongly attractive [13]. The values retained in this paper are a 25 MeV repulsive potential for the  $K^+$  and a 60 MeV attractive potential for the  $K^-$ . It was decided to consider the same potential for  $K^0$  and  $\bar{K}^0$  as for  $K^+$  and  $K^-$ , respectively, with a slight correction (10 MeV) due to Coulomb repulsion. This choice is consistent with experimental measurements summarized in Ref. [13]. The potential for  $\Sigma$  particles is extremely difficult to measure because of nuclear effects. A repulsive potential of 16 MeV for all  $\Sigma$ 's is used in INCL based on a recent study [14].

In summary, the potentials used in this new version of INCL are

$$\begin{array}{ll} 25 \text{ MeV,} & K^+, \\ 15 \text{ MeV,} & K^0, \\ -60 \text{ MeV,} & K^-, \\ -50 \text{ MeV,} & \bar{K}^0, \\ 16 \text{ MeV,} & \Sigma's, \\ [-28, -41] \text{ MeV,} & \Lambda \text{ (Ref. [12]).} \end{array} \quad (1)$$

The difference between the potentials for kaons and antikaons potentials is not considered in other codes that we use for our comparisons. This produces significant differences in cross-section prediction near threshold energies.

### C. Interactions

As mentioned previously, the methods used to determine production, scattering, and absorption cross sections and the angular distributions of outgoing particles are described in detail in Ref. [9]. In Table I, we summarize the considered reactions based on experimental data.

In addition to the reactions listed in Table I, we include two other types of reactions, which are listed in Table II. First are the  $\Delta N$  reactions, which are expected to contribute significantly to strangeness production according to the study of Tsushima *et al.* [15]. The second type is the strangeness production in reactions where many particles are produced in the final state but no measurements are available. This second type is needed to get the correct inclusive strangeness production cross section. Since the cross sections given in Table II are coming from models, larger uncertainties are expected.

Strangeness production from baryonic-resonances-induced reactions seems to be negligible, or at least of second order, compared to the total strangeness production and the large widths of some resonances make their propagation questionable; therefore, the resonances are not propagated in INCL, except the  $\Delta(1232)$ , which has been intensively studied. This

TABLE I. List of considered reactions involving strangeness based on experimental data.

|                                |                                 |
|--------------------------------|---------------------------------|
| $NN \rightarrow N\Lambda K$    | $N\bar{K} \rightarrow N\bar{K}$ |
| $\rightarrow N\Sigma K$        | $\rightarrow \Lambda\pi$        |
| $\rightarrow N\Lambda K\pi$    | $\rightarrow \Sigma\pi$         |
| $\rightarrow N\Sigma K\pi$     | $\rightarrow N\bar{K}\pi$       |
| $\rightarrow N\Lambda K\pi\pi$ | $\rightarrow \Lambda\pi\pi$     |
| $\rightarrow N\Sigma K\pi\pi$  | $\rightarrow \Sigma\pi\pi$      |
| $\rightarrow NNK\bar{K}$       | $\rightarrow N\bar{K}\pi\pi$    |
| $\pi N \rightarrow \Lambda K$  | $NK \rightarrow NK$             |
| $\rightarrow \Sigma K$         | $\rightarrow NK\pi$             |
| $\rightarrow \Lambda K\pi$     | $\rightarrow NK\pi\pi$          |
| $\rightarrow \Sigma K\pi$      | $N\Lambda \rightarrow N\Lambda$ |
| $\rightarrow \Lambda K\pi\pi$  | $\rightarrow N\Sigma$           |
| $\rightarrow \Sigma K\pi\pi$   | $N\Sigma \rightarrow N\Lambda$  |
| $\rightarrow NK\bar{K}$        | $\rightarrow N\Sigma$           |

is justified *a posteriori* (see Sec. IV). However, the contributions of resonances as intermediate state of binary collisions are taken into account in an unresolved manner in the cross sections of the reactions summarized in Tables I and II.

#### D. Post-cascade treatment

The implementation of strangeness can lead to a new situation at the end of the intranuclear cascade; a hyperremnant can be produced in which at least one strange particle is still inside the target nucleus at the end of the intranuclear cascade. Therefore, we had to decide how to treat the remaining strange particle after the end of the cascade.

Owing to the repulsive average nuclear potential for the kaons, we decided to eject the trapped kaons at the end of the cascade and to correct their kinematics accordingly to their potential.

All  $\Sigma$ 's and antikaons have high absorption cross sections ( $N\Sigma \rightarrow N\Lambda$  and  $N\bar{K} \rightarrow \Lambda\pi$ ) at low energy. Therefore, we decided to absorb all of them when they are trapped inside the nucleus and to convert the excess of energy and the mass energy of the possible pions into nucleus excitation energy.

After kaon emission and  $\Sigma$  and antikaon absorption, the hyperremnant contains only protons, neutrons, and  $\Lambda$ 's. The hyperremnant is then de-excited using a new version of the ABLA07 code [2], which will be presented in a future paper.

TABLE II. List of the reactions involving strangeness and requiring information that has been taken exclusively from models. The  $X$  stands for all possible reactions excluding the reaction summarized in Table I.

|                                   |                           |
|-----------------------------------|---------------------------|
| $\Delta N \rightarrow N\Lambda K$ | $NN \rightarrow K + X$    |
| $\rightarrow N\Sigma K$           |                           |
| $\rightarrow \Delta\Lambda K$     | $\pi N \rightarrow K + X$ |
| $\rightarrow \Delta\Sigma K$      |                           |
| $\rightarrow NNK\bar{K}$          |                           |

### III. VARIANCE REDUCTION METHODS

Strangeness production is a rare process in spallation reactions, especially for energies below 5–6 GeV. However, the user of INC models might have an interest in strangeness physics. Therefore, to have enough statistics, a calculation will require a large number of events, an event being defined as the simulation of one cascade in INCL. This would take a lot of computational time and result in a large amount of useless information: taking a lot of space on a hard drive and slowing down the analysis.

The solution proposed here is to use variance reduction methods, which are based on the importance sampling method. In INCL + +6, we implemented a variance reduction scheme (VRS), which artificially increases the statistics for the observables linked to the strangeness production, keeping the calculation time and the output file size unchanged.

Note that INCL has an option enabling to write into the final ROOT file only if a certain condition is fulfilled. This makes the output size problem marginal for us. However, the information that does not fulfill the condition is lost with this option. With the newly implemented VRS all the information is kept, which we consider a major benefit.

This section describes the variance reduction scheme and its operation. The reader only interested in the calculation results of INCL dealing with strangeness production can skip this section. The variance reduction method developed here, and used in INCL, aims at getting results with lower uncertainties, and in some cases is the only way to get results within a reasonable computation time, but it does not change the conclusions. Readers eager to understand the method will find in the following some generalities (Sec. III A) presenting the topic, the constraints and the difficulties within INCL, the scheme developed, and the technical aspects (Sec. III B), the case of correlations (Sec. III C), the reliability of the method (Sec. III D), and some examples of uses of the VRS (Sec. III E). Readers interested by the main lines and aspects, and not by the method itself, can go directly to Secs. III D and III E.

#### A. Generalities

Basically, a fully operational important sampling method consists of two successive steps. In a first step, the rules of the simulation, i.e., the physics ingredients, are modified (biased) to improve the sampling. In our approach, we increase the strangeness production cross section. In a second step, the simulation result is accordingly corrected. The goal of the variance reduction methods is to obtain the true observable of interest (e.g., a cross section) with a reduced variance and therefore with reduced uncertainties within the same computational time.

For this study, all the processes associated with strangeness production are of interest. Therefore, the VRS of INCL was developed to increase the realization probability for all of them, independently of the observable. However, even though the current VRS globally modifies the probabilities for processes involving strangeness, other schemes are in principle possible. The scheme could also be adapted to be

more restrictive (e.g., forward kaon production) or to work for another particle type (e.g.,  $\eta$  meson production) or physics (e.g., peripheral collisions).

The reason for the low global strangeness production rates in INCL is the low elementary strangeness production cross sections. For example, the strangeness production represents 0.014% (0.15%) of the total cross section in proton-proton collision at kinetic energies of 2(3) GeV. Therefore, in a first step, the VRS modifies the hadron-hadron cross sections to increase the strangeness production. In a second step, it corrects *a posteriori* the bias introduced to obtain the unbiased observable estimators with reduced uncertainties.

In the general case, the second step (i.e., the bias reversal) consists in determining the ratio of the probability to make an observation in the nonbiased version to the probability to make the same observation in the biased version. This ratio is called the importance or the weight and can be calculated for a complete event (a cascade), a specific particle, or for any observation during the event. Next, the importance is used to weight the contribution of observations. This number gives the information about how much this observable is biased. The importance for an observable  $X$  can be written as

$$W_X = \frac{P(X|\text{no-bias})}{P(X|\text{bias})}. \quad (2)$$

This expression of the importance leads to a first requirement that events contributing to a variable of interest should have a nonzero probability of realization in the biased version (the version using the variance reduction scheme); otherwise, it will result in an arithmetic exception. In other words, every strange event in INCL that can be produced in the standard version must be attainable in the version using the variance reduction scheme. Only the probability of realization can be changed. Because of this constraint and because it is not trivial to know whether strangeness production could occur later during an event, no channel cross section can be reduced to zero at the binary collision level in INCL.

The treatment of a particle during an INCL event can be subdivided into a three-step cycle. First, the particles are propagated freely inside the nucleus. This step ends when two particles collide. A collision occurs when the distance between two particles is below a maximal interaction distance based on their total interaction cross sections. Second, the type of the reaction of the binary collision is randomly chosen based on the respective reaction cross sections. In the last step, the phase space and the charge repartition is randomly generated either based on differential cross sections (if available) or on phase space generators. Then, the cycle is repeated until the end of the intranuclear cascade.

The propagation along straight trajectories between collision events is deterministic. According to this, an artificially decrease of a total interaction cross section can lead to a situation of particles flying past each other where they would have collided using the original cross section. From such an event onwards, the subsequent cascade is outside the universe of possibilities based on the unbiased total interaction cross section. Therefore, the importance of this event would be null according to Eq. (2). The same argument can be made for an increased total interaction cross section. Hence, the

total interaction cross sections for binary collisions must be preserved.

Thanks to the random treatment of the reaction choice and the phase space generation, both steps can be biased; changing the probability of realization. In our case, only the step selecting the type of reaction needs to be biased to increase the global strangeness production. However, if the user is interested in the production of a particle in a specific phase space (e.g., backward production), phase space generation can be biased under minor modification in the code.

The two constraints, no cancellation of any particular channel and the preservation of total cross sections for binary collision, will be crucial for the INCL variance reduction scheme.

With the variance reduction scheme, the new weighted estimators and the associated variances of an observable  $X$  [16] are given by

$$E(X) = \frac{\sum_{i=1}^M w_i x_i}{N}, \quad (3)$$

$$V(X) = \frac{\sum_{i=1}^M [w_i x_i - E(X)]^2}{M}, \quad (4)$$

with  $w_i$  being the importance of the  $i$ th observation  $x_i$  of the observable  $X$  and  $N$  being the total number of events. The natural summation is such a case corresponds to summing over the events ( $M = N$ ), with  $w_i$  the importance of the  $i$ th event and  $x_i$  the number of observations corresponding to the observable  $X$  (e.g., kaons) in the corresponding event.

An alternative summation is actually possible and preferred in our case. We can sum over the final particles ( $M \neq N$ ) with  $w_i$  the importance of the  $i$ th particle and  $x_i = 1$  or 0 if the particle corresponds to the observable or not, respectively. With the natural summation, we estimate the mean value of the number of strange particles produced per reaction. In the alternative summation, we estimate the number of strange particles produced and we normalize by the number of reactions. It clearly appears the observables estimated in the two approaches correspond.

For INCL, it is simpler to sum over the particles considering the output generated. Moreover, for sake of simplicity, the relative uncertainties displayed in figures below do not derive from Eq. (4) but are calculated using the formula

$$\text{relat. unc.} = \frac{\sqrt{\sum (w_i x_i)^2}}{\sum w_i x_i}, \quad (5)$$

which is equivalent to the standard relative uncertainty equals to the inverse square root of the number of observations in the Monte Carlos simulations and which is an approximation of the relative uncertainty that can be obtained using Eq. (4).

It is worth emphasising that, the observable  $X$  can be anything. For example, it can be the presence or absence of a particle of a certain type, the number of particles of a certain type produced during the cascade, the presence of many correlated particles, an entire event, or a part of an event.

The minimum variance for a given number of events is achieved when all the strange particle importances are equal according to Eqs. (3) and (4). However, these importances are not always equal in INCL for reasons explained further below.

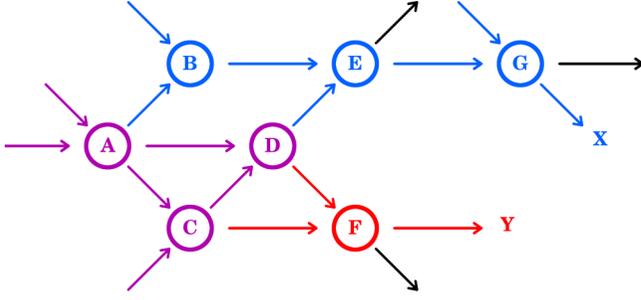


FIG. 1. Simplified example of an intranuclear cascade represented as a time ordered graph. Circles and arrows represent binary collisions and the propagation of the particles, respectively. The blue (red) and purple part of the graph is the history of the  $X$  ( $Y$ ) particle.

Therefore, the objective is to keep them as close as possible to minimize the variance and therefore uncertainties.

### B. Variance reduction scheme in INCL

For a better understanding of the VRS used with INCL, a cascade can be seen as a time ordered graph where the edges (arrows) and vertices correspond to the particles propagating and to the binary collisions, respectively. Thus, a cascade is fully defined by the set of its vertices. The information about what happened before (e.g., the projectile type, the impact parameter, etc.) is hidden in the definition of the vertices initial state. A schematic example is displayed in Fig. 1.

As explained in the previous subsection, the VRS must increase the strangeness production in a way that minimizes the spread of the strange particle importances to achieve minimum variance and therefore to obtain the smallest possible uncertainties for a given number of events.

In our variance reduction scheme, if the particle  $Y$  shown in Fig. 1 is a strange particle, then the event containing the set of vertices  $A$ ,  $C$ ,  $D$ , and  $F$  should have an increased probability of realization. If the particle  $X$  is a strange particle, then the events containing the set of vertices  $A$ ,  $B$ ,  $C$ ,  $D$ ,  $E$ , and  $G$  should be promoted. If both particles are strange particles, then both paths should be promoted in the same way. Two solutions can be considered. The first one is to promote the vertices  $A$ ,  $C$ , and  $D$ , which are common to both paths, and to continue as a standard cascade. In such an approach, it actually means only the first vertex can be biased. The problem is that we do not know whether a strange particle will be produced in any of the binary collision represented by a later vertex. As a consequence, only strangeness production in the first collision can be promoted. This would lead to differences between the importances for the events that produced strangeness in the first collision, and the events that produced strangeness in secondary collisions. Thus, it would be a new source of variance, what should be avoided. The second solution is to bias each vertex (each elementary binary collision) along the entire cascade in such a way the same importance is obtained for every strange particle. The second solution was chosen for INCL and the importance associated to each final particle is provided in the final ROOT file.

In our solution, the promotion of strangeness production can change from one binary collision to another. However, if different channels are producing strangeness in a given binary collision, then they should be promoted in the same way to get the same importance as the other strange particles whatever the channel chosen. Once again, this should be done to minimize the variance.

With the INCL variance reduction scheme, we introduce a new input parameter to INCL. This input is a scalar, which is used to define the required importance of strange particles. This allows to have the same importance for strange particles in the different events, which are independent. This input scalar will be called the *bias factor* in the following. The VRS implemented in INCL tries to force the final importances of strange particles to be the reciprocal value of the bias factor. Consequently, the bias factor is a multiplication factor of the probability to create a strange particle. However, the effective increase is lower in some cases. This is discussed in Sec. III E with the way to optimize the bias factor.

Before explaining how the VRS forces the final importances of strange particles to be a specific value, we describe how the importances of particles are calculated. Doing so, we start with the general formula,

$$P(N \cap M) = P(N|M) \times P(M), \quad (6)$$

with  $P(N \cap M)$  the probability of realization of  $N$  and  $M$ ,  $P(N|M)$  the probability of realization of  $N$  knowing that  $M$  is realized, and  $P(M)$  the probability of realization of  $M$ .

In Fig. 1 the vertex  $B$  can happen only if the vertex  $A$  has been realized, therefore,

$$P(A|B) = 1. \quad (7)$$

Combining Eqs. (6) and (7), we get

$$P(A \cap B) = P(A|B) \times P(B) = P(B). \quad (8)$$

For the same reason,

$$P(A|F) = P(C|F) = P(D|F) = 1, \quad (9)$$

$$\Rightarrow P(A \cap C \cap D \cap F) = P(F). \quad (10)$$

Therefore, in the version using the VRS the probability of producing the final particle  $Y$ , which is produced in vertex  $F$ , must be corrected by the importance  $W_Y$ , which is given by

$$\begin{aligned} W_Y = W_F &= \frac{P(F|\text{no-bias})}{P(F|\text{bias})}, \\ &= \frac{P(A \cap C \cap D \cap F|\text{no-bias})}{P(A \cap C \cap D \cap F|\text{bias})}, \\ &= \frac{P(A|\text{no-bias})}{P(A|\text{bias})} \times \frac{P(C|A, \text{no-bias})}{P(C|A, \text{bias})} \\ &\quad \times \frac{P(D|A, C, \text{no-bias})}{P(D|A, C, \text{bias})} \times \frac{P(F|A, C, D, \text{no-bias})}{P(F|A, C, D, \text{bias})}, \\ &= \frac{\sigma(A)/\sigma_{\text{tot}}(A)}{\sigma'(A)/\sigma_{\text{tot}}(A)} \times \frac{\sigma(C)/\sigma_{\text{tot}}(C)}{\sigma'(C)/\sigma_{\text{tot}}(C)} \\ &\quad \times \frac{\sigma(D)/\sigma_{\text{tot}}(D)}{\sigma'(D)/\sigma_{\text{tot}}(D)} \times \frac{\sigma(F)/\sigma_{\text{tot}}(F)}{\sigma'(F)/\sigma_{\text{tot}}(F)}, \end{aligned}$$

$$\begin{aligned}
&= \frac{\sigma(A)}{\sigma'(A)} \times \frac{\sigma(C)}{\sigma'(C)} \times \frac{\sigma(D)}{\sigma'(D)} \times \frac{\sigma(F)}{\sigma'(F)}, \\
&= \text{CSR}(A) \times \text{CSR}(C) \times \text{CSR}(D) \times \text{CSR}(F), \quad (11)
\end{aligned}$$

with  $\text{CSR}(I)$  the cross-section ratio of vertex  $I$ ,  $\sigma(I)$  and  $\sigma'(I)$  the standard and biased cross section of the reaction that took place in vertex  $I$ , and  $\sigma_{\text{tot}}(I)$  the total interaction cross for the vertex  $I$ .  $W_Y = W_F$  because the propagation between collisions is not biased.

The cross-section ratios are easily determined during an event since both terms in the ratio  $\sigma/\sigma'$  are known for all the possible reactions when a binary collision happens. Some examples are discussed in the Appendix.

In the VRS of INCL, the cross-section ratio of the vertices are stored. Therefore, whenever a binary collision happens, the cross-section ratio of the previous vertices are known. This allows to calculate what should be the cross-section ratio of the vertex representing the aforementioned binary collision [e.g., the vertex  $F$  for the case in Eq. (11)] to match the importance of the outgoing particles (the particle  $Y$  in our example) to the desired importance. This calculated cross-section ratio defines how the strangeness production should be promoted (see examples in the Appendix).

In the special case of a binary collision between a strange particle and another particle, a strange particle will be present in the final state regardless of the channel chosen in INCL. Therefore, the strangeness cannot be promoted and the cross-section ratio will be equal to 1. Such a binary collision can result in a strange particle that does not have the aimed importance of the strange particles. No solution was found to solve this problem. However, the dispersion of particle importances due to this phenomenon in most of cases is marginal thanks to the precautions discussed below and does not introduce a new significant source of variance.

The main source of importance dispersion for strange particles in INCL, which is itself source of variance, is due to the two constraints discussed before. First, the nonstrange cross sections cannot be null but no lower bound is fixed. Therefore, when a reaction cross section is drastically reduced but the corresponding reaction is chosen anyway, the cross-section ratio of the binary collision is extremely large. Second, the total cross section must remain unchanged. This fixes a lower bound for the cross-section ratio of a vertex that produces strangeness ( $=\sigma^{\text{strange}}/\sigma^{\text{tot}}$ ). Therefore, if the cross-section ratio required for a vertex is below this limit, then the effective strangeness promotion will not correspond to the one required.

The first constraint can lead to extremely large cross-section ratios, which cannot be counterbalanced by the following ones because of the second constraint. If a strange particle is produced on such a path, then it would result in a strange particle with a large importance compared to others and, therefore, it will strongly contribute to the observables. Such high contributions produced by single particles will make the convergence slower, lead to variance *jumps* and, therefore, to large uncertainties for the observables (see Sec. III E). If pronounced *jumps* are seen, then all the observables estimated with the corresponding INCL results should not be trusted. In such a case, the associated uncertainty estimated accordingly

to Eq. (4) or Eq. (5) might be highly underestimated because of the bad sampling. This situation means the bias factor was chosen way too large and paths important for strangeness production were suppressed too much.

To avoid these problems, a compromise should be found between the increase of the statistics, obtained using high bias factors, and the maintain of equal importances for strange particles, which is most easily achieved using small bias factors.

Therefore, a safeguard was implemented in the variance reduction scheme. This safeguard aims at optimising the convergence efficiency for observables of interest by modifying the bias factor used. The safeguard does it by preventing the decrease of channel cross sections below the half of the initial cross sections. Therefore, a vertex cross-section ratio cannot be higher than 2. This strongly limits the product of vertex cross-section ratios for a given history and it will be easier to counterbalance it. At the end of an event, this procedure strongly reduces the variance *jumps* even if the bias factor is chosen too high. A less restrictive safeguard has been tested but the actual one presents better compromises. However, this safeguard is not perfect. If a particle has a history with numerous vertices with cross-section ratios between 1 and 2, then it can have a large importance anyway and will result in variance *jumps* and a slow convergence.

### C. Event importance reversal

As previously mentioned, Eq. (3) is used to estimate the observables. Two types of summation can be used: a sum over the events using the importance of the events where  $x_i$  is an integer (not used in our case) or a sum over the particles using the importance of the particles where  $x_i = 0$  or 1. However, the dynamic adjustment of the bias factor depending on the history of events and their importances introduces some sort of dependency between particles of a same event. If working at the level of particle importances as we do, Eq. (3) is probably not fully correct any more, since correlations have been taken into account (e.g., when looking at cross section of hyperon emission in coincidence with a kaon).

There is an alternative possible type of summation for Eq. (3). The summation can be done over the particles ( $x_i = 0$  or 1) but using the event importances instead of the individual particle importances.

The event importance is equal to the product of every vertex cross-section ratio of the entire event. It is, *a priori*, different from the importance of an observation  $X$  (let us say a particle) because it includes the contribution of extra vertices, which are not on the path of the particle  $X$ , and are therefore not relevant. However, the expected value of the cross-section ratio of a vertex  $A$  is

$$\begin{aligned}
E[\text{CSR}(A)] &= \sum_{\text{reac}} \frac{\sigma(\text{reac})}{\sigma'(\text{reac})} \times P(\text{reac}|\text{init. state, bias}), \\
&= \sum_{\text{reac}} \frac{\sigma(\text{reac})}{\sigma_{\text{tot}}} = 1, \quad (12)
\end{aligned}$$

with the same notation as in Eq. (11). Noteworthy, this equation is true only if no reaction of the nonbiased version is

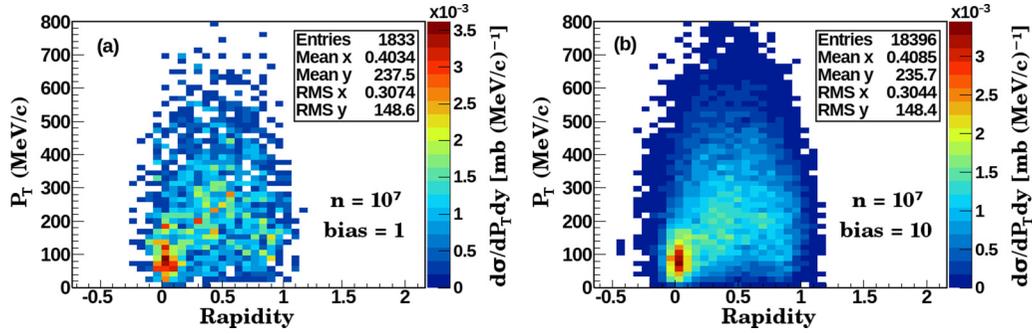


FIG. 2.  $\Lambda$  transverse momentum versus rapidity distribution in  $p(1.7 \text{ GeV}) + \text{Ca}$  collision. Both plots are obtained using the same number of events ( $10^7$ ). (a) No bias. (b) Bias factor = 10.

forbidden in the biased version, which is the case in INCL. The same argument can be used for any substructure in a cascade, which have no constraints on its final state. In particular, the extra vertices, which are not on the path of the particle  $X$ , have no constraint on their final state, since they play no role in the production of the particle  $X$ . Therefore, the contribution of extra vertices is statistically null and the expected value of the event importance knowing the particle  $X$  has been produced is equal to the importance of this particle  $X$ .

Then, it is trivial to prove that the expected value of the estimator is the same using  $w_i [=E(w_{\text{event}})]$  or  $w_{\text{event}}$  taking into account  $x_i = 0$  or 1 in Eq. (3) when using the summation over the particles.

The event importance is provided in the final ROOT file. The convergence being slower or equal, it is recommended to use the particle importance for the reversal of independent observables. However, it was decided to conserve only a scalar importance for each particle and a event importance to limit the quantity of information recorded in the output file. Doing so, the correlations are lost. Therefore, if the observable is a coincidence production, then the event importances must be used regardless of the type of summation. Additionally, if INCL is connected to another program (e.g., a transport code) that is not able to keep track of the importance of the individual particles, then this program can use the event importance to weight its own results.

#### D. Validation

As a first test of whether the VRS works reliably and as expected, the results obtained using the variance reduction method are compared with results obtained using the nonbiased version. Not only the observables obtained with the VRS must be the same as the observables obtained without it (no residual bias), but the uncertainties must also be reduced for the same calculation time (variance reduction). Various illustrative comparisons are discussed in the following.

A comparison of simulations using or not the VRS indicates that the calculation time is largely independent of the bias factor used. Therefore, the number of events can be used as a measure of the computational time and/or to compare the efficiency of calculations using different bias factors.

Figure 2 shows a comparison of calculations with and without bias. The two calculations have identical inputs except for the bias factor (1 and 10, respectively). It can be clearly seen that the calculation using the bias factor = 10 [Fig. 2(b)] produced a much more precise distribution using the same binning thanks to the statistic increased by a factor around 10. Additionally, no significant difference shows up for the amplitude, the mean values, and the standard deviations displayed for both bias factors in the frames. Consequently, the objectives of the VRS are perfectly fulfilled.

A test of convergence was carried out for the reaction  $p(1.7 \text{ GeV}) + {}^{12}\text{C}$  using different bias factors (no bias, 10, and 100) using 4500 simulations with various numbers of events per calculation. The observable chosen to test the convergence efficiency is the  $K^+$  mean momentum. What is considered as the true value for the  $K^+$  mean momentum was estimated with an additional calculation using no bias and  $10^9$  events, which resulted in an uncertainty of 0.1% for the true value. We then compared the estimator of the  $K^+$  mean momentum for the 4500 simulations to the true value. The absolute value of the difference between an estimator and the true value is the error. Figure 3 shows the mean relative error for the different bias factors as a function of the number of events, which corresponds to a computational time, which is independent on the bias factor. Fits of the form  $a n^b$  are plotted and the parameters and their uncertainties are shown.

Since the true value has been estimated with an uncertainty of 0.1%, a mean relative error of below 0.001 cannot be interpreted in Fig. 3.

A horizontal line was added in Fig. 3. It represents a relative error of 1.5%. According to this, the precision of 1.5% is obtained with about  $5 \times 10^6$  events without the VRS and with only  $5 \times 10^5$  ( $1.8 \times 10^5$ ) events with a bias factor 10(100). Therefore, the 1.5% precision is reached with about 10 times less events using a bias factor 10, while using a bias factor of 100 the gain in time is not 100, but only 30. This reduction of the computational time needed to reach the same precision is called the effective bias factor. The effective bias factor 30 for the bias factor 100 indicates that the optimal bias is probably around 30 and the safeguard (see Sec. III B) forced the effective bias factor to be closer to this value.

In Fig. 3, a decreasing mean relative error that is inversely proportional to the square root of the number of events is observed for every bias factor used. This includes the bias

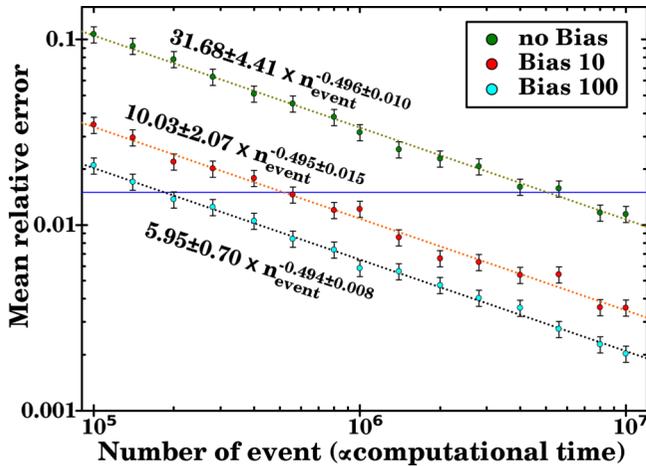


FIG. 3. Mean relative error of the  $K^+$  mean momentum estimator as a function of the number of events. The number of events corresponds to a simulation time. The true value taken for the kaons mean momentum is estimated using a  $10^9$  unbiased event calculation. The considered reaction is  $p(1.7 \text{ GeV}) + {}^{12}\text{C}$  with *bias factor* = 10 (red), 100 (cyan) and not using a variance reduction method (green). Dotted lines are fits of the form  $a n_{\text{event}}^b$ . The horizontal blue line is here to guide eyes (see text).

factor that is clearly above the optimal bias factor. Additionally, the dispersion of the results for the 4500 calculations follows a normal distribution around the true  $K^+$  mean momentum (estimated using the calculation with  $10^9$  events) with respect to their statistical uncertainties. This confirms that every calculation converges to the same limit regardless of the bias factor used. Therefore, the VRS (bias + reversal) does not introduce any bias in the final observables. They all converge to the true value.

### E. Examples

Let us review the important vocabulary introduced in this section:

*Bias factor*: The augmentation of the statistics of strange particles required at the beginning of the calculation.

*Effective bias*: The real augmentation of the statistics of strange particles observed.

*Optimal bias factor*: The bias factor minimising the uncertainties linked to strange particles for a given number of events.

The use of the VRS of INCL ++6 requires two steps in addition to the standard use of INCL. Firstly, the user must provide the bias factor and second, the user must weight the final observables with the corresponding importances stored in the output file.

The *a priori* bias factor optimisation is not trivial because the optimal bias factor strongly depends on the initial parameters, like the target size and the projectile kinetic energy, but it depends also on the final observables. However, the bias factor does not need to be perfectly optimised; a sensible choice for the bias factor already helps significantly to speed up the convergence of estimator of observables involving strangeness

production. By starting with a sensible choice, the safeguard presented above is able to finalise the optimisation. A work on the bias optimisation was carried out. However, the observables multiplicity and the possibilities of initial parameters being too large, a simple way to determine the optimal bias factor cannot be provided by the authors. The simplest way to evaluate the optimal bias factor is an interpolation taking into account the energy and the mass number of the target, given that it is safer to be below the optimal bias factor. Examples of sensible choices can be found in Sec. IV for each figure plotted.

In Fig. 4, we show the results of the study on the limits of the variance reduction methods implemented in INCL. The limits discussed are problematic only for some *extreme* (high projectile energy and large target mass) reactions where the choice of the bias factor was not sensible. However, this case illustrates perfectly the problem of the importance dispersion, which produces large uncertainties in terms of variance *jumps* even when the statistics is increased.

Here, we studied the mass distribution of hyperremnants in  $p(10 \text{ GeV}) + {}^{208}\text{Pb}$  collisions. The bias factors used were 1 (no variance reduction method used), 2, 5, 10, 20, and 50. The number of events was the same ( $10^7$ ) for all the calculations and the computational time was equivalent. The Figs. 4(a) and 4(b) show the results obtained using the variance reduction scheme, with a bias factor 2 and 10, respectively, together with the result from the standard calculation. The increase of the statistics ( $n_{\text{Hyp}}$ ) and the normalized mean uncertainty as a function of the bias factor are displayed in Figs. 4(c) and 4(d). The left panels clearly show the minimal variance (i.e. the smallest uncertainties) is not achieved with the highest statistics in this case.

An interpolation (shown as a blue line) of the normalized uncertainties predicts an optimal bias factor of around 2.5. This is illustrated by the significantly better description of the hyperremnants spectrum obtained using a bias factor 2 compared to the calculation using a bias factor 10. The large uncertainties obtained using a bias factor 10 are purely linked to the dispersion of hyperremnant importances. It is interesting to note that the interpolation of the effective bias factor (blue line) in Fig. 4(c) deviates from the nominal bias factor (dashed brown line) around the optimal bias factor (2.5). The effective bias factor is reduced by the safeguard discussed before and the reduction starts to be significant slightly above the optimal bias factor. This demonstrates that the safeguard works well in this case, though it is not perfect.

Generally, variance *jumps* result in a global underestimation compared to the standard calculation and in some strongly overestimated bins. This can be dangerous if an underestimated bin is studied because the error bars for this bin are not significantly large but they can be far from the true value [see Fig. 4(b)]. This is why it is crucial to minimize the importance dispersion. Remember, however, that every calculation converges to the same limit. Once again, the variance *jumps* just make the convergence slower. The problem is that the variance, and by extension the error bars, might be underestimated because of the bad sampling.

Figure 5 shows a reaction in which the number of nucleons in the target is too low to produce variance *jumps* thanks to the

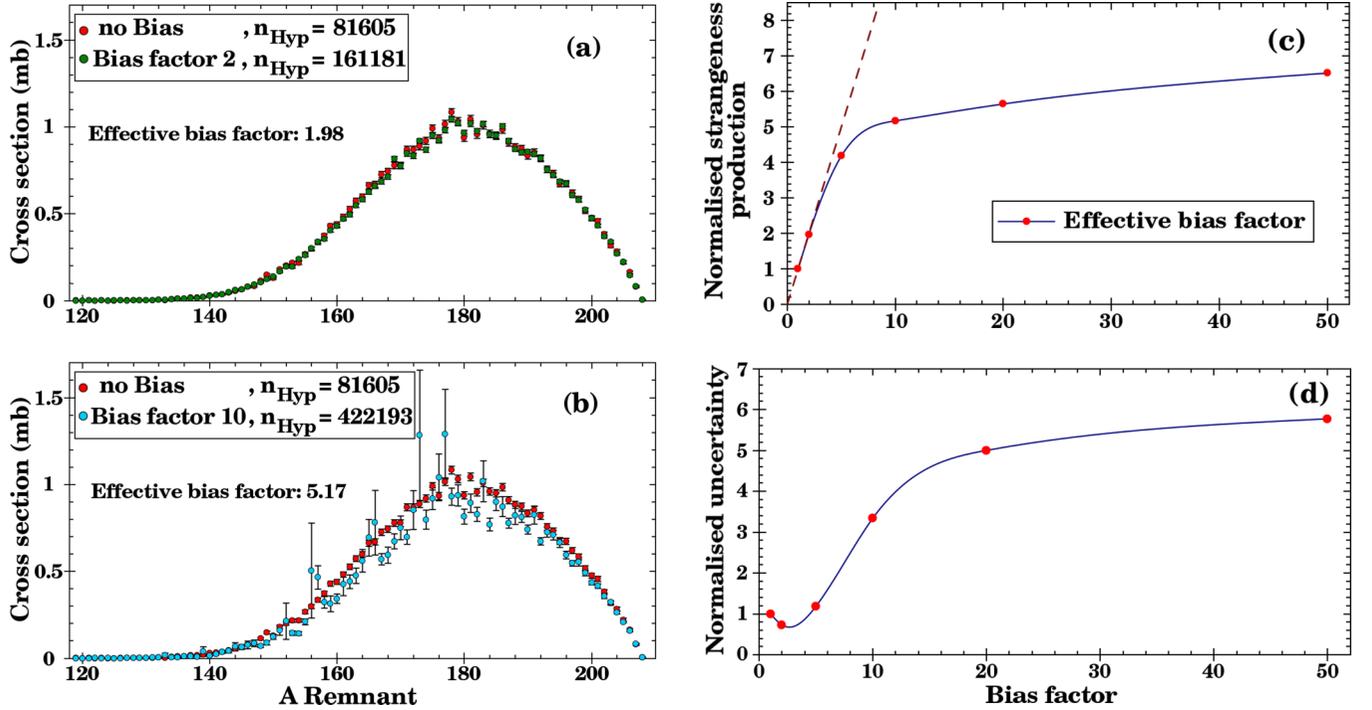


FIG. 4. Left: Hyperremnant mass distribution in  $p(10 \text{ GeV}) + {}^{208}\text{Pb}$  collisions with  $10^7$  shots. Calculations with bias factors = 2 (a) and 10 (b) are compared to the calculation without VRS used. Right: Evolution of uncertainties and statistics as a function of the bias factor (see text). The uncertainties and statistics are normalized to 1 for bias factor=1 (no VRS used). The dashed brown line in Fig. 4(c) represents the perfect case where the effective bias factor equals the nominal bias factor.

safeguard regardless of the bias factor. The number of target nucleons being low, this leads to short intranuclear cascades with a low number of vertices. Additionally, the cross-section ratio of vertices is limited by the safeguard. This results in strong constraints for the importances of final particles.

In such cases associated to low energy, the safeguard matches automatically the effective bias factor to the optimal bias factor when the bias factor has been chosen too high. This is well illustrated in Fig. 5. Although the nominal bias factor ( $10^4$ ) is clearly above the optimal bias factor, which is 60, the spectrum obtained using the VRS exhibits no variance *jump*.

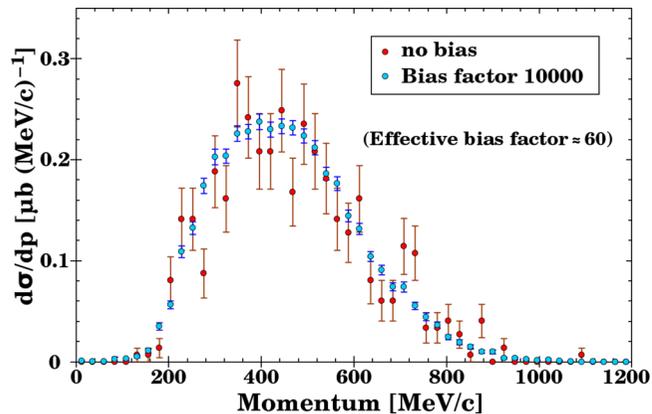


FIG. 5.  $K^+$  momentum in  $p(1.6 \text{ GeV}) + {}^{12}\text{C}$  with  $10^7$  events. INCL using a bias factor  $10^4$  (blue) is compared to INCL using no bias (red).

#### IV. RESULTS

In this section, results of INCL ++6 calculations are compared to experimental data and to other models to facilitate the interpretation and the analysis. The different strengths and weaknesses observed in INCL ++6 are discussed. Hypernucleus production is not presented here because it depends strongly on the de-excitation stage of the spallation reaction. A forthcoming paper will be dedicated to the strange degree of freedom in the de-excitation model and the hypernucleus production will be discussed in this publication. Here, we focus on the intranuclear cascade and single strange particle production.

The studied particles are the charged kaons ( $K^+$  and  $K^-$ ), the  $\Lambda$ , and the neutral kaons. The  $K^+$  was clearly the most studied particle in the past. Experimental data exist near the threshold, and even in the subthreshold region, up to high energies ( $\approx 14 \text{ GeV}$ ). Additionally, various targets were studied and the emitted  $K^+$  were observed at different angles. Moreover, the  $K^+$  is the only particle together with the  $K^0$  in our energy range that carries a positive hypercharge. Therefore, every strangeness production results in kaon production. Because the production modes of the two kaons are similar, studying the  $K^+$  also gives an estimate of the reliability of the total strangeness production in INCL ++6. The  $K^-$  was less intensively studied because of the lower production rate (about two orders of magnitude compared to the  $K^+$ ). However, double differential experimental data can be found in the literature. The  $\Lambda$  and the neutral kaons having no electric charge, their detection is more complex and less

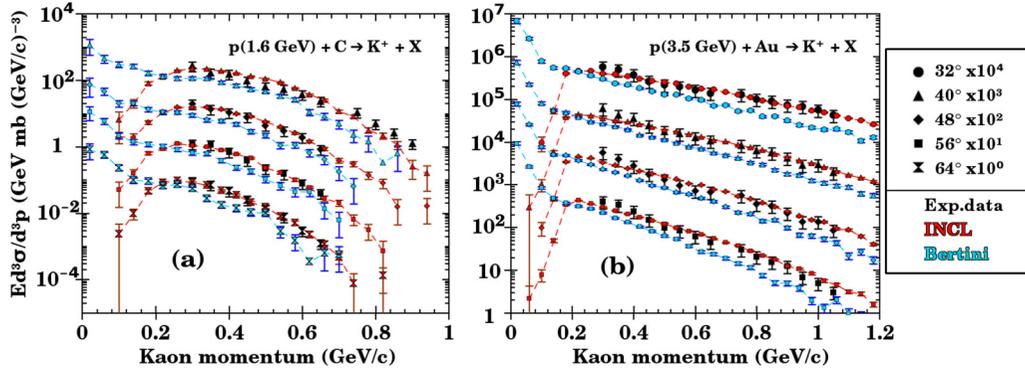


FIG. 6.  $K^+$  invariant cross section for various angles in (a)  $p(1.6 \text{ GeV}) + \text{C}$  and (b)  $p(3.5 \text{ GeV}) + \text{Au}$  collisions. Experimental data [17] (black symbols) are compared to INCL (red) and to the Bertini cascade model [18] (blue). Bias factor used: 10.

experimental data are available. However, their analysis can help to understand the different processes in competition in the strangeness production.

The bias factors (see Sec. III) used to obtain the INCL + +6 results are given in the captions of the corresponding figures.

### A. KaoS

The Kaon Spectrometer (KaoS) [17] experiment was performed at the heavy-ion synchrotron SIS at GSI in Darmstadt. The KaoS collaboration measured the  $K^+$  and  $K^-$  production in  $p + \text{C}$  and  $p + \text{Au}$  collisions at 1.6, 2.5, and 3.5 GeV proton beam kinetic energies. The kaon momentum was measured from  $p_{\text{lab}} = 0.3$  to 1.1 GeV/c.

Most of the data measured for the  $K^+$  production are well described by INCL. Figure 6 shows the  $K^+$  production invariant cross section in two configurations measured by the KaoS collaboration in comparison to INCL and to the Bertini cascade model [18]. Either for Fig. 6(a) with  $K^+$  production near threshold on a light nucleus (carbon) or for Fig. 6(b) with  $K^+$  production at higher energy on a heavy nucleus (gold), the results of INCL match very well the experimental data. The comparison with the KaoS data indicates a reliable total strangeness production cross section for a relatively large range of nuclei, energies, and angles. It can be seen that the Bertini cascade model gives a shape similar to INCL in the momentum range of the experimental data but underestimates them by roughly 40%. This difference can be explained by the  $\Delta$ -induced strangeness production, which is not included in the Bertini model. Figure 6 depicts for the low momenta region huge differences between the predictions from the INCL and Bertini models. This is due to the different values used for the  $K^+$  potential. In our approach, the  $K^+$  repulsive potential in INCL reduces drastically the invariant cross section at low momenta. Experimental data at lower momenta would help to test the  $K^+$  potential.

Figure 7 shows the invariant cross section for  $K^-$  production for the reaction proton on Au at 3.5 GeV. It can be seen that the data are well described by INCL for momenta above 0.5 GeV/c. However, below this energy, the cross section is clearly underestimated and INCL does not reproduce the shape of the experimental data. In Ref. [17], the authors

concluded that the  $K^-$  production is mostly due to the  $NY \rightarrow K^-NN$  reaction and to a lesser extent to the  $\pi Y \rightarrow K^-N$  reaction. Both reactions are not considered in INCL because of the lack of experimental data. This is probably the explanation. The low energy production being dominated by secondary reactions and the high-energy production by primary reactions, our result seems consistent with a predominance of antikaon production through strangeness exchange ( $NY \rightarrow K^-NN$  and  $\pi Y \rightarrow K^-N$ ) for low momenta and through direct production ( $NN \rightarrow NNK\bar{K}$ ) for high momenta. The Bertini cascade model, which includes the  $NY \rightarrow K^-NN$  (but not the  $\pi Y \rightarrow K^-N$ ) reaction, shows a significantly better description over the entire momentum range covered by the experimental data. This would be in favour of adding the extra strangeness exchange reactions even when considering the bad quality of experimental data.

### B. ITEP

The study carried out at the Institute of Theoretical and Experimental Physics (ITEP) accelerator in Russia [19] measured the  $K^+$  production in proton-nucleus collisions. The nuclei studied were Be, Al, Cu, and Ta. This choice covers

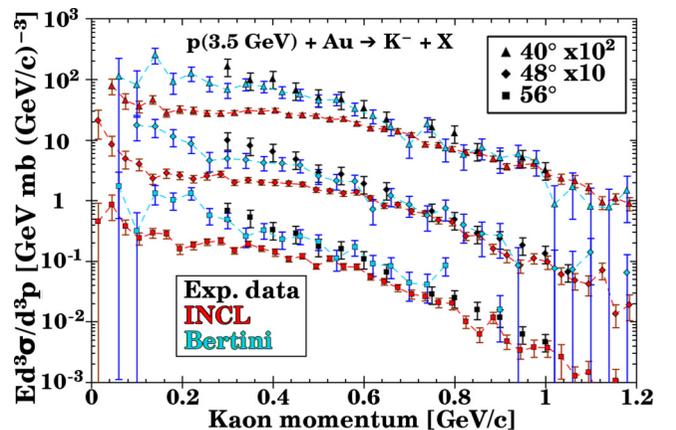


FIG. 7.  $K^-$  invariant cross section at  $40^\circ$ ,  $48^\circ$ , and  $56^\circ$  in  $p(3.5 \text{ GeV}) + \text{Au}$  collisions. Experimental data [17] (black) are compared to INCL (red) and to the Bertini cascade model [18] (blue). Bias factor used: 10.

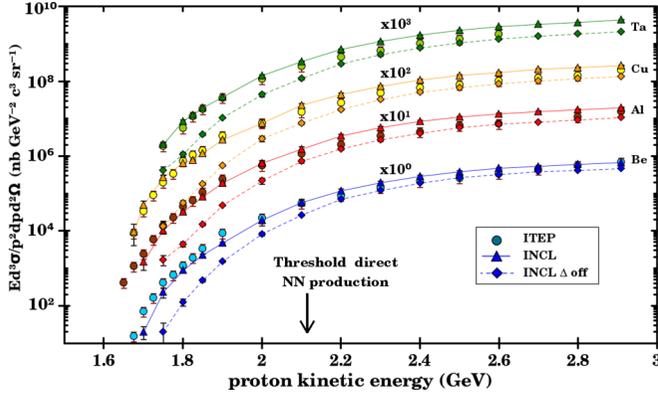


FIG. 8.  $K^+$  invariant cross section in  $p + A \rightarrow K^+ + X$  reactions for kaons emitted with a momentum of  $1.280 \pm 0.014$  GeV/c at  $\theta = 10.5^\circ$ . The experimental data from Ref. [19] (circles) are compared to INCL with (triangles) and without  $\Delta$ -induced strangeness production (squares). Bias factor used: 50 for INCL without  $\Delta$ -induced strangeness production for  $T_p < 2$  GeV, 20 otherwise.

a mass range from  $A = 9$  to 181 ( $Z = 4$  to 73). The projectile energy range from 1.65 up to 2.91 GeV. The experiment measured the production of  $K^+$  with a momentum  $p = 1.280 \pm 0.014$  GeV/c and with an emission angle  $\theta = 10.5^\circ$ . This very specific phase space constraint was used to drastically reduce the contribution of the  $K^+$  production in  $\pi N \rightarrow YK$  secondary reactions (see Ref. [19] for details). This simplified their analysis by considering only the  $NN \rightarrow NYK$  primary reactions.

In our analysis, we tested the implemented  $\Delta$ -induced strangeness production, which is not based on experimental data, but the corresponding cross sections are based on theoretical calculations from Tsushima *et al.* [15] (see Ref. [9] for more details). In Fig. 8 we compare INCL calculations with and without  $\Delta$ -induced strangeness production to experimental data [19]. This allows to study the impact of  $\Delta N \rightarrow NYK$  secondary reactions compared to  $NN \rightarrow NYK$  primary reactions; the  $\pi N \rightarrow YK$  secondary reactions being naturally suppressed by the phase space constraint. The Bertini model, which has no variance reduction method, is not plotted in Fig. 8 because of the unreasonable computing time needed to get comparable uncertainties. The threshold for the direct production of  $K^+$  with a momentum  $p = 1.280 \pm 0.014$  GeV/c is  $T = 2.115$  GeV in nucleon-nucleon collisions. The sub-threshold production is allowed thanks to effects of structure like the Fermi motion, and also to secondary reactions (e.g.,  $\Delta N$ ) as explained below.

Figure 8 shows that the standard INCL calculations (with  $\Delta$ -induced strangeness production) reproduce relatively well the shape and the absolute values of the experimental data, especially for energies below 2.1 GeV, which correspond to the subthreshold production. However, the standard INCL model overestimates the 2.1–2.9 GeV energy region by around 50%. In comparison, INCL without  $\Delta$ -induced strangeness production underestimates the experimental data over the entire energy range. Above  $T = 2$  GeV, the underestimation is around 20%. This underestimation reaches an order of magnitude for the lowest momenta. This demonstrates the crucial role of  $\Delta$ 's

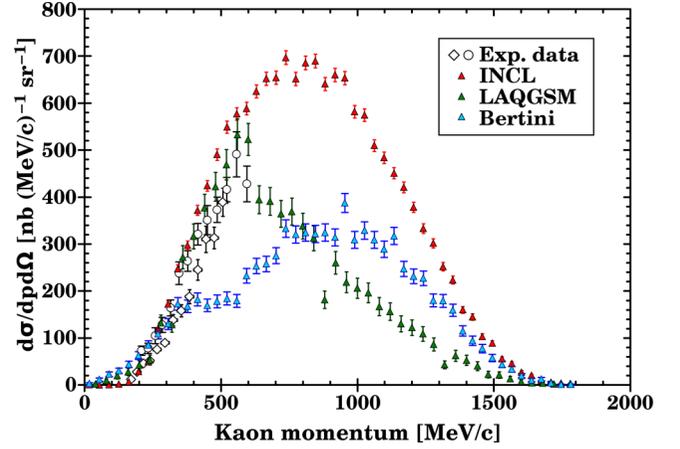


FIG. 9.  $K^+$  momentum spectrum in  $p(2.3 \text{ GeV}) + {}^{12}\text{C}$  collisions within the angular acceptance of the ANKE experiment. Two sets of experimental data [20] (circles and squares) are compared to INCL (red), LAQGSM [21] (green), and the Bertini cascade model [18] (blue). Bias factor used: 10.

in the strangeness production. Going deeper in the analysis, the overestimation in the 2.1–2.9 GeV energy range indicates that the  $\Delta$ -induced kaon production in this region is likely overestimated. This seems consistent with observations made in Ref. [15] where the authors observed an overestimation of the cross sections for these reactions with center-of-mass energies 200 MeV above the threshold.

The assumption that cross sections are very well described near the threshold but are overestimated at higher energy could be due to the neglect of hyperonic resonances in the model used in Ref. [15]. This choice was made because of the low confidence level associated to these resonances. However, they could play a significant role, notably at high energies.

A precise evaluation of  $\Delta$ -induced strangeness production cross sections with INCL is difficult to realize because of the complexity of the spallation process. However, the impossibility to measure the  $\Delta$ -induced strangeness production cross sections and the limitations of the theoretical calculations make simulation models like INCL suitable candidates to estimate and/or test such cross sections.

### C. ANKE

The ANKE experiment [20] investigated the production of  $K^+$  in the forward direction in  $p + A$  collisions with proton kinetic energies between 1 and 2.3 GeV. The targets were  $A = {}^2\text{H}, {}^{12}\text{C}, \text{Cu}, \text{Ag},$  and  ${}^{197}\text{Au}$ . The experiment took place at the COoler SYnchrotron COSY-Jülich in Germany. The angular acceptance was  $\pm 12^\circ$  horizontally and  $\pm 7^\circ$  vertically.

In Fig. 9, calculations from LAQGSM [21], Bertini [18], and INCL are compared to ANKE experimental data. It can be seen that the three models fit relatively well the experimental data at low momenta (below 300 MeV/c) with INCL being slightly closer to the experimental data than the other two. At higher momenta, INCL and LAQGSM still reproduces the data while Bertini underestimates them. At higher energies where no experimental data exist, every model gives a

different shape. Whereas LAQGSM decreases quickly, INCL continues to increase and Bertini predicts a bump. The maxima are around 600 MeV/c, 800 MeV/c, and 900 MeV/c for LAQGSM, INCL, and Bertini, respectively. An extrapolation from Fig. 8 suggests that INCL overestimates the production cross section for  $K^+$  with  $p_{K^+} = 1.280$  GeV/c in forward direction by roughly 30%. This would be compatible with the Bertini's value at  $p_{K^+} = 1.280$  GeV/c. Again, the INCL  $\Delta$ -induced kaon production is probably overestimated.

#### D. LBL

The experiment carried out at the Lawrence Berkeley Laboratory (LBL) [24] studied the inclusive  $K^+$  production using projectiles with  $T = 2.1$  GeV/nucleon. Several projectile types and targets were tested but a direct comparison between measured and modelled data is only possible for the reactions  $p + {}^{208}\text{Pb}$  and  ${}^2\text{H} + {}^{208}\text{Pb}$ .

To estimate the inclusive  $K^+$  production cross section, the collaboration measured  $K^+$  spectra at four different angles:  $\theta = 15^\circ, 35^\circ, 60^\circ,$  and  $80^\circ$  for the  $p + {}^{208}\text{Pb}$  reaction and  $\theta = 15^\circ, 25^\circ, 35^\circ,$  and  $60^\circ$  for the  ${}^2\text{H} + {}^{208}\text{Pb}$  reaction. Spectra were measured for kaon momenta from 350 to 750 MeV/c.

The two panels in Fig. 10 shows that there is good agreement between INCL predictions and experimental data both for proton- and deuteron-induced reactions. It must be mentioned this experiment is very similar to the one carried out by the KaoS collaboration. Though, the angles studied, the energies, the projectiles, and the targets are slightly different but the conclusions are the same. Again, the good agreement not only in the absolute values but also in the energy dependence of the data demonstrates the reliability of the handling of strange particles in the new INCL version. This is a first validation that INCL can handle light clusters as projectile.

#### E. Hades

The High Acceptance DiElectron Spectrometer (HADES) collaboration studied  $\Lambda$  and  $K_s^0$  productions in  $p + p$  and  $p + \text{Nb}$  collisions at 3.5 GeV [23,24]. The Experiment took place at GSI in Germany. The  $\Lambda$  particles were measured in the  $[0.1, 1.3]$  rapidity ( $y$ ) range in the laboratory frame and the  $K_s^0$  particles were measured in the  $[-0.85, 0.55]$  rapidity range in the nucleon-nucleon center of mass. In this second case, the rapidity of the laboratory is  $y_{\text{lab}} = -1.118$ .

The hypernucleus production is strongly correlated to the  $\Lambda$  production since most of the observed hypernuclei involve one or more  $\Lambda$ 's. Therefore, the  $\Lambda$  production must be tested before studying the hypernucleus production, which implies more complex processes.

In Fig. 11 we compare the  $\Lambda$  production yield as function of rapidity calculated using INCL to predictions from UrQMD [25] and GiBUU [26] and to experimental data [23]. The original plot is taken from [23]. The Bertini cascade model does not handle  $\Sigma^0$ 's decay, which plays a role in the  $\Lambda$  production. Therefore, results from the Bertini code are not plotted here.

Figure 11 depicts a rather good agreement between INCL and the experimental data when taking into account

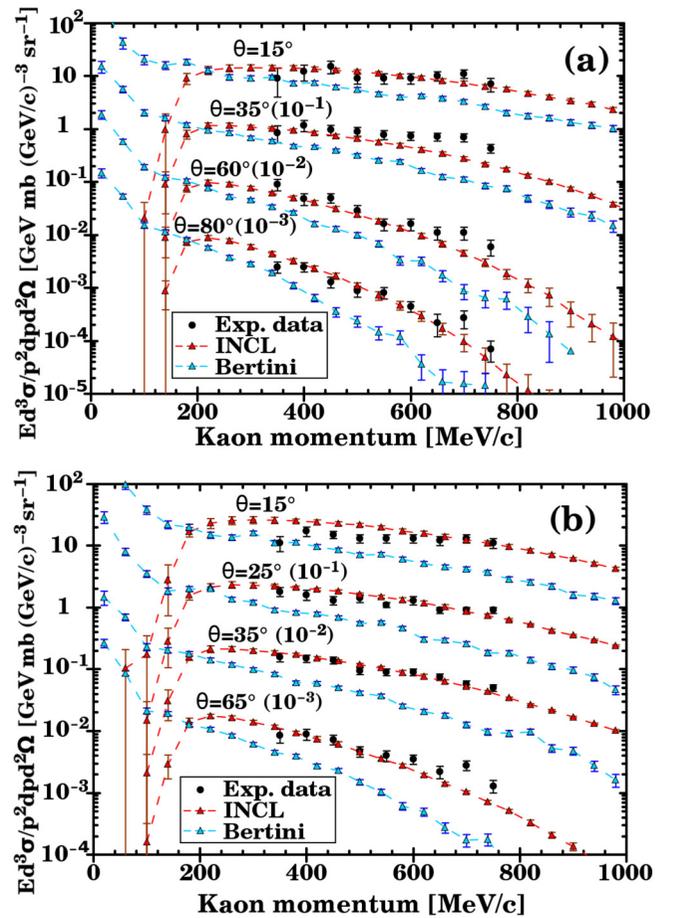


FIG. 10.  $K^+$  invariant cross section at 2.1 GeV/Nucleon for the reactions  $p + {}^{208}\text{Pb}$  (a) and  ${}^2\text{H} + {}^{208}\text{Pb}$  (b). The experimental data [22] (black dots) measured at the LBL are compared to the Bertini cascade model [18] (blue) and INCL (red). Bias factor used: 10.

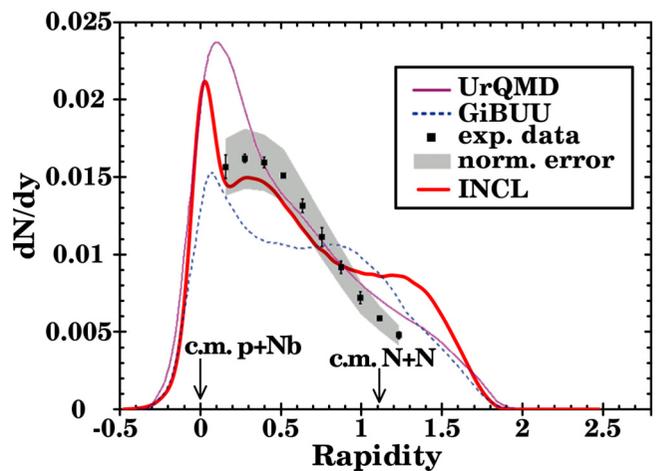


FIG. 11.  $\Lambda$  production yield in  $p(3.5 \text{ GeV}) + \text{Nb}$  collisions as a function of rapidity. The HADES experimental data (black square) are compared to GiBUU (blue dashed line), UrQMD (purple line), and INCL (red line) model predictions. The original plot can be found in [23]. Reprinted by permission from Springer Nature Customer Service Centre GmbH. Bias factor used: 10.

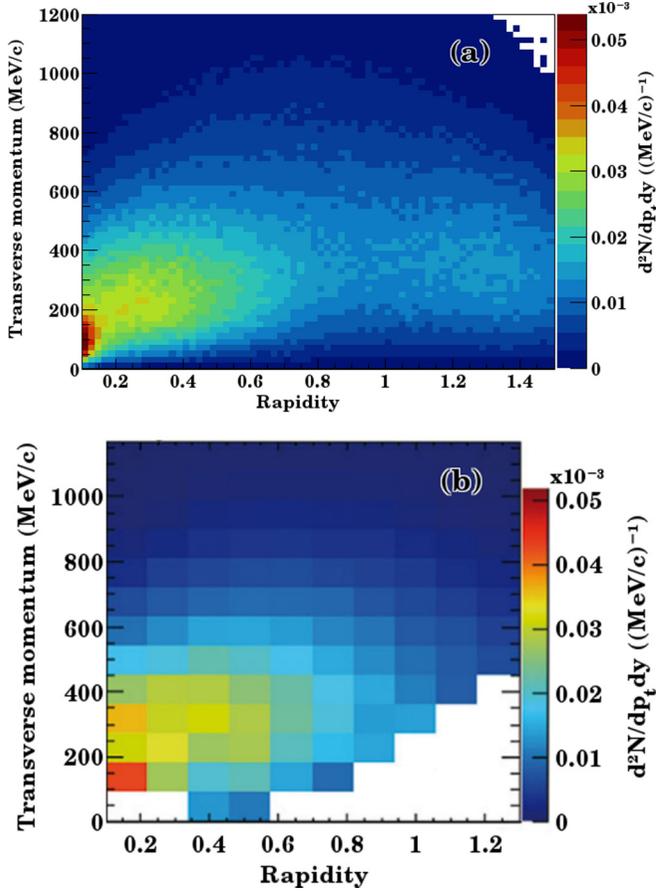


FIG. 12.  $\Lambda$  rapidity versus transverse momentum yield in  $p(3.5 \text{ GeV}) + \text{Nb}$  collisions. (a) INCL predictions, (b) experimental data from Ref. [23] after correction of efficiency. Reprinted by permission from Springer Nature Customer Service Centre GmbH.

systematic errors except for the rapidities higher than 0.9. The bump observed in the INCL predictions around  $y = 1.3$  can also be seen with the GiBUU model, but at lower rapidity. The UrQMD model does not exhibit such a behavior and is close to the experimental data, but it misses the first two data points at the lowest rapidities.

Aiming at understanding the INCL bump, we show in Fig. 12 the transverse momentum versus the rapidity of  $\Lambda$  particles. It can be seen that Fig. 12(a) (INCL) and Fig. 12(b) (HADES data) match well. For the problematic rapidities (above  $y = 0.9$ ) one can find an overabundance of  $\Lambda$  particles of INCL compared to HADES in the rapidity range [1,1.4] and the transverse momentum range [200,500] MeV/c. This corresponds to emission angles between 5 and 19 degrees when the experiment acceptance is between 18 to 85 degrees.

The acceptance of the HADES experiment being limited, data for the unmeasured regions visible in Fig. 12(b) have been estimated by extrapolating the transverse momentum spectra using Maxwell-Boltzmann distributions. The so determined data are in disagreement with the INCL prediction. Consequently, the bump around  $y = 1.3$  predicted by INCL is in a phase-space region not measured experimentally.

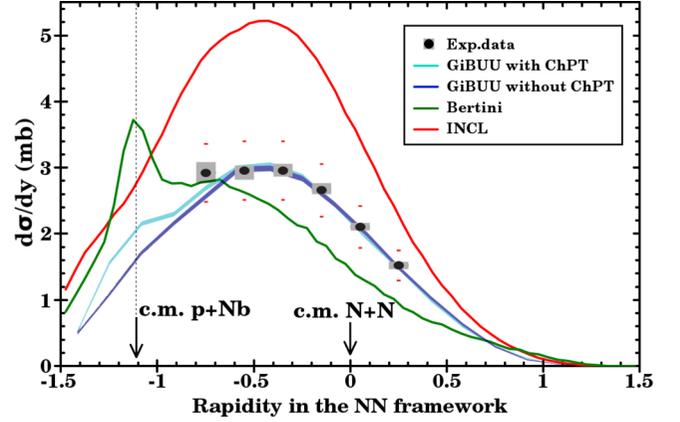


FIG. 13.  $K_s^0$  production cross section in  $p(3.5 \text{ GeV}) + \text{Nb}$  collisions in function of the rapidity in the nucleon-nucleon center of mass. HADES experimental data [24] (black circles) are compared to GiBUU with (cyan) and without (blue) a chiral potential, the Bertini cascade model (green), and INCL (red). The little red dots represent the experimental systematics linked to the normalization. The original plot comes from Ref. [24]. Bias factor used: 10.

Therefore, there is no strict contradiction between experimental data and INCL model predictions

In the region around  $y = 0$ , every model predicts a peak. This peak is high and broad using UrQMD, narrow and smaller using GiBUU, and high and narrow using INCL. A comparison between INCL with and without  $\Sigma^0$  in-flight decay showed the peak in INCL predictions is entirely due to the  $\Lambda$  production inside the nucleus. Additionally, studying the origin of the  $\Lambda$ 's indicated that this peak is the result of the hyperremnant de-excitation for INCL.

The second particle measured by the HADES collaboration is the  $K_s^0$ . In Fig. 13 the experimental data [24] are compared to GiBUU calculations with and without a chiral potential, to the Bertini cascade model, and to INCL. It can be seen that the best description of the experimental data are obtained using GiBUU. However, the version of GiBUU used here is a modified version in which the  $K^0$  production has been artificially reduced to fit the  $p + p$  HADES experimental data (see Ref. [24] for details). Additionally, this GiBUU version is not the same as the one used in Ref. [23], although the reaction studied is the same. This makes their results difficult to be interpreted. INCL gives a good shape but overestimates the experimental data by roughly 65%. Similarly to the case for the  $K^+$  production, the  $\Delta$ -induced reactions could be an explanation for the overestimation but some part of this overestimation might also be due to the normalization. The total reaction cross section  $\sigma_{\text{tot}}^{p+Nb}$  used by HADES is  $848 \pm 127 \text{ mb}$ , while INCL calculates a value of 1048 mb. A measurement of the total reaction cross section for the same system at a lower energy (1.2 GeV instead of 3.5 GeV) gave  $1063 \pm 40 \text{ mb}$  [27], which is consistent with the INCL value. The Bertini cascade model does not reproduce the energy dependence of the experimental data, but the predicted absolute yield of  $K_s^0$  corresponds to the experimental data. The bump around the laboratory rapidity ( $y_{\text{lab}} = -1.118$ ) can also be seen in GiBUU predictions with the chiral potential. This

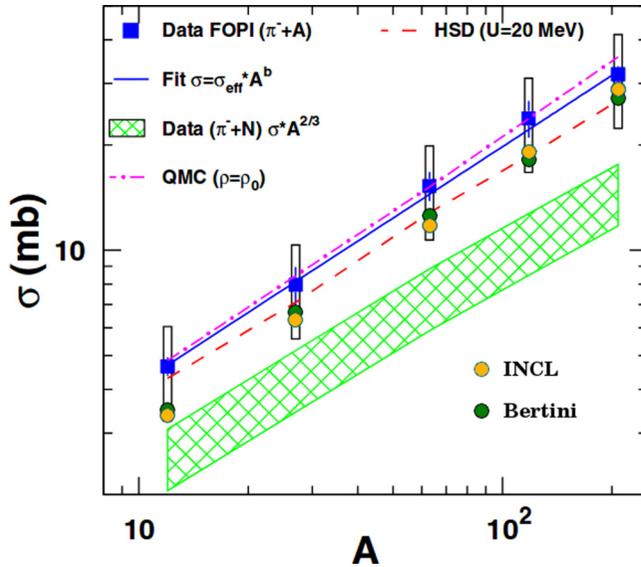


FIG. 14. Neutral kaon production cross section in the collision  $\pi^-$  (1020 MeV) +  $A$  ( $p_\pi = 1150$  MeV/ $c$ ). The FOPI experimental data [28] (blue squares) are compared to QMC [29] (pink dash-dotted line), HSD [30,31] (red dashed line), Bertini [18] (green circles), and INCL (orange circles) models. Open rectangular bars are the FOPI systematic uncertainties. The meaning of the hatched area is explained in the text. The original figure comes from Ref. [28]. Bias factor used: 10.

is likely due to the attractive potential used by these models for the  $K^0$ , whereas INCL uses a repulsive potential.

### F. FOPI

The FOPI collaboration [28] measured the in-medium neutral kaon inclusive cross sections in  $\pi^-$ -induced reactions on various targets: C, Al, Cu, Sn, and Pb. The pion beam had a kinetic energy  $T \approx 1.02$  GeV ( $p_\pi = 1150$  MeV/ $c$ ). The experiment took place at the heavy-ion synchrotron SIS at GSI. The geometrical acceptance of the detector was restricted to the polar angles in the range  $25^\circ < \theta < 150^\circ$ . An *a posteriori* correction was applied by the FOPI collaboration to obtain the total inclusive cross sections.

Figure 14 depicts the production of neutral kaons measured by the FOPI collaboration. In the original paper [28], the authors compared their experimental data with two models: the quark-meson coupling model (QMC) [29] and the hadron-string-dynamics model (HSD) [30,31]. By fitting the data, the FOPI collaboration found that the inclusive cross sections can be described by a  $A$ -dependent cross section of the form  $\sigma(A) = \sigma_{\text{eff}} \times A^b$ , with  $b = 0.67 \pm 0.03$ . This would indicate that  $K^0$  production is dominated by peripheral collision. Therefore, they developed the function:

$$\sigma(A) = \sigma(\pi^-(1150 \text{ MeV}/c) + N \rightarrow K^0 + X) \times A^{2/3}, \quad (13)$$

with  $N$  a target nucleon. The nucleon cross section was obtained by summing all  $\sigma(\pi^- + N \rightarrow K^0 + Y)$  processes weighted with the relative proton and neutron numbers of the target nucleus. The so developed function is represented as

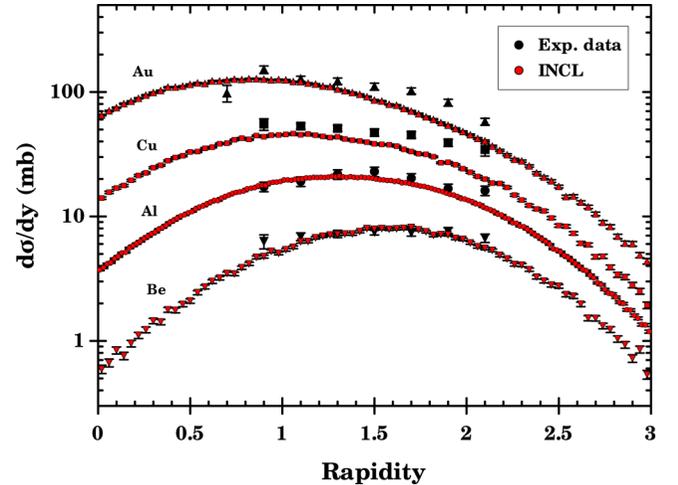


FIG. 15.  $K^+$  production cross sections in  $p(14.6 \text{ GeV}/c) + A$  collisions as a function of rapidity. The experimental data [32] (black) are compared to INCL predictions (red). No variance reduction method was used.

a hatched band in Fig. 14. The bandwidth corresponds to an uncertainty of 20%. It appears that the experimental data do not enter in the region defined by this band, which indicates that the simple dependence expressed in Equation 13 cannot explain the observed cross sections. Therefore, there must be an additional process.

The INCL calculations slightly underestimate the experimental data but they are within the experimental systematic uncertainties (open rectangular bars). The inclusive cross sections predicted by INCL is proportional to  $A^{3/4}$ , which promote a different interpretation on the  $K^0$  production processes with larger contributions of secondary reactions and a strangeness production deeper in the nucleus. The Bertini cascade model [18] shows a result similar to INCL but with a slightly smaller slope. Both INCL and Bertini cross sections are consistent with Equation 13 using a  $A^{3/4}$  dependence instead of the  $A^{2/3}$  dependence proposed.

### G. E-802

The INCL model has been extended to work up to incident energies of about  $\approx 15$  GeV. However, the experiment studying strangeness production focus below 3.5 GeV. A notable exception is the experiment realized by the E-802 collaboration [32] at the Brookhaven National Laboratory (U.S.) using a proton beam momentum of 14.6 GeV/ $c$  ( $T_p \approx 13.7$  GeV). This experiment measured various particle production cross sections in proton-nucleus reactions and, notably, is also measured the  $K^+$  production cross sections. Therefore, it provides a good opportunity to test INCL strangeness physics for the highest energies. The nuclei studied by the E-802 collaboration were Be, Al, Cu, and Au.

Figure 15 shows a comparison of INCL predictions with E-802 experimental data. The Bertini model predictions are not plotted because the upper energy limit for this model is about 10 GeV. Considering that the experimental data are close to the upper limit of INCL, the shape as well as the

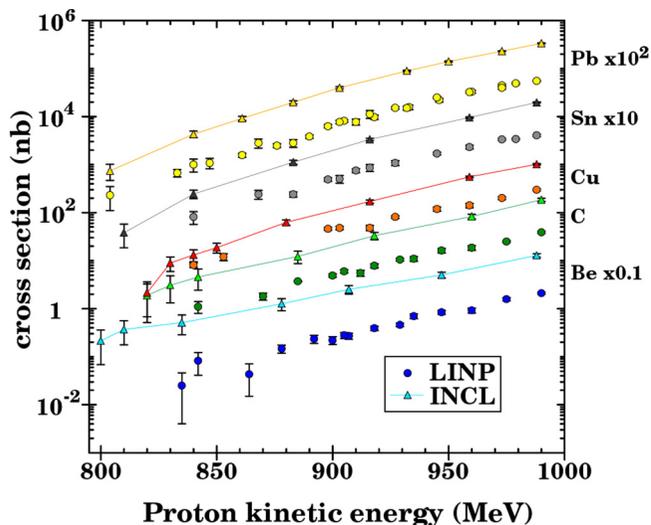


FIG. 16. Subthreshold  $K^+$  production cross sections in  $p+A$  collisions as a function of the initial proton kinetic energy. The experimental data [33] (circles) are compared to INCL predictions (triangles).

absolute values are in excellent agreement. However, a slight underestimation at high rapidities can be observed for Cu and Au. The observed deviations are difficult to analyze because they are only for two target elements and the discrepancies differ: a bump for Au and a linear deviation for Cu.

#### H. LINC

The experiment carried out at the Leningrad Institute of Nuclear Physics (LINC) [33] studied the subthreshold  $K^+$  production on various nuclei, from beryllium to uranium. A measurement of the total cross sections as a function of energy was performed in the 800–1000 MeV energy range for Be, C, Cu, Sn, and Pb.

In Fig. 16 INCL predictions are compared with the experimental data. The first observation is a clear overestimation of the experimental data. This overestimation is around a factor 6 for Be and about a factor 4 for the other nuclei. The prediction of the slope is good, especially for C, Sn, and Pb. While the agreement between model predictions and experimental data is only fair, the major success of these calculations is the variance reduction (see Sec. III). The effective bias factors of these calculations are around 1000 depending on the configuration with a maximum at 2000. This allowed to obtain cross sections below the nanobarn scale using a relatively short calculation time (half a day for the entire set using parallelization). Such a result cannot be obtained in a reasonable time without a variance reduction method.

Different explanations are possible to explain the overestimation. To understand the problem, it is important to remember that the beam energies in this experiment are far below the  $K^+$  production threshold for  $p+p$  collisions (1.582 GeV). Therefore, the strangeness production requires strong effects of structure, which are complex and not very well understood. Such effects include effects related to the momentum spectra of particles inside the nucleus and effects related to successive

collisions. Additionally,  $\Delta$  particles seem to play a crucial role in this experiment. An explanation for the overestimation could therefore be the treatment of the  $\Delta$  particles in INCL, especially for the production of strange particles in  $\Delta N$  reactions. Another explanation could be the semiclassical description of the nucleus by INCL, which does not consider every quantum effect, which could play a significant role in subthreshold processes. A following study has been carried out to better understand the discrepancies. The effect and limitations of the semiclassical approach cannot be tested but modifications related to the physics of  $\Delta$  particles or to effects of structure showed a sensible reduction of the strangeness production. None of the modifications, however, could fully explain the discrepancies. Therefore, the overestimation is likely due to a combination of the three explanations proposed above. However, it is difficult to prove or reject this hypothesis considering the complexity of the processes involved.

#### V. CONCLUSION AND FUTURE

In this paper, a full description and a detailed analysis of the new version of INCL, INCL++6, were presented. First, the main new capabilities of INCL, which are the strangeness production and the variance reduction methods, were described. Then, the model predictions were compared to experimental data.

In this second part, different projectiles (proton, deuteron, and pion) were tested at various kinetic energies: from 1.6 to 13.7 GeV for protons, 1.02 GeV for pions, and 2.1 GeV/nucleon for deuteron. Various targets ranging from  $^9\text{Be}$  to  $^{208}\text{Pb}$  were considered and a wide range of angles was covered.

For most of the studied cases, there is a good agreement between experimental data and INCL predictions. Notably, the dependence of experimental data on either energy or target mass number is often well reproduced, which demonstrates that the strangeness physics is well implemented. However, slight discrepancies exist. In some cases experimental data are overestimated. Different explanations were proposed to understand these discrepancies, but the lack of experimental data prevents definitive conclusions. The dominant problem is likely in the  $\Delta$ -induced reactions. Though, implementing  $\Delta$ -induced reactions significantly improved the description of strangeness production, the cross sections used are suspected to be overestimated at high energies, which is consistent with the conclusions in Ref. [15].

With the validation of INCL++6, the problem of the statistics and computational time arose. This problem has been solved with the implementation of a variance reduction scheme in INCL, which has been described in detail. The developed method increases artificially the strangeness production respecting some mathematical constraints, which results in a better statistics for observables linked to the strangeness physics, and therefore, to a reduction of the required calculation time.

In addition, our new version of INCL is now implemented in the transport code Geant4 [34]. Thus, it can be used in the simulation of macroscopic systems, which allows other collaborations to have a full access to the last version of our

model. They can use it for the design of new experiments dedicated to the study of strange particles and hypernuclei in the foreseeable future, such as for the HypHI [35], Panda [36], and CBM [37] experiments at the FAIR facility.

In the future, a more detailed study of  $\Delta$ -induced strangeness production should be carried out to improve the corresponding cross-section parametrizations. The  $K^-$  production could be improved by the implementation of additional strangeness exchange channels. Concerning the variance reduction methods, the safeguard can be improved to reduce the number of edge cases exhibiting convergence issues, what would simplify its use.

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### APPENDIX: VERTEX CROSS-SECTION RATIO

In this Appendix some examples of cross-section ratio calculations are discussed. This plans to explain the biasing steps of the variance reduction scheme.

#### 1. Example 1

In the first example (Fig. 17) we consider two successive binary collisions. The two nonstrange particles at the origin of the vertex  $A$  are considered without any history. Let us assume that the total cross section for the vertex  $A$  is 20 mb, with the reaction cross section for the production of strangeness is 0.1 mb and therefore the nonstrangeness production cross section is 19.9 mb. Our goal is an increase of the total strangeness production by a factor 10 (= bias factor).

According to Eq. (2), to increase the probability of a particle production by a factor 10 is to give it an importance  $W = 1/10$ .

Since initial particles in  $A$  have no history, the probability of strangeness production at this time of the cascade is directly proportional to the strangeness production cross section.

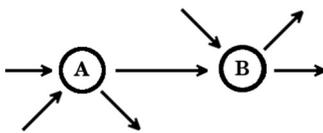


FIG. 17. Example of basic intranuclear cascade represented as a time ordered graph. Circles and arrows represent binary collisions and the particles propagating, respectively.

Consequently, to increase the probability of strangeness production by a factor 10, the corresponding cross section should be multiplied by the same factor 10. Therefore, the modified strangeness production cross section is 1 mb and, because the total cross section should be conserved, the modified nonstrangeness production cross section is 19 mb.

At this stage of the variance reduction scheme, the cascade is biased. Now, the vertex cross-section ratio of the vertex  $A$  should be determined. The vertex cross-section ratio is the ratio of the event importance before versus after this vertex. This ratio is directly equal to the inverse of the cross-section multiplying factor of the chosen channel. This means the vertex cross-section ratio will be equal, in this case, to  $0.1 \text{ mb}/1 \text{ mb} = 0.1$  if strangeness is produced and equal to  $19.9 \text{ mb}/19 \text{ mb} \approx 1.047$  otherwise. Let us now assume that no strange particle is produced. At this stage, the probability of realization for this cascade is decreased by a factor 1.047 and the importances of outgoing particles are increased by 1.047. In INCL, the cross-section ratio is registered for future uses.

Let us assume that one of the particles produced in the vertex  $A$  collides with another nonstrange particle without history in the vertex  $B$ . Remember that we want to increase the probability of creating strange particles by a factor 10. Therefore, the importance of a strange particle produced in the vertex  $B$  should be  $W = 0.1$ . The importance is the product of every cross-section ratio on a path:

$$0.1 = W = \text{CSR}(A) \times \text{CSR}(B). \quad (\text{A1})$$

Therefore, if a strange particle is produced, then the cross-section ratio of vertex  $B$  should be  $\text{CSR}(B)_S = 1/[10 \times \text{CSR}(A)] \approx 0.095$ . This determines the modified strangeness production cross section of this vertex ( $\sigma'_S = \sigma_S/\text{CSR}(B)_S$ ), the modified nonstrangeness production cross section ( $\sigma'_{\text{NS}} = \sigma_{\text{tot}} - \sigma'_S$ ), and the cross-section ratio of vertex  $B$  in case of nonstrangeness production [ $\text{CSR}(B)_{\text{NS}} = \sigma_{\text{NS}}/\sigma'_{\text{NS}}$ ].

In the general case, the cross-section ratio of a vertex following a set  $K$  of vertices and with strange particles in its final state should be

$$\text{CSR}(X) = \frac{1}{\text{bias ratio} \times \prod_{I \in K} \text{CSR}(I)}. \quad (\text{A2})$$

#### 2. Example 2

The second example shown in Fig. 18 illustrates the case of branch recombination. Continuing the case presented in Appendix A 1, a particle coming from vertex  $B$ , which is itself induced by a particle from vertex  $A$ , collides with a particle directly produced by vertex  $A$ .

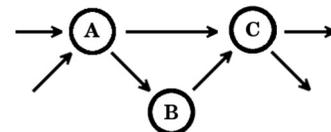


FIG. 18. Same as described in the caption of Fig. 17, except with three vertices and a branch recombination.

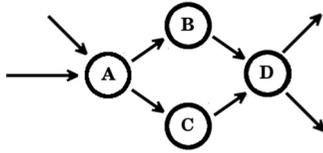


FIG. 19. Same as described in the caption of Fig. 18, except with four vertices.

Equation (A2) gives  $CSR(C)$  in case of strangeness production:

$$CSR(C) = \frac{1}{\text{bias ratio} \times CSR(A) \times CSR(B)}. \quad (\text{A3})$$

We also know  $CSR(B)$ , which produced strangeness:

$$CSR(B) = \frac{1}{\text{bias ratio} \times CSR(A)}. \quad (\text{A4})$$

Therefore,  $CSR(C) = 1$ . Consequently, the cross sections are not modified because the probability to reach the initial

state of vertex  $C$  has already been increased by a factor equal to the bias ratio.

### 3. Example 3

The last example shown in Fig. 19 is a more complex case of branch recombination. Let us assume that no strange particles were produced in the vertices  $A$  and  $B$  but that strangeness production happened in vertex  $C$ .  $CSR(A)$  and  $CSR(B)$  are determined as described in Appendix A1 in the case of nonstrangeness production.  $CSR(C)$  follows the Eq. (A2) because of strangeness production:

$$CSR(C) = \frac{1}{\text{bias ratio} \times CSR(A)}. \quad (\text{A5})$$

If strangeness is produced in vertex  $D$ , then  $CSR(D)$  also follows Eq. (A2):

$$\begin{aligned} CSR(D) &= \frac{1}{\text{bias ratio} \times CSR(A) \times CSR(B) \times CSR(C)} \\ &= \frac{1}{CSR(B)}. \end{aligned} \quad (\text{A6})$$

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