

Fermi Constant from Muon Decay Versus Electroweak Fits and Cabibbo-Kobayashi-Maskawa Unitarity

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The Fermi constant G_F is extremely well measured through the muon lifetime, defining one of the key fundamental parameters in the standard model (SM). Therefore, to search for physics beyond the SM (BSM) via G_F , the constraining power is determined by the precision of the second-best independent determination of G_F . The best alternative extractions of G_F proceed either via the global electroweak (EW) fit or from superallowed β decays in combination with the Cabibbo angle measured in kaon, τ , or D decays. Both variants display some tension with G_F from muon decay, albeit in opposite directions, reflecting the known tensions within the EW fit and hints for the apparent violation of Cabibbo-Kobayashi-Maskawa unitarity, respectively. We investigate how BSM physics could bring the three determinations of G_F into agreement using SM effective field theory and comment on future perspectives.

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Introduction.—The numerical value of the Fermi constant G_F is conventionally defined via the muon lifetime within the standard model (SM). Even though this measurement is extremely precise [1–3] at

$$G_F^\mu = 1.1663787(6) \times 10^{-5} \text{ GeV}^{-2}, \quad (1)$$

at the level of 0.5 ppm, its determination of the Fermi constant is not necessarily free of beyond the SM (BSM) contributions. In fact, one can only conclude the presence or absence of BSM effects by comparing G_F^μ to another independent determination. This idea was first introduced by Marciano in Ref. [4], concentrating on Z-pole observables and the fine-structure constant α . In addition to a lot of new data that have become available since 1999, another option already mentioned in Refs. [4,5]—the determination of G_F from β and kaon decays using Cabibbo-Kobayashi-Maskawa (CKM) unitarity—has become of particular interest due to recent hints for the (apparent) violation of first-row CKM unitarity. These developments motivate a fresh look at the Fermi constant: in particular, on its extraction from a global electroweak (EW) fit and via CKM unitarity, as will be discussed in the first part of this Letter.

The comparison of the resulting values for G_F then serves as a model-independent measure of possible BSM effects. It shows that with modern input, the two independent extractions are close in precision yet still lagging behind muon decay by almost three orders of magnitude. Therefore, the BSM sensitivity is governed by the uncertainty of these two indirect determinations. Since the different G_F measurements turn out to display some disagreement beyond their quoted uncertainties, the second part of this Letter is devoted to a systematic analysis of possible BSM contributions in SM effective field theory (SMEFT) [6,7] to see which scenarios could account for these tensions without being excluded by other constraints. This is important to identify BSM scenarios that could be responsible for the tensions, which will be scrutinized with forthcoming data in the next years.

Determinations of G_F .—Within the SM, the Fermi constant G_F is defined by, and is most precisely determined from, the muon lifetime [2]

$$\frac{1}{\tau_\mu} = \frac{(G_F^\mu)^2 m_\mu^5}{192\pi^3} (1 + \Delta q), \quad (2)$$

where Δq includes the phase space, QED, and hadronic radiative corrections. The resulting numerical value in Eq. (1) is so precise that its error can be ignored in the following. To address the question whether G_F^μ subsumes BSM contributions, however, alternative independent determinations of G_F are indispensable; and their precision limits the extent to which BSM contamination in G_F^μ can be detected.

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In Ref. [4], the two best independent determinations were found as

$$G_F^{Z\ell^+\ell^-} = 1.1650(14) \binom{+11}{-6} \times 10^{-5} \text{ GeV}^{-2},$$

$$G_F^{(3)} = 1.1672(8) \binom{+18}{-7} \times 10^{-5} \text{ GeV}^{-2}, \quad (3)$$

where the first variant uses the width for $Z \rightarrow \ell^+\ell^-(\gamma)$, while the second employs α and $\sin^2\theta_W$ together with the appropriate radiative corrections. Since the present uncertainty in $\Gamma[Z \rightarrow \ell^+\ell^-(\gamma)] = 83.984(86) \text{ MeV}$ [8] is only marginally improved compared to the one available in 1999, the update

$$G_F^{Z\ell^+\ell^-} = 1.1661(16) \times 10^{-5} \text{ GeV}^{-2} \quad (4)$$

does not lead to a gain in precision, but the shift in the central value improves agreement with G_F^μ . The second variant, $G_F^{(3)}$, is more interesting because, here, the main limitation arose from the radiative corrections, which have seen significant improvements regarding the input values for the masses of the top quark m_t , the Higgs boson M_H , the strong coupling α_s , and the hadronic running of α . In fact, with all EW parameters determined, it now makes sense to use the global EW fit (for which G_F^μ is usually a key input quantity) instead as a tool to determine G_F in a completely independent way.

The EW observables included in our fit (W mass, $\sin^2\theta_W$, and Z -pole observables [9,10]) are given in Table I, with the other input parameters summarized in Table II. Here, the hadronic running $\Delta\alpha_{\text{had}}$ is taken from e^+e^- data, which are insensitive to the changes in $e^+e^- \rightarrow$ hadrons cross sections [11–18] recently suggested by lattice QCD [19], as long as these changes are concentrated at low energies [20–23]. We perform the global EW fit in a Bayesian framework using the publicly available HEPfit package [24], whose Markov chain Monte Carlo determination of posteriors is powered by the Bayesian Analysis Toolkit [25]. As a result, we find

$$G_F^{\text{EW}}|_{\text{full}} = 1.16716(39) \times 10^{-5} \text{ GeV}^{-2}, \quad (5)$$

which is a gain in precision over $G_F^{(3)}$ in Eq. (3) by a factor of 5. As depicted in Fig. 1, this value lies above G_F^μ by 2σ , reflecting the known tensions within the EW fit [26,27]. For comparison, we also considered a closer analog of $G_F^{(3)}$ by only keeping $\sin^2\theta_W$ from Table I in the fit, which gives

$$G_F^{\text{EW}}|_{\text{minimal}} = 1.16728(83) \times 10^{-5} \text{ GeV}^{-2}, \quad (6)$$

which is consistent with Eq. (5) but with a larger uncertainty. The pull of G_F^{EW} away from G_F^μ is mainly driven by M_W , $\sin^2\theta_W$ from hadron colliders, A_ℓ , and $A_{\text{FB}}^{0,\ell}$.

As a third possibility, one can determine the Fermi constant from superallowed β decays, taking V_{us} from kaon or τ decays and assuming CKM unitarity ($|V_{ub}|$ is also

TABLE I. EW observables included in our global fit together with their current experimental values.

M_W (GeV)	[8]	80.379(12)
Γ_W (GeV)	[8]	2.085(42)
BR($W \rightarrow$ had)	[8]	0.6741(27)
BR($W \rightarrow$ lep)	[8]	0.1086(9)
$\sin^2\theta_{\text{eff}}(Q_{\text{FB}})$	[8]	0.2324(12)
$\sin^2\theta_{\text{eff}}(\text{Tevatron})$	[28]	0.23148(33)
$\sin^2\theta_{\text{eff}}(\text{LHC})$	[29–32]	0.23129(33)
Γ_Z (GeV)	[10]	2.4952(23)
σ_h^0 (nb)	[10]	41.541(37)
P_τ^{pol}	[10]	0.1465(33)
A_ℓ	[10]	0.1513(21)
R_ℓ^0	[10]	20.767(25)
$A_{\text{FB}}^{0,\ell}$	[10]	0.0171(10)
R_b^0	[10]	0.21629(66)
R_c^0	[10]	0.1721(30)
$A_{\text{FB}}^{0,b}$	[10]	0.0992(16)
$A_{\text{FB}}^{0,c}$	[10]	0.0707(35)
A_b	[10]	0.923(20)
A_c	[10]	0.670(27)

needed, but the impact of its uncertainty is negligible). This is particularly interesting, given recent hints for the apparent violation of first-row CKM unitarity known as the Cabibbo angle anomaly (CAA). The significance of the tension crucially depends on the radiative corrections applied to β decays [45–52] but also on the treatment of tensions between $K_{\ell 2}$ and $K_{\ell 3}$ decays [53] and the constraints from τ decays [54]; see Ref. [55] for more details. In the end, quoting a significance around 3σ should give a realistic representation of the current situation; for definiteness, we will thus use the estimate of first-row CKM unitarity from Ref. [8]:

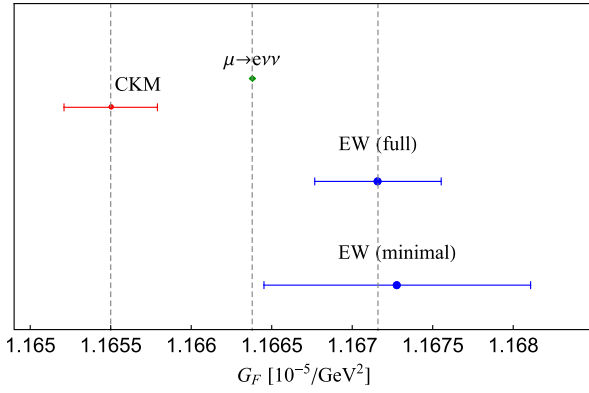
$$|V_{ud}|^2 + |V_{us}|^2 + |V_{ub}|^2 = 0.9985(5). \quad (7)$$

In addition, we remark that there is also a deficit in the first-column CKM unitarity relation [8]

$$|V_{ud}|^2 + |V_{cd}|^2 + |V_{td}|^2 = 0.9970(18), \quad (8)$$

TABLE II. Parameters of the EW fit together with their (Gaussian) priors.

Parameter	Prior
$\alpha \times 10^3$ [8]	7.2973525664(17)
$\Delta\alpha_{\text{had}} \times 10^4$ [16,17]	276.1(1.1)
$\alpha_s(M_Z)$ [8,33]	0.1179(10)
M_Z (GeV) [8,34–37]	91.1876(21)
M_H (GeV) [8,38–40]	125.10(14)
m_t (GeV) [8,41–44]	172.76(30)


 FIG. 1. Values of G_F from the different determinations.

which is less significant than Eq. (7) but suggests that if the deficits were due to BSM effects, they would likely be related to β decays. For the numerical analysis, we will continue to use Eq. (7), given the higher precision. The deficit in Eq. (7) translates to

$$G_F^{\text{CKM}} = 0.99925(25) \times G_F^\mu = 1.16550(29) \times 10^{-5} \text{ GeV}^{-2}. \quad (9)$$

Comparing the three independent determinations of G_F in Fig. 1, one finds the situation in which G_F^{EW} lies above G_F^μ by 2σ , G_F^{CKM} lies below G_F^μ by 3σ , and the tension between G_F^{EW} and G_F^{CKM} amounts to 3.4σ . To bring all three determinations into agreement within 1σ , an effect in at least two of the underlying processes is thus necessary. This leads us to study BSM contributions to (1) $\mu \rightarrow e\nu\nu$ transitions; (2) $Z \rightarrow \ell\ell, \nu\nu$, α_2/α , and M_Z/M_W ; and (3) superallowed β decays, where the second point gives the main observables in the EW fit, with α_2/α as a proxy for the ratio of $SU(2)_L$ and $U(1)_Y$ couplings. We do not consider the possibility of BSM effects in kaon, τ , or D decays because this would require a correlated effect with a relating symmetry. Furthermore, as shown in Ref. [55], the sensitivity to a BSM effect in superallowed β decays is enhanced by a factor of $|V_{ud}|^2/|V_{us}|^2$ compared to kaon, τ , or D decays. This can also be seen from Eq. (7) because $|V_{ud}|$ gives the dominant contribution.

BSM explanations of the discrepancies between these determinations of G_F have been studied in the literature in the context of the CAA [55–66]. In this Letter, we will analyze possible BSM effects in all three G_F determinations using an effective-field-theory approach with gauge-invariant dimension-6 operators [6,7].

SMEFT analysis.—Dimension-6 operators that can explain the differences among the determinations of G_F can be grouped into the following classes: (A) four-fermion operators in $\mu \rightarrow e\nu\nu$, (B) four-fermion operators in $u \rightarrow d e \nu$, (C) modified W - u - d couplings, (D) modified W - ℓ - ν couplings, and (E) other operators affecting the EW fit.

Global fits to a similar set of effective operators have been considered in Refs. [67–72]; here, we will concentrate directly on the impact on G_F determinations, following the conventions of Ref. [7].

Four-fermion operators in $\mu \rightarrow e\nu\nu$.—Not counting flavor indices, there are only two operators that can generate a neutral current involving four leptons:

$$Q_{\ell\ell}^{ijkl} = \bar{\ell}_i \gamma^\mu \ell_j \bar{\ell}_k \gamma^\mu \ell_l, \quad Q_{\ell e}^{ijkl} = \bar{\ell}_i \gamma^\mu \ell_j \bar{e}_k \gamma^\mu e_l. \quad (10)$$

Not all flavor combinations are independent, e.g., $Q_{\ell\ell}^{ijkl} = Q_{\ell\ell}^{klij} = Q_{\ell\ell}^{iljk} = Q_{\ell\ell}^{jilk}$ due to Fierz identities and $Q_{\ell\ell}^{ijkl} = Q_{\ell\ell(e)}^{ijkl*}$ due to Hermiticity. Instead of summing over flavor indices, it is easiest to absorb these terms into a redefinition of the operators whose latter two indices are 12, which contribute directly to $\mu \rightarrow e\nu\nu$. Therefore, we have to consider nine different flavor combinations for both operators: (1) $Q_{\ell\ell}^{2112}$ contributes to the SM amplitude (its coefficient is real by Fierz identities and Hermiticity). Therefore, it can give a constructive or destructive effect in the muon lifetime and does not affect the Michel parameters [73–80]. In order to bring G_F^{CKM} and G_F^μ into agreement at 1σ , we need

$$C_{\ell\ell}^{2112} \approx -1.4 \times 10^{-3} G_F. \quad (11)$$

This Wilson coefficient is constrained by LEP searches for $e^+e^- \rightarrow \mu^+\mu^-$ [9]:

$$-\frac{4\pi}{(9.8 \text{ TeV})^2} < C_{\ell\ell}^{1221} < \frac{4\pi}{(12.2 \text{ TeV})^2}, \quad (12)$$

which is a factor of 8 weaker than preferred by the CAA but within reach of future e^+e^- colliders. (2) Even though $Q_{\ell e}^{2112}$ has a vectorial Dirac structure, it leads to a scalar amplitude after applying Fierz identities. Its interference with the SM amplitude is usually expressed in terms of the Michel parameter $\eta = \text{Re } C_{\ell e}^{2112}/(2\sqrt{2}G_F)$, leading to a correction of $1 - 2\eta m_e/m_\mu$. In the absence of right-handed neutrinos, the restricted analysis from Ref. [78] applies, constraining the shift in G_F^μ to 0.68×10^{-4} , which is well below the required effect to obtain 1σ agreement with G_F^{CKM} or G_F^{EW} . (3) The operators $Q_{\ell\ell(e)}^{1212}$ could contribute to muon decay as long as the neutrino flavors are not detected. To reconcile G_F^{CKM} and G_F^μ within 1σ , we need $|C_{\ell\ell}^{1212}| \approx 0.045 G_F$ or $|C_{\ell e}^{1212}| \approx 0.09 G_F$. Both solutions are excluded by muonium-antimuonium oscillations ($M = \mu^+e^-$) [81]

$$\mathcal{P}(\bar{M}-M) < \frac{8.3 \times 10^{-11}}{S_B}, \quad (13)$$

with correction factors of $S_B = 0.35$ ($C_{\ell\ell}^{1212}$) and $S_B = 0.78$ ($C_{\ell e}^{1212}$) for the extrapolation to zero magnetic field. By comparing to the prediction for the rate [82–84],

$$\begin{aligned} \mathcal{P}(\bar{M}-M) &= \frac{8(\alpha\mu_{\mu e})^6\tau_\mu^2 G_F^2}{\pi^2} |C_{\ell\ell(e)}^{1212}/G_F|^2 \\ &= 3.21 \times 10^{-6} |C_{\ell\ell(e)}^{1212}/G_F|^2, \end{aligned} \quad (14)$$

with reduced mass $\mu_{\mu e} = m_\mu m_e / (m_\mu + m_e)$, the limits become $|C_{\ell\ell(e)}^{1212}| < 8.6(5.8) \times 10^{-3} G_F$. (4) For $Q_{\ell\ell(e)}^{1112}$, again, numerical values of $|C_{\ell\ell(e)}^{1112}| \approx 0.09 G_F$ are preferred (as for all the remaining Wilson coefficients in this list). Both operators give tree-level effects in $\mu \rightarrow 3e$, e.g.,

$$\text{Br}[\mu \rightarrow 3e] = \frac{m_\mu^5 \tau_\mu}{768\pi^3} |C_{\ell\ell}^{1112}|^2 = 0.25 \left| \frac{C_{\ell\ell}^{1112}}{G_F} \right|^2, \quad (15)$$

which exceeds the experimental limit on the branching ratio of 1.0×10^{-12} [85] by many orders of magnitude (the result for $C_{\ell e}^{1112}$ is smaller by a factor of 1/2). (5) The operators $Q_{\ell\ell(e)}^{2212}$ and $Q_{\ell\ell(e)}^{3312}$ contribute at the one-loop level to $\mu \rightarrow e$ conversion and $\mu \rightarrow 3e$, as well as at the two-loop level to $\mu \rightarrow e\gamma$ [86]. Here, the current best bounds come from $\mu \rightarrow e$ conversion. Using table 3 in Ref. [86], we have

$$\begin{aligned} |C_{\ell\ell}^{3312}| &< 6.4 \times 10^{-5} G_F, \\ |C_{\ell\ell}^{2212}| &< 2.8 \times 10^{-5} G_F, \end{aligned} \quad (16)$$

excluding again a sizable BSM effect, and this is similar for $Q_{\ell e}^{3312}$ and $Q_{\ell e}^{2212}$. (6) $Q_{\ell\ell(e)}^{2312}$, $Q_{\ell\ell(e)}^{3212}$ and $Q_{\ell\ell(e)}^{1312}$, $Q_{\ell\ell(e)}^{3112}$ contribute to $\tau \rightarrow \mu\mu e$ and $\tau \rightarrow \mu ee$, respectively, which excludes a sizable effect in analogy to $\mu \rightarrow 3e$ above [54,87,88].

The other four-quark operators can only contribute via loop effects, which leads us to conclude that the only viable mechanism proceeds via a modification of the SM operator $Q_{\ell\ell}^{2112}$.

Four-fermion operators in $d \rightarrow ue\nu$.—First of all, the operators $Q_{\ell equ}^{(1)1111}$ and $Q_{\ell equ}^{(3)1111}$ give rise to $d \rightarrow ue\nu$ scalar amplitudes. Such amplitudes lead to enhanced effects in $\pi \rightarrow \mu\nu/\pi \rightarrow e\nu$ with respect to β decays, and therefore can only have a negligible impact on the latter once the stringent experimental bounds [8,89] are taken into account. Furthermore, the tensor amplitude generated by $Q_{\ell equ}^{(3)ijkl}$ has a suppressed matrix element in β decays.

Therefore, we are left with $Q_{\ell q}^{(3)1111}$, for which we only consider the flavor combination that leads to interference with the SM. The CAA prefers $C_{\ell q}^{(3)1111} \approx 0.7 \times 10^{-3} G_F$. Via $SU(2)_L$ invariance, this operator generates effects in neutral-current (NC) interactions

$$\mathcal{L}_{\text{NC}} = C_{\ell q}^{(3)1111} (\bar{d}\gamma^\mu P_L d - \bar{u}\gamma^\mu P_L u) \bar{e}\gamma_\mu P_L e. \quad (17)$$

Note that since the SM amplitude for $\bar{u}u(\bar{d}d) \rightarrow e^+e^-$, at high energies, has a negative (positive) sign, we have constructive interference in both amplitudes. Therefore, the latest nonresonant dilepton searches by ATLAS [90] naively lead to

$$C_{\ell q}^{(3)1111} \lesssim 1.6 \times 10^{-3} G_F, \quad (18)$$

which implies that bringing G_F^{CKM} into 1σ agreement with G_F^μ via four-fermion operators affecting $d \rightarrow ue\nu$ transitions is still possible. However, ATLAS derived the bound for the $SU(2)_L$ singlet operator, which means that the actual constraint for triplet operators is stronger because it leads to constructive interference in both the up- and the down-quark channels. In consequence, the required value of $C_{\ell q}^{(3)1111}$ lies at the border of the ATLAS constraint.

Modified W - u - d couplings.—There are only two operators modifying the W couplings to quarks:

$$\begin{aligned} Q_{\phi q}^{(3)ij} &= \phi^\dagger i D_\mu \phi \bar{q}_i \gamma^\mu \tau^I q_j, \\ Q_{\phi ud}^{ij} &= \tilde{\phi}^\dagger i D_\mu \phi \bar{u}_i \gamma^\mu d_j. \end{aligned} \quad (19)$$

First of all, $Q_{\phi ud}^{ij}$ generates right-handed W -quark couplings [91–95], which can solve the CAA. In addition, a right-handed W - u - s coupling could also account for the difference between $K_{\ell 2}$ and $K_{\ell 3}$ decays [96]. $Q_{\phi q}^{(3)ij}$ generates modifications of the left-handed W -quark couplings, and data prefer

$$C_{\phi q}^{(3)11} \approx 0.7 \times 10^{-3} G_F. \quad (20)$$

Due to $SU(2)_L$ invariance, in general, effects in $D^0-\bar{D}^0$ and $K^0-\bar{K}^0$ mixing (as well as in $\Gamma[Z \rightarrow \text{hadrons}]/\Gamma[Z \rightarrow \text{leptons}]$) are generated. However, the former bounds can be avoided by a $U(2)$ flavor symmetry, and the latter by simultaneous contributions to $Q_{\phi q}^{(1)ij}$. For a detailed analysis, we refer the reader to Ref. [97].

Modified W - ℓ - ν couplings.—Only the operator

$$Q_{\phi\ell}^{(3)ij} = \phi^\dagger i D_\mu \phi \bar{\ell}_i \gamma^\mu \tau^I \ell_j \quad (21)$$

generates modified W - ℓ - ν couplings at tree level. In order to avoid the stringent bounds from charged lepton flavor violation, the off-diagonal Wilson coefficients (in particular, $C_{\phi\ell}^{(3)12}$) must be very small. Since they also do not generate amplitudes interfering with the SM ones, their effect can be neglected. While $C_{\phi\ell}^{(3)11}$ affects G_F^μ and G_F^{CKM} in the same way, $C_{\phi\ell}^{(3)22}$ only enters in muon decay. Therefore, agreement between G_F^μ and G_F^{CKM} can be achieved by $C_{\phi\ell}^{(3)11} < 0$, $C_{\phi\ell}^{(3)22} > 0$, and $|C_{\phi\ell}^{(3)22}| < |C_{\phi\ell}^{(3)11}|$ without violating lepton flavor universality tests such as $\pi(K) \rightarrow \mu\nu/\pi(K) \rightarrow e\nu$ or $\tau \rightarrow \mu\nu/\tau(\mu) \rightarrow e\nu$ [55,58,98]. However, $C_{\phi\ell}^{(3)ij}$ also affects Z couplings to leptons and neutrinos, which enter the global EW fit.

Electroweak fit.—The impact of modified gauge-boson-lepton couplings on the global EW fit, generated by $Q_{\phi\ell}^{(3)ij}$ and

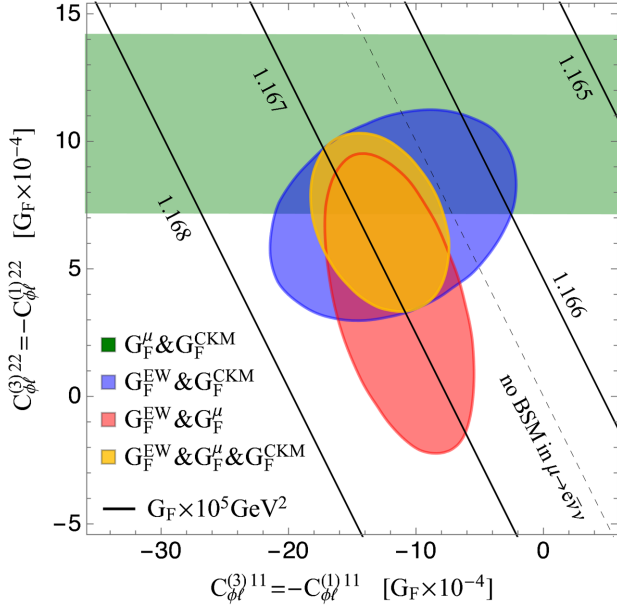


FIG. 2. Example of complementarity between G_F determinations from muon decay G_F^μ , CKM unitarity G_F^{CKM} , and global EW fit G_F^{EW} in case of $C_{\phi\ell}^{(3)ii} = -C_{\phi\ell}^{(1)ii}$, corresponding to modifications of neutrino couplings to gauge bosons (EW fit also includes $\tau \rightarrow \mu\nu\nu/\tau(\mu) \rightarrow e\nu\nu$ [8,54,98]). Here, we show preferred 1σ regions obtained by requiring two or all three G_F determinations agree. Contour lines show value of Fermi constant extracted from muon decay once BSM effects are taken into account.

$$Q_{\phi\ell}^{(1)ij} = \phi^\dagger i \overleftrightarrow{D}_\mu \phi \bar{\ell} \gamma^\mu \ell_j, \quad (22)$$

can be minimized by only affecting $Z\nu\nu$ but not $Z\ell\ell$ by imposing $C_{\phi\ell}^{(1)ij} = -C_{\phi\ell}^{(3)ij}$. In this way, in addition to the Fermi constant, only the Z width to neutrinos changes and the fit improves significantly compared to the SM [58]; see Fig. 2 for the preferred parameter space. One can even further improve the fit by assuming $C_{\phi\ell}^{(1)11} = -C_{\phi\ell}^{(3)11}$ and $C_{\phi\ell}^{(1)22} = -3C_{\phi\ell}^{(3)22}$, which leads to a better description of $Z \rightarrow \mu\mu$ data [61,62]. Furthermore, the part of the tension between G_F^{EW} and G_F^μ driven by the W mass can be alleviated by the operator $Q_{\phi WB} = \phi^\dagger \tau^I \phi W_{\mu\nu}^I B^{\mu\nu}$.

Conclusions and outlook.—Since the Fermi constant is determined extremely precisely by the muon lifetime, Eq. (1), constraining BSM effects via G_F , is limited entirely by the precision of the second-best determination. In this Letter, we derived in a first step two alternative independent determinations from the EW fit [Eq. (5)] and superallowed β decays using CKM unitarity [Eq. (9)]. The latter determination is more precise than the one from the EW fit, even though the precision of G_F^{EW} increased by a factor of 5 compared to Ref. [4]. Furthermore, as shown in Fig. 1, the two determinations display tensions of 2σ and 3σ compared to G_F^μ , respectively.

In a second step, we investigated how these hints of BSM physics can be explained within SMEFT. For BSM physics in G_F^μ , we were able to rule out all four-fermion operators (except for $Q_{\ell\ell}^{2112}$, which generates a SM-like amplitude) and modified $W\text{-}\ell\text{-}\nu$ couplings (from $Q_{\phi\ell}^{(3)ij}$). Therefore, both constructive and destructive interferences are possible, which would bring G_F^μ into agreement with G_F^{CKM} or G_F^{EW} , respectively, at the expense of increasing the tension with the other determination. To achieve a better agreement among the three different values of G_F , BSM effects in G_F^{CKM} and/or G_F^{EW} are also necessary. In the case of G_F^{CKM} only a single four-fermion operator, $Q_{\ell q}^{(3)1111}$, and $Q_{\phi\ell}^{(3)ij}$, $Q_{\phi q}^{(3)ij}$ remain. Finally, modified gauge-boson-lepton couplings, via $Q_{\phi\ell}^{(3)ij}$ and $Q_{\phi\ell}^{(1)ij}$, can not only change G_F^{CKM} and G_F^μ but also affect the EW fit via the Z -pole observables, which can further improve the global agreement; see Fig. 2. This figure also demonstrates the advantage of interpreting the tensions in terms of G_F , defining a transparent benchmark for comparison both in SMEFT and concrete BSM scenarios; and it allows one to constrain the amount of BSM contributions to muon decay.

Our study highlights the importance of improving the precision of the alternative independent determinations of G_F^{CKM} and G_F^{EW} in order to confirm or refute BSM contributions to the Fermi constant. Concerning G_F^{CKM} , improvements in the determination of $|V_{ud}|$ should arise from advances in nuclear-structure [99,100] and EW radiative corrections [101], while experimental developments [102–108] could make the determination from neutron decay [109–111] competitive and, in combination with $K_{\ell 3}$ decays, add another complementary constraint on $|V_{ud}|/|V_{us}|$ via pion β decay [112,113]. Furthermore, improved measurements of $|V_{cd}|$ from D decays [114] could bring the precision of the first-column CKM unitarity relation close to the first-row one, which, in turn, could be corroborated via improved $|V_{us}|$ determinations from $K_{\ell 3}$ decays [115–117]. The precision of G_F^{EW} will profit in the near future from LHC measurements of m_t and M_W , in the midterm from the Belle-II EW precision program [118], as well as in the long term from future e^+e^- colliders such as the FCC-ee [119], ILC [120], CEPC [121], or CLIC [122], which could achieve a precision at the level of 10^{-5} .

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- [1] D. M. Webber, V. Tishchenko, Q. Peng, S. Battu, R. M. Carey *et al.* (MuLan Collaboration), *Phys. Rev. Lett.* **106**, 041803 (2011).
- [2] V. Tishchenko, S. Battu, R. M. Carey, D. B. Chitwood, J. Crnkovic *et al.* (MuLan Collaboration), *Phys. Rev. D* **87**, 052003 (2013).
- [3] T. P. Gorringer and D. W. Hertzog, *Prog. Part. Nucl. Phys.* **84**, 73 (2015).
- [4] W. J. Marciano, *Phys. Rev. D* **60**, 093006 (1999).
- [5] B. W. Lee and R. E. Shrock, *Phys. Rev. D* **16**, 1444 (1977).
- [6] W. Buchmüller and D. Wyler, *Nucl. Phys.* **B268**, 621 (1986).
- [7] B. Grzadkowski, M. Iskrzyński, M. Misiak, and J. Rosiek, *J. High Energy Phys.* **10** (2010) 085.
- [8] P. A. Zyla *et al.* (Particle Data Group Collaboration), *Prog. Theor. Exp. Phys.* **2020**, 083C01 (2020).
- [9] S. Schael *et al.* (ALEPH, DELPHI, L3, OPAL, and LEP Electroweak Collaborations), *Phys. Rep.* **532**, 119 (2013).
- [10] S. Schael *et al.* (ALEPH, DELPHI, L3, OPAL, SLD, LEP Electroweak Working Group, SLD Electroweak Group, and SLD Heavy Flavour Group Collaborations), *Phys. Rep.* **427**, 257 (2006).
- [11] T. Aoyama *et al.*, *Phys. Rep.* **887**, 1 (2020).
- [12] M. Davier, A. Hoecker, B. Malaescu, and Z. Zhang, *Eur. Phys. J. C* **77**, 827 (2017).
- [13] A. Keshavarzi, D. Nomura, and T. Teubner, *Phys. Rev. D* **97**, 114025 (2018).
- [14] G. Colangelo, M. Hoferichter, and P. Stoffer, *J. High Energy Phys.* **02** (2019) 006.
- [15] M. Hoferichter, B.-L. Hoid, and B. Kubis, *J. High Energy Phys.* **08** (2019) 137.
- [16] M. Davier, A. Hoecker, B. Malaescu, and Z. Zhang, *Eur. Phys. J. C* **80**, 241 (2020); **80**, 410(E) (2020).
- [17] A. Keshavarzi, D. Nomura, and T. Teubner, *Phys. Rev. D* **101**, 014029 (2020).
- [18] B.-L. Hoid, M. Hoferichter, and B. Kubis, *Eur. Phys. J. C* **80**, 988 (2020).
- [19] S. Borsanyi *et al.*, *Nature (London)* **593**, 51 (2021).
- [20] A. Crivellin, M. Hoferichter, C. A. Manzari, and M. Montull, *Phys. Rev. Lett.* **125**, 091801 (2020).
- [21] A. Keshavarzi, W. J. Marciano, M. Passera, and A. Sirlin, *Phys. Rev. D* **102**, 033002 (2020).
- [22] B. Malaescu and M. Schott, *Eur. Phys. J. C* **81**, 46 (2021).
- [23] G. Colangelo, M. Hoferichter, and P. Stoffer, *Phys. Lett. B* **814**, 136073 (2021).
- [24] J. De Blas *et al.*, *Eur. Phys. J. C* **80**, 456 (2020).
- [25] A. Caldwell, D. Kollar, and K. Kroninger, *Comput. Phys. Commun.* **180**, 2197 (2009).
- [26] M. Baak, J. Cúth, J. Haller, A. Hoecker, R. Kogler, K. Mönig, M. Schott, and J. Stelzer (Gitter Group Collaboration), *Eur. Phys. J. C* **74**, 3046 (2014).
- [27] J. de Blas, M. Ciuchini, E. Franco, S. Mishima, M. Pierini, L. Reina, and L. Silvestrini, *J. High Energy Phys.* **12** (2016) 135.
- [28] T. A. Aaltonen *et al.* (CDF, D0 Collaborations), *Phys. Rev. D* **97**, 112007 (2018).
- [29] R. Aaij *et al.* (LHCb Collaboration), *J. High Energy Phys.* **11** (2015) 190.
- [30] G. Aad *et al.* (ATLAS Collaboration), *J. High Energy Phys.* **09** (2015) 049.
- [31] The ATLAS collaboration, Report No. ATLAS-CONF-2018-037, 2018, <http://cds.cern.ch/record/2630340>.
- [32] A. M. Sirunyan *et al.* (CMS Collaboration), *Eur. Phys. J. C* **78**, 701 (2018).
- [33] S. Aoki *et al.* (Flavour Lattice Averaging Group Collaboration), *Eur. Phys. J. C* **80**, 113 (2020).
- [34] R. Barate *et al.* (ALEPH Collaboration), *Eur. Phys. J. C* **14**, 1 (2000).
- [35] G. Abbiendi *et al.* (OPAL Collaboration), *Eur. Phys. J. C* **19**, 587 (2001).
- [36] P. Abreu *et al.* (DELPHI Collaboration), *Eur. Phys. J. C* **16**, 371 (2000).
- [37] M. Acciarri *et al.* (L3 Collaboration), *Eur. Phys. J. C* **16**, 1 (2000).
- [38] G. Aad *et al.* (ATLAS and CMS Collaborations), *Phys. Rev. Lett.* **114**, 191803 (2015).
- [39] M. Aaboud *et al.* (ATLAS Collaboration), *Phys. Lett. B* **784**, 345 (2018).
- [40] A. M. Sirunyan *et al.* (CMS Collaboration), *J. High Energy Phys.* **11** (2017) 047.
- [41] V. Khachatryan *et al.* (CMS Collaboration), *Phys. Rev. D* **93**, 072004 (2016).
- [42] CDF and D0 Collaborations, Report No. FERMILAB-CONF-16-298-E, 2016, <https://inspirehep.net/literature/1479761>.
- [43] M. Aaboud *et al.* (ATLAS Collaboration), *Eur. Phys. J. C* **79**, 290 (2019).
- [44] A. M. Sirunyan *et al.* (CMS Collaboration), *Eur. Phys. J. C* **79**, 313 (2019).
- [45] W. J. Marciano and A. Sirlin, *Phys. Rev. Lett.* **96**, 032002 (2006).
- [46] C.-Y. Seng, M. Gorchtein, H. H. Patel, and M. J. Ramsey-Musolf, *Phys. Rev. Lett.* **121**, 241804 (2018).
- [47] C. Y. Seng, M. Gorchtein, and M. J. Ramsey-Musolf, *Phys. Rev. D* **100**, 013001 (2019).
- [48] M. Gorchtein, *Phys. Rev. Lett.* **123**, 042503 (2019).
- [49] A. Czarnecki, W. J. Marciano, and A. Sirlin, *Phys. Rev. D* **100**, 073008 (2019).
- [50] C.-Y. Seng, X. Feng, M. Gorchtein, and L.-C. Jin, *Phys. Rev. D* **101**, 111301(R) (2020).
- [51] L. Hayen, *Phys. Rev. D* **103**, 113001 (2021).
- [52] J. C. Hardy and I. S. Towner, *Phys. Rev. C* **102**, 045501 (2020).
- [53] M. Moulson, *Proc. Sci. CKM2016* (2017) 033.
- [54] Y. S. Amhis *et al.* (HFLAV Collaboration), *Eur. Phys. J. C* **81**, 226 (2021).
- [55] A. Crivellin and M. Hoferichter, *Phys. Rev. Lett.* **125**, 111801 (2020).
- [56] B. Belfatto, R. Beradze, and Z. Berezhiani, *Eur. Phys. J. C* **80**, 149 (2020).
- [57] D. A. Bryman and R. Shrock, *Phys. Rev. D* **100**, 073011 (2019).
- [58] A. M. Coutinho, A. Crivellin, and C. A. Manzari, *Phys. Rev. Lett.* **125**, 071802 (2020).
- [59] B. Capdevila, A. Crivellin, C. A. Manzari, and M. Montull, *Phys. Rev. D* **103**, 015032 (2021).
- [60] M. Endo and S. Mishima, *J. High Energy Phys.* **08** (2020) 004.

- [61] A. Crivellin, F. Kirk, C. A. Manzari, and M. Montull, *J. High Energy Phys.* **12** (2020) 166.
- [62] M. Kirk, *Phys. Rev. D* **103**, 035004 (2021).
- [63] A. K. Alok, A. Dighe, S. Gangal, and J. Kumar, [arXiv:2010.12009](https://arxiv.org/abs/2010.12009).
- [64] A. Crivellin, C. A. Manzari, M. Alguero, and J. Matias, *Phys. Rev. Lett.* **127**, 011801 (2021).
- [65] A. Crivellin, F. Kirk, C. A. Manzari, and L. Panizzi, *Phys. Rev. D* **103**, 073002 (2021).
- [66] A. Crivellin, D. Müller, and L. Schnell, *Phys. Rev. D* **103**, 115023 (2021).
- [67] Z. Han and W. Skiba, *Phys. Rev. D* **71**, 075009 (2005).
- [68] V. Cirigliano, J. Jenkins, and M. González-Alonso, *Nucl. Phys.* **B830**, 95 (2010).
- [69] A. Falkowski and F. Riva, *J. High Energy Phys.* **02** (2015) 039.
- [70] A. Falkowski and K. Mimouni, *J. High Energy Phys.* **02** (2016) 086.
- [71] J. Ellis, C. W. Murphy, V. Sanz, and T. You, *J. High Energy Phys.* **06** (2018) 146.
- [72] W. Skiba and Q. Xia, [arXiv:2007.15688](https://arxiv.org/abs/2007.15688).
- [73] L. Michel, *Proc. Phys. Soc. London Sect. A* **63**, 514 (1950).
- [74] L. Michel and A. Wightman, *Phys. Rev.* **93**, 354 (1954).
- [75] T. Kinoshita and A. Sirlin, *Phys. Rev.* **107**, 593 (1957).
- [76] F. Scheck, *Phys. Rep.* **44**, 187 (1978).
- [77] W. Fetscher, H. J. Gerber, and K. F. Johnson, *Phys. Lett. B* **173**, 102 (1986).
- [78] N. Danneberg *et al.*, *Phys. Rev. Lett.* **94**, 021802 (2005).
- [79] R. P. MacDonald, R. Bayes, J. Bueno, Y. I. Davydov, P. Depommier *et al.* (TWIST Collaboration), *Phys. Rev. D* **78**, 032010 (2008).
- [80] R. Bayes, J. F. Bueno, A. Hillairet, Y. I. Davydov, P. Depommier *et al.* (TWIST Collaboration), *Phys. Rev. Lett.* **106**, 041804 (2011).
- [81] L. Willmann, P. V. Schmidt, H. P. Wirtz, R. Abela, V. Baranov *et al.*, *Phys. Rev. Lett.* **82**, 49 (1999).
- [82] G. Feinberg and S. Weinberg, *Phys. Rev.* **123**, 1439 (1961).
- [83] W.-S. Hou and G.-G. Wong, *Phys. Lett. B* **357**, 145 (1995).
- [84] K. Horikawa and K. Sasaki, *Phys. Rev. D* **53**, 560 (1996).
- [85] U. Bellgardt *et al.* (SINDRUM Collaboration), *Nucl. Phys.* **B299**, 1 (1988).
- [86] A. Crivellin, S. Davidson, G. M. Pruna, and A. Signer, *J. High Energy Phys.* **05** (2017) 117.
- [87] K. Hayasaka *et al.*, *Phys. Lett. B* **687**, 139 (2010).
- [88] J. P. Lees *et al.* (BABAR Collaboration), *Phys. Rev. D* **81**, 111101 (2010).
- [89] A. Aguilar-Arevalo, M. Aoki, M. Blecher, D. I. Britton, D. A. Bryman *et al.* (PiENU Collaboration), *Phys. Rev. Lett.* **115**, 071801 (2015).
- [90] G. Aad *et al.* (ATLAS Collaboration), *J. High Energy Phys.* **11** (2020) 005.
- [91] V. Bernard, M. Oertel, E. Passemar, and J. Stern, *J. High Energy Phys.* **01** (2008) 015.
- [92] A. Crivellin, *Phys. Rev. D* **81**, 031301(R) (2010).
- [93] A. J. Buras, K. Gemmler, and G. Isidori, *Nucl. Phys.* **B843**, 107 (2011).
- [94] A. Crivellin and S. Pokorski, *Phys. Rev. Lett.* **114**, 011802 (2015).
- [95] S. Alioli, V. Cirigliano, W. Dekens, J. de Vries, and E. Mereghetti, *J. High Energy Phys.* **05** (2017) 086.
- [96] Y. Grossman, E. Passemar, and S. Schacht, *J. High Energy Phys.* **07** (2020) 068.
- [97] B. Belfatto and Z. Berezhiani, [arXiv:2103.05549](https://arxiv.org/abs/2103.05549).
- [98] A. Pich, *Prog. Part. Nucl. Phys.* **75**, 41 (2014).
- [99] V. Cirigliano, A. Garcia, D. Gazit, O. Naviliat-Cuncic, G. Savard, and A. Young, [arXiv:1907.02164](https://arxiv.org/abs/1907.02164).
- [100] M. S. Martin, S. R. Stroberg, J. D. Holt, and K. G. Leach, [arXiv:2101.11826](https://arxiv.org/abs/2101.11826).
- [101] X. Feng, M. Gorchtein, L.-C. Jin, P.-X. Ma, and C.-Y. Seng, *Phys. Rev. Lett.* **124**, 192002 (2020).
- [102] J. Fry *et al.*, *Europhys. J. Web Conf.* **219**, 04002 (2019).
- [103] T. Soldner, H. Abele, G. Konrad, B. Märkisch, F. M. Piegsa, U. Schmidt, C. Theroine, and P. T. Sánchez, *Europhys. J. Web Conf.* **219**, 10003 (2019).
- [104] X. Wang *et al.* (PERC Collaboration), *Europhys. J. Web Conf.* **219**, 04007 (2019).
- [105] A. P. Serebrov, O. M. Zhrebtsov, A. N. Murashkin, G. N. Klyushnikov, and A. K. Fomin, *Phys. At. Nucl.* **82**, 98 (2019).
- [106] D. Gaisbauer, I. Konorov, D. Steffen, and S. Paul, *Nucl. Instrum. Methods Phys. Res., Sect. A* **824**, 290 (2016).
- [107] V. F. Ezhov, V. L. Ryabov, A. Z. Andreev, B. A. Bazarov, A. G. Glushkov, and V. A. Knyaz'kov, *Tech. Phys. Lett.* **44**, 602 (2018).
- [108] N. Callahan, C. Y. Liu, F. Gonzalez, E. Adamek, J. D. Bowman *et al.* (UCN τ Collaboration), *Phys. Rev. C* **100**, 015501 (2019).
- [109] R. W. Pattie Jr. *et al.*, *Science* **360**, 627 (2018).
- [110] B. Märkisch *et al.*, *Phys. Rev. Lett.* **122**, 242501 (2019).
- [111] A. Czarnecki, W. J. Marciano, and A. Sirlin, *Phys. Rev. Lett.* **120**, 202002 (2018).
- [112] D. Počanić, E. Frlež, V. A. Baranov, W. Bertl, C. Bronnimann *et al.*, *Phys. Rev. Lett.* **93**, 181803 (2004).
- [113] A. Czarnecki, W. J. Marciano, and A. Sirlin, *Phys. Rev. D* **101**, 091301(R) (2020).
- [114] M. Ablikim *et al.* (BESIII Collaboration), *Chin. Phys. C* **44**, 040001 (2020).
- [115] O. P. Yushchenko *et al.* (OKA Collaboration), *JETP Lett.* **107**, 139 (2018).
- [116] A. A. Alves Junior *et al.*, *J. High Energy Phys.* **05** (2019) 048.
- [117] D. Babusci *et al.* (KLOE-2 Collaboration), *Phys. Lett. B* **804**, 135378 (2020).
- [118] W. Altmannshofer *et al.* (Belle-II Collaboration), *Prog. Theor. Exp. Phys.* **2019**, 123C01 (2019); **2020**, 029201(E) (2020).
- [119] A. Abada *et al.* (FCC Collaboration), *Eur. Phys. J. Special Topics* **228**, 261 (2019).
- [120] H. Baer *et al.*, [arXiv:1306.6352](https://arxiv.org/abs/1306.6352).
- [121] F. An *et al.*, *Chin. Phys. C* **43**, 043002 (2019).
- [122] M. Aicheler *et al.*, Reports No. CERN-2012-007, No. SLAC-R-985, No. KEK-Report-2012-1, No. PSI-12-01, No. JAI-2012-001, 2012, <https://doi.org/10.5170/CERN-2012-007>.